

# Review of Lectures 9 to 16

## AOSC 433/633 & CHEM 433

Ross Salawitch

- Exam on Thursday:
  - Conceptual questions only: no calculators
  - Closed book; no notes
- ELMS gradebook should be current
  - please let us know of any issues
- Please note:
  - Last day to drop a course with a "W" is 12 April
  - Students dropping a course on Testudo have until 11pm on 12 April
  - Drop/Add services for Spring 2017 will no longer be available after 11pm on 12 April
  - Students who need to come in person to drop a course have until 5 pm on 12 April

**11 April 2017**

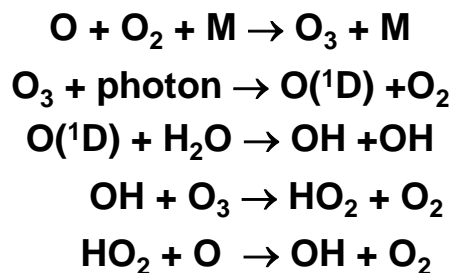
# Importance of Radicals

- With a few exceptions, the only reactions between molecules that proceed at appreciable rates are those involving at least one radical
- Radicals require significant energy to form: a bond must be broken
- Radical formation is tied to absorption of photons that “photodissociate” a compound, leading to radical formation

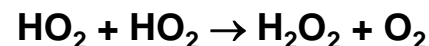
## Initiation



## Propagation

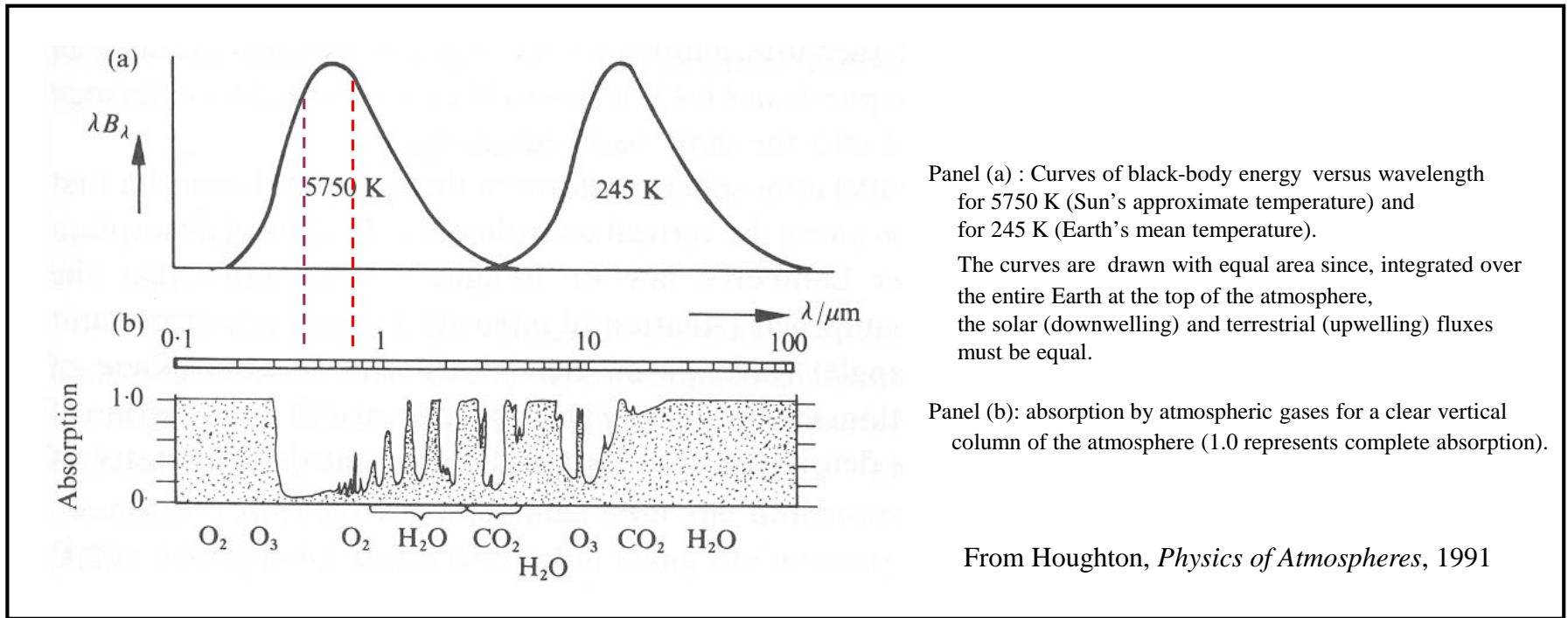


## Termination



# Atmospheric Radiation

- Solar irradiance (downwelling) at top of atmosphere occurs at wavelengths between ~200 and 2000 nm (~5750 K “black body” temperature)

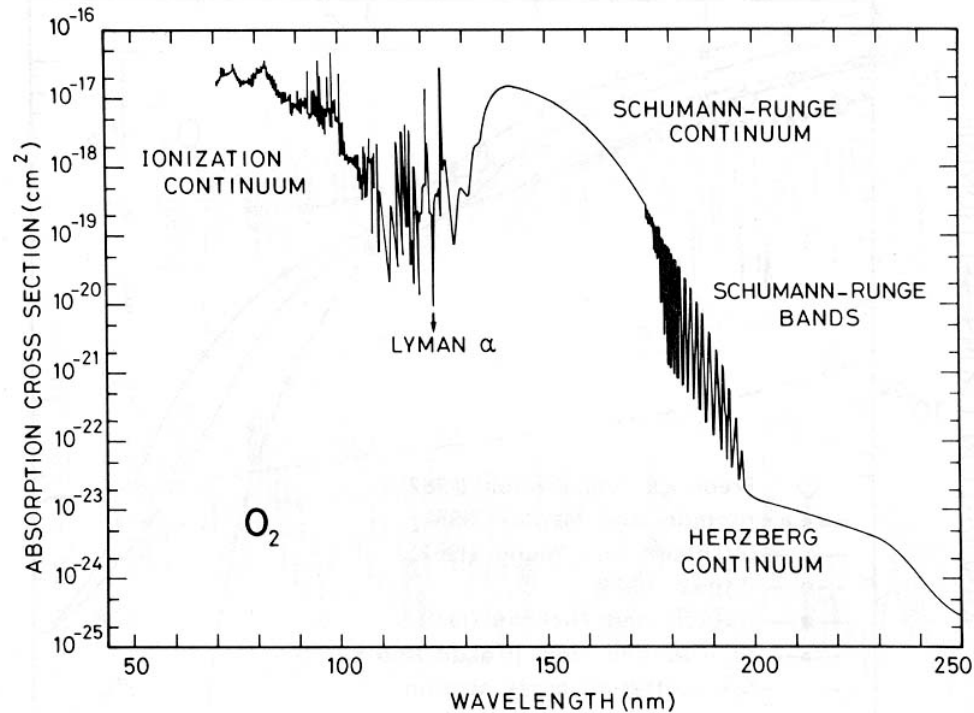


Panel (a) : Curves of black-body energy versus wavelength for 5750 K (Sun's approximate temperature) and for 245 K (Earth's mean temperature).  
The curves are drawn with equal area since, integrated over the entire Earth at the top of the atmosphere, the solar (downwelling) and terrestrial (upwelling) fluxes must be equal.  
Panel (b): absorption by atmospheric gases for a clear vertical column of the atmosphere (1.0 represents complete absorption).

From Houghton, *Physics of Atmospheres*, 1991

- Absorption and photodissociation in the UV occurs due to changes in the electronic state (orbital configuration) of molecules

# Absorption Cross Section of O<sub>2</sub>



From Brasseur & Solomon, *Aeronomy of the Middle Atmosphere*, 1986

- O<sub>2</sub> can not dissociate longward of ~250 nm
- All absorption shown above is dissociative (e.g., leads to production of two O atoms)
- Structure in the O<sub>2</sub> cross section is related to whether the initial transition involves an unbound electronic state (smooth) or involves a specific vibrational level of an electronic state (banded, due to requirement of specific quanta of energy)

# Optical Depth of O<sub>2</sub> Absorption

Recall the *Beer-Lambert Law*:

$$F(z, \lambda) = F_{\text{TOA}}(\lambda) e^{-\tau(z, \lambda)} \quad (\text{TOA : Top of Atmosphere})$$

where:

$$\tau(z, \lambda) = m \int_z^{\infty} \sigma_{\lambda} [C] dz' \quad (\tau: \text{optical depth})$$

Also:

$$\int_0^{\infty} [\text{O}_2] dz' \approx 4 \times 10^{24} \text{ molecules/cm}^2$$

O <sub>2</sub> Optical Depth for $\theta = 0^\circ$ , $z = 0$ km			
	$\sigma_{\text{max}}$ (cm <sup>2</sup> )	$\tau$ (0 km)	$e^{-\tau}$ (0 km)
Schumann-Runge Continuum	$10^{-17}$	$4 \times 10^7$	0.
Schumann-Runge Bands	$10^{-20}$	$4 \times 10^4$	0.
	$3 \times 10^{-23}$	120	$7.6 \times 10^{-53}$
Herzberg Continuum	$10^{-23}$	40	$4.2 \times 10^{-18}$

# Optical Depth of O<sub>3</sub> Absorption

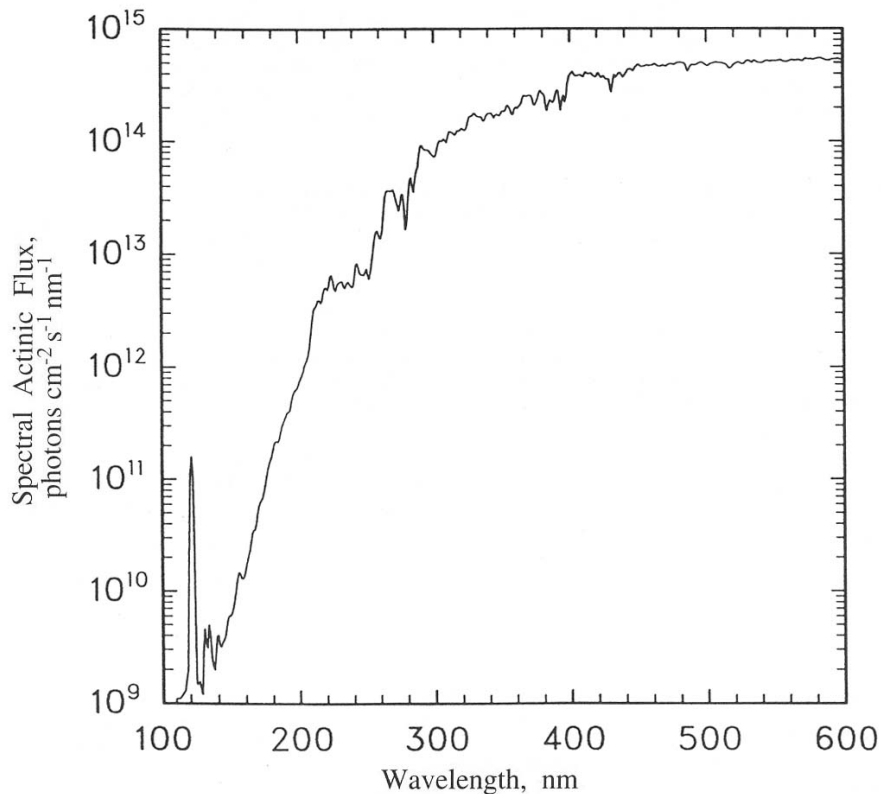
A typical mid-latitude column abundance for O<sub>3</sub> is 300 Dobson units (DU):

$$1 \text{ DU} = 2.687 \times 10^{16} \text{ molecules/cm}^2; \quad 300 \text{ DU} = 8 \times 10^{18} \text{ molecules/cm}^2$$

Aside: 
$$\frac{\text{Column O}_3}{\text{Column Air}} = \frac{8 \times 10^{18}}{2 \times 10^{25}} = 0.4 \text{ parts per million} \Rightarrow \text{Ozone is a trace species!}$$

O <sub>3</sub> Optical Depth for $\theta = 0^\circ$ , $z = 0 \text{ km}$				
	$\sigma_{\text{max}} \text{ (cm}^2\text{)}$	$\tau \text{ (0 km)}$	$e^{-\tau \text{ (0 km)}}$	O <sub>3</sub> Column, $\tau = 1.0$
Hartley (~220 to 280 nm)	$10^{-17}$	80	$1.8 \times 10^{-35}$	3.7 DU
Huggins (~310 to 330 nm)	$10^{-19}$	0.8	0.45	372 DU
Chappuis (~500 to 700 nm)	$3 \times 10^{-21}$	0.024	~1.0	12,400 DU

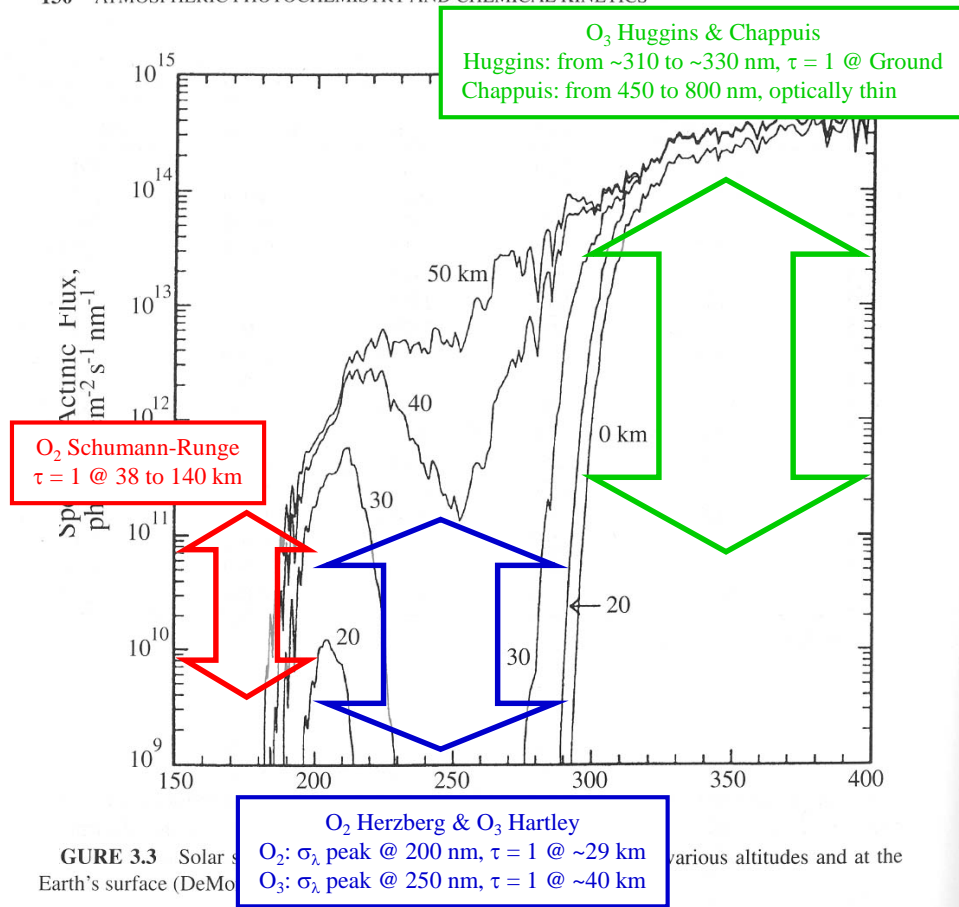
# Solar Spectral Actinic Flux



**FIGURE 6.** Solar spectral actinic flux (photons  $\text{cm}^{-2} \text{s}^{-1} \text{nm}^{-1}$ ) at the top of Earth's atmosphere.

From DeMore et al., *Chemical Kinetics and Photochemical Data for Use in Stratospheric Modeling*, Evaluation No. 11, 1994.

## 130 ATMOSPHERIC PHOTOCHEMISTRY AND CHEMICAL KINETICS



**FIGURE 3.3** Solar spectral actinic flux at various altitudes and at the Earth's surface (DeMore et al., 1994).

From Seinfeld and Pandis, *Atmospheric Chemistry and Physics*, 1998.

# Photolysis Frequency

For a specific spectral interval, the photolysis frequency (*partial J value*) of a gas is given by the product of its absorption cross section and the solar irradiance:

$$J_{\text{gas}}(z, \lambda) = \text{Quantum\_Yield}(\lambda) \sigma_{\text{gas}}(\lambda, T) F(z, \lambda)$$

Units:  $\text{s}^{-1} \text{ nm}^{-1}$

The total *photolysis frequency* (*J value*) is found by integrating  $J_{\text{gas}}(z, \lambda)$  over all wavelengths for which the gas photodissociates:

$$J_{\text{gas}}(z) = \int_{\lambda_{\text{min}}}^{\lambda_{\text{max}}} J_{\text{gas}}(z, \lambda) d\lambda$$

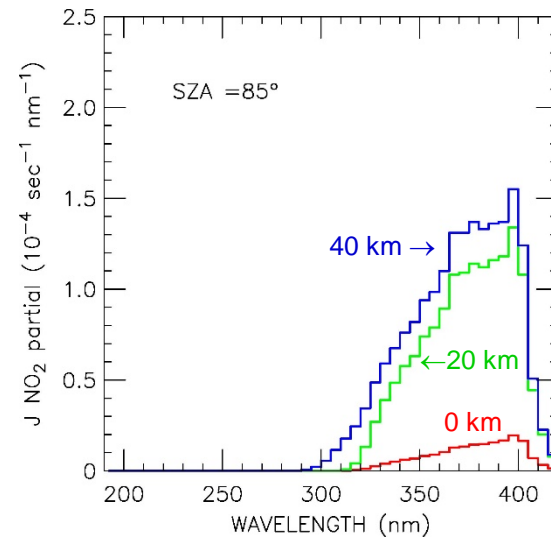
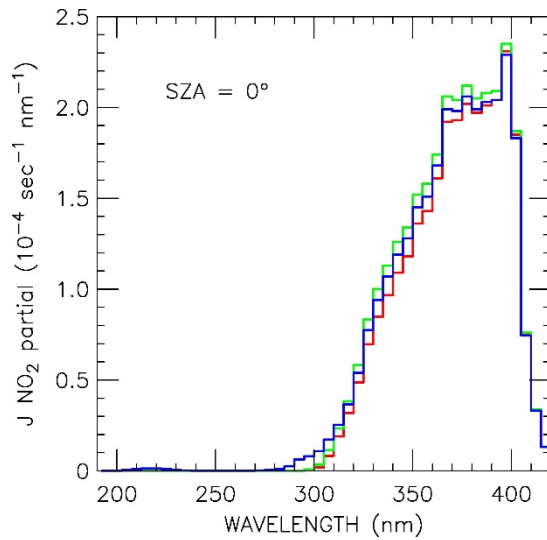
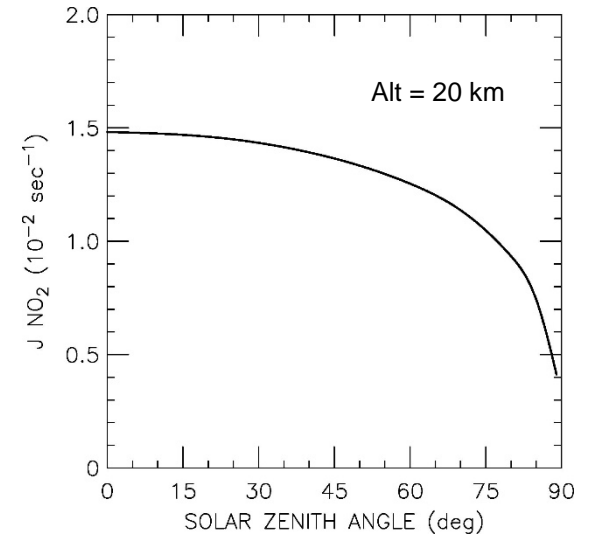
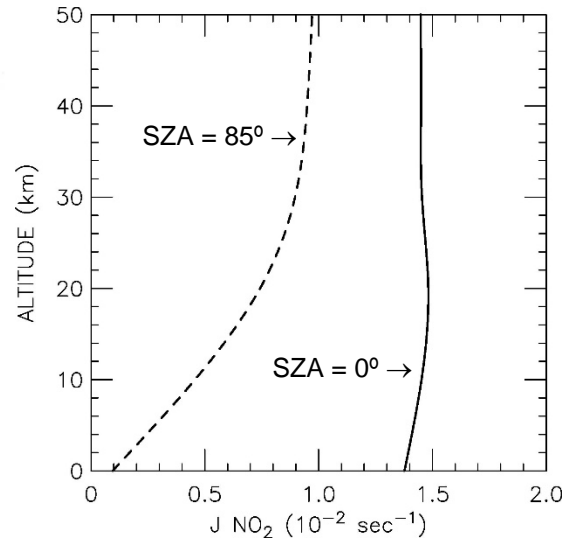
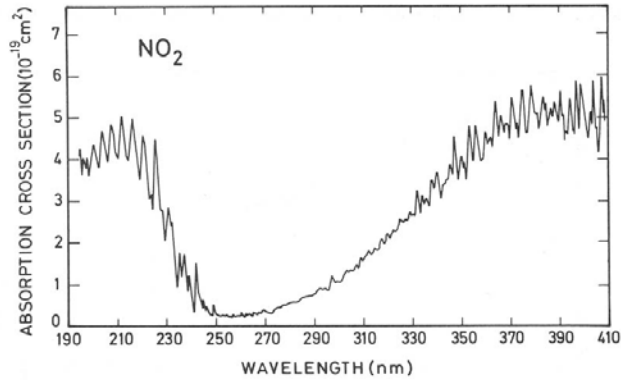
Units:  $\text{s}^{-1}$

$$\text{Rate of Reaction} = \frac{d\text{O}_3}{dt} = J [\text{O}_3]; \text{ Units of } J \text{ are } \text{s}^{-1}$$

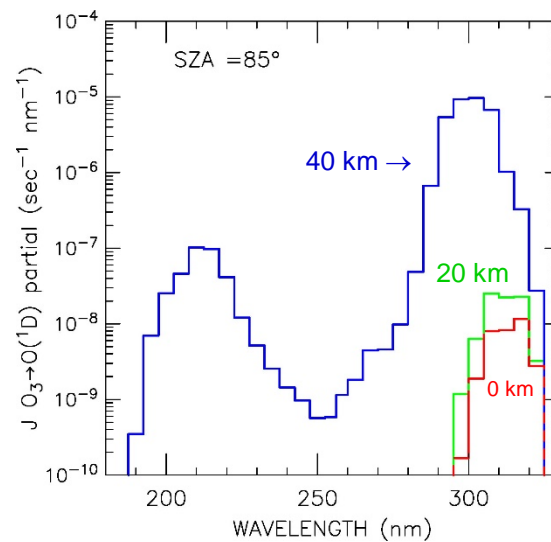
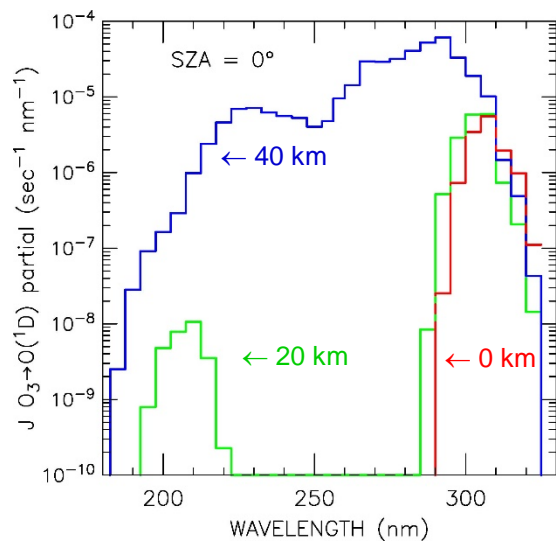
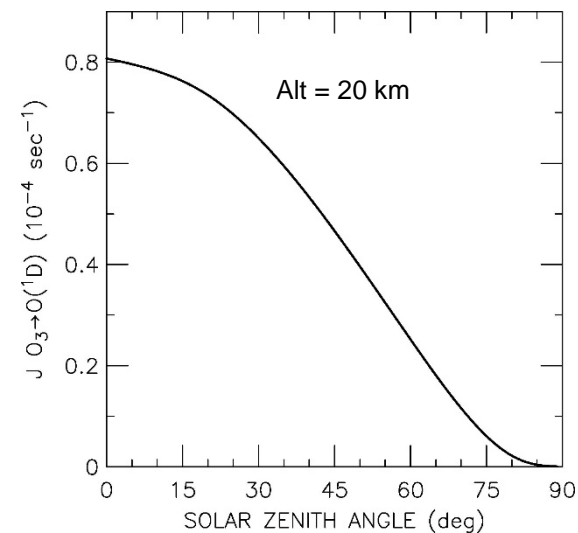
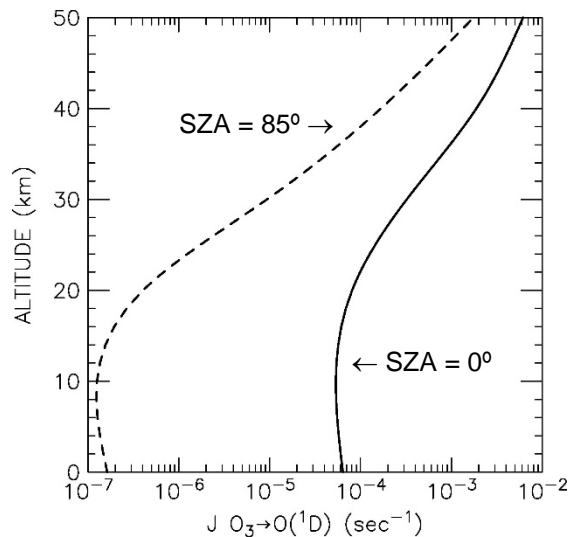
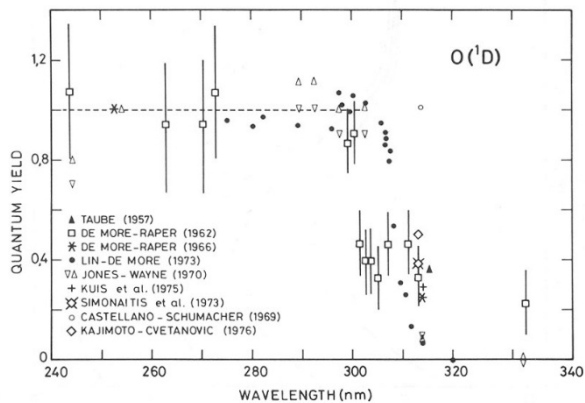
*More precisely, calculations of photolysis frequencies consider the “spectral actinic flux”, which represents the amount of available photons integrated over all angles, rather than “solar irradiance”. These two quantities differ because of scattering of solar radiation by gases and aerosols, and reflection of radiation by clouds and the surface.*



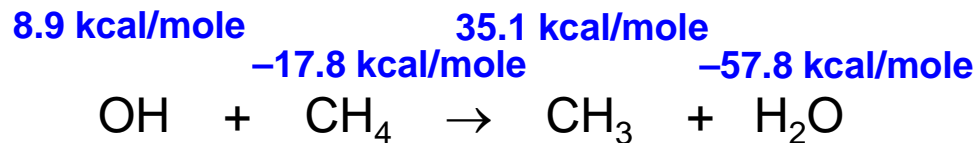
# NO<sub>2</sub> Photolysis



# O<sub>3</sub> → O(<sup>1</sup>D) Photolysis



# Bimolecular Gas Phase Reactions



$\Delta\text{Enthalpy} = -13.8 \text{ kcal/mole}$

**Exothermic !**

$$\text{Rate of Reaction} = \frac{d\text{CH}_4}{dt} = k [\text{OH}][\text{CH}_4]$$

$E_A / R \Rightarrow$  Activation Energy / Gas Constant

Arrhenius Expression for rate constant:

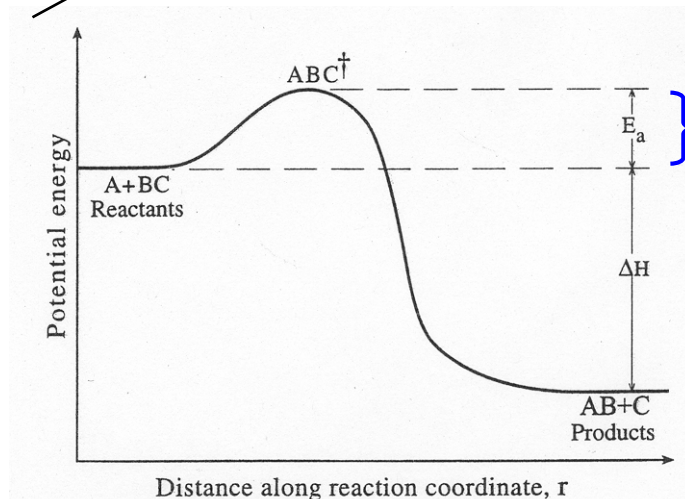
$$k = 2.45 \times 10^{-12} \times e^{-1775/T} \text{ cm}^3 \text{ sec}^{-1}$$

**Energy Term**

**A factor**

$$k_c e^{\Delta S / R}$$

**Entropy Term**



$$\begin{aligned}
 R &= 8.3143 \times 10^7 \text{ erg / (K mole)} \\
 &= 2.87 \times 10^6 \text{ erg / (K gm) for air}
 \end{aligned}$$

Figure 3.3 Barrier energies for the forward reaction ( $E_a$ ) and the reverse reaction ( $E_a + \Delta H$ ).

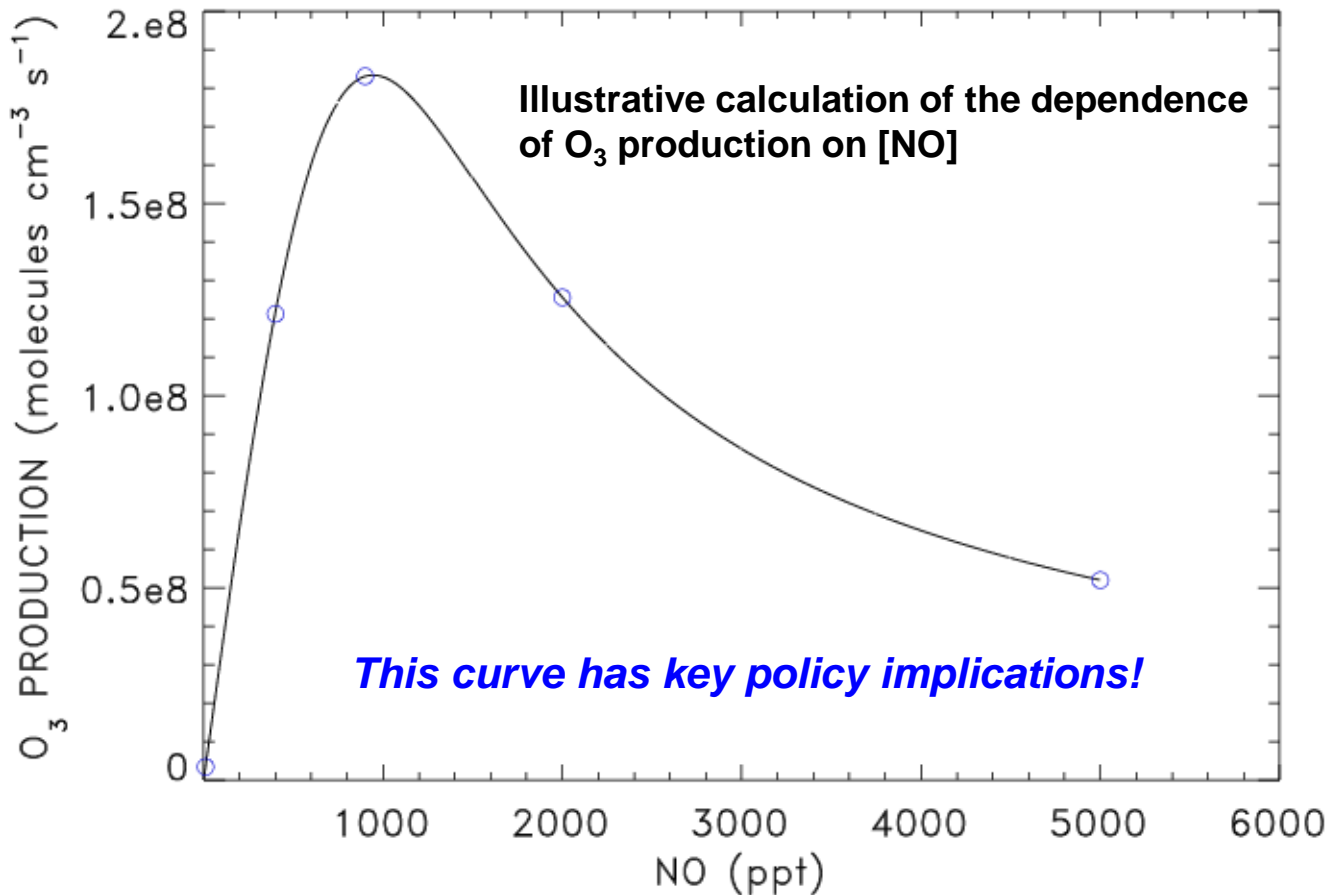
Yung and DeMore, *Photochemistry of Planetary Atmospheres*, Oxford, 1999.

# Tropospheric Ozone Production versus NO

Production of Tropospheric O<sub>3</sub> limited by: \_\_\_\_\_ ?

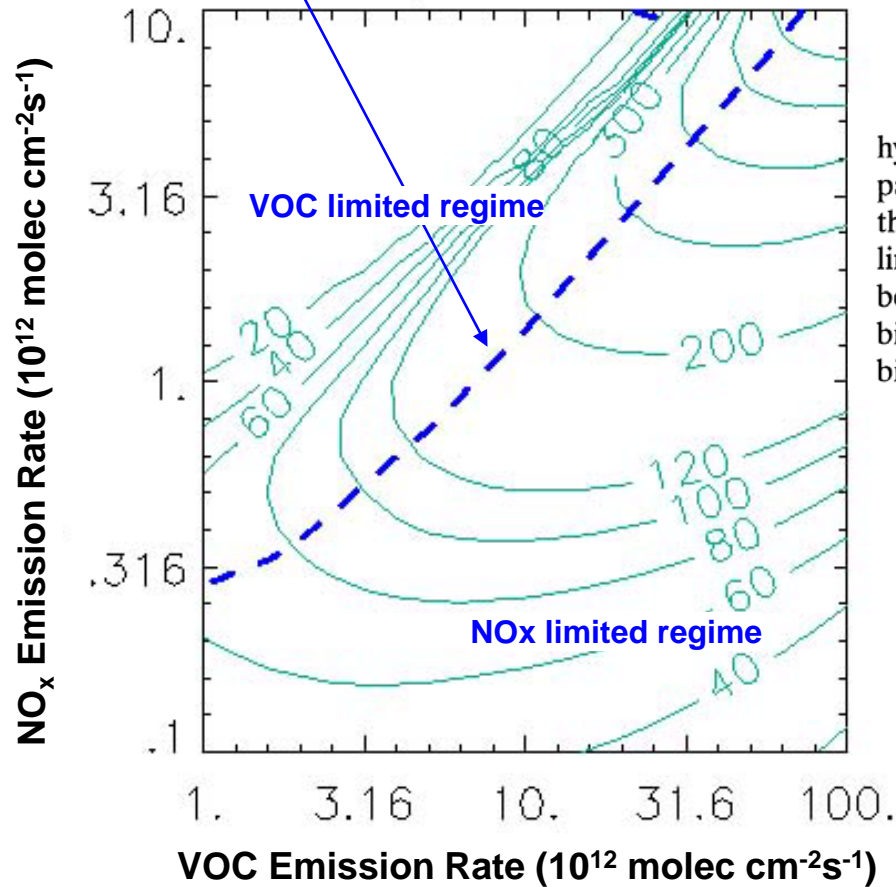
As NO<sub>x</sub> rises:

[HO<sub>2</sub>] falls faster than [NO] rises,  
leading to a decrease in the value \_\_\_\_\_



# Tropospheric Ozone Production versus $\text{NO}_x$ and VOCs

Ridge: local maximum for  $\text{O}_3$  that separates the  $\text{NO}_x$ -limited regime from and VOC limited regime



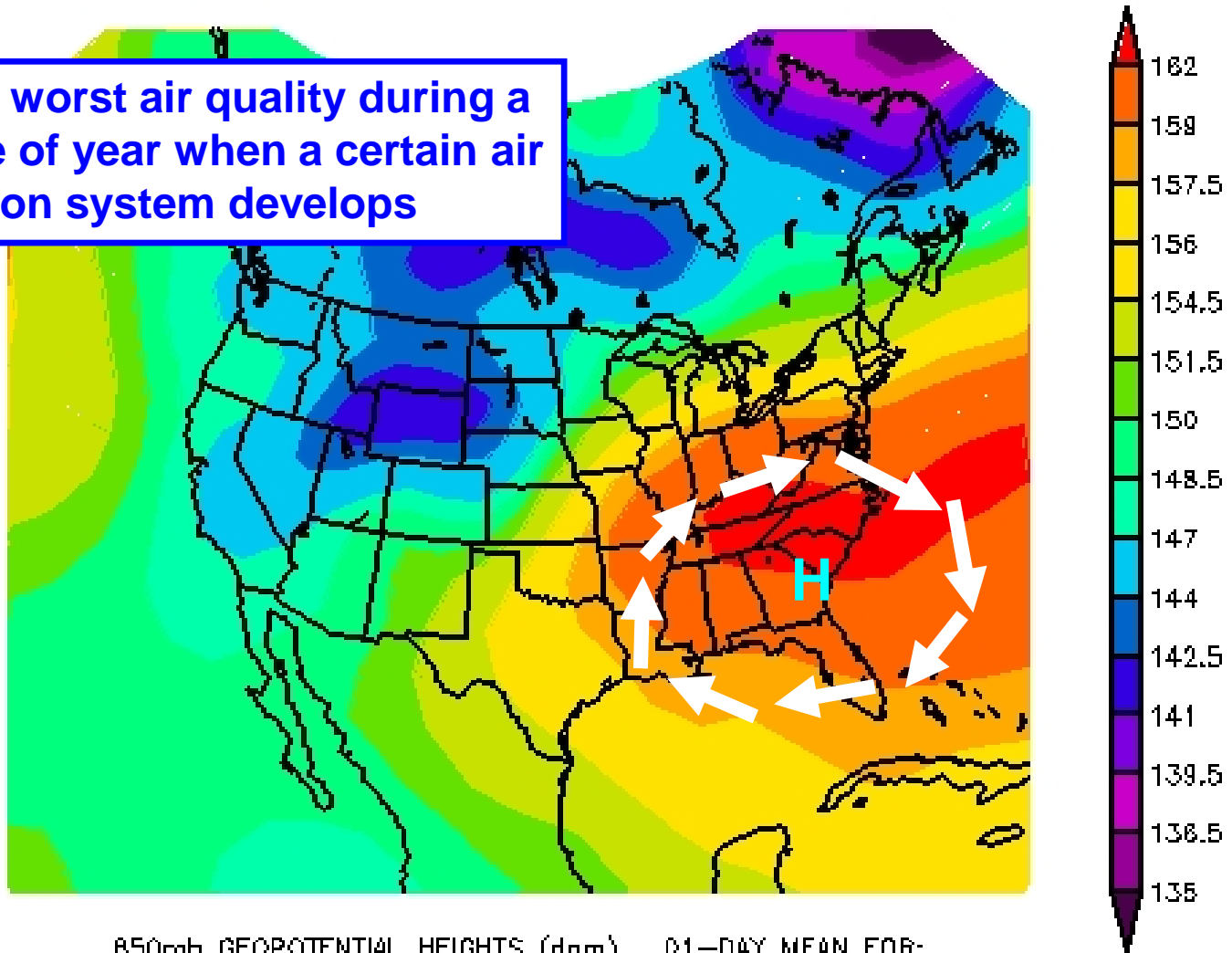
An important discovery in the past decade is that the focus on hydrocarbon emission controls to combat  $\text{O}_3$  pollution may have been partly misdirected. Measurements and model calculations now show that  $\text{O}_3$  production over most of the United States is primarily  $\text{NO}_x$  limited, not hydrocarbon limited. The early models were in error in part because they underestimated emissions of hydrocarbons from automobiles, and in part because they did not account for natural emission of biogenic hydrocarbons from trees and crops.

Jacob, Chapter 12, Introduction to Atmospheric Chemistry, 1999

Figure: <http://www-personal.umich.edu/~sillman/ozone.htm>

# Day-to-day meteorology (weather!) affects severity and duration of pollution episodes

Maryland has worst air quality during a particular time of year when a certain air circulation system develops

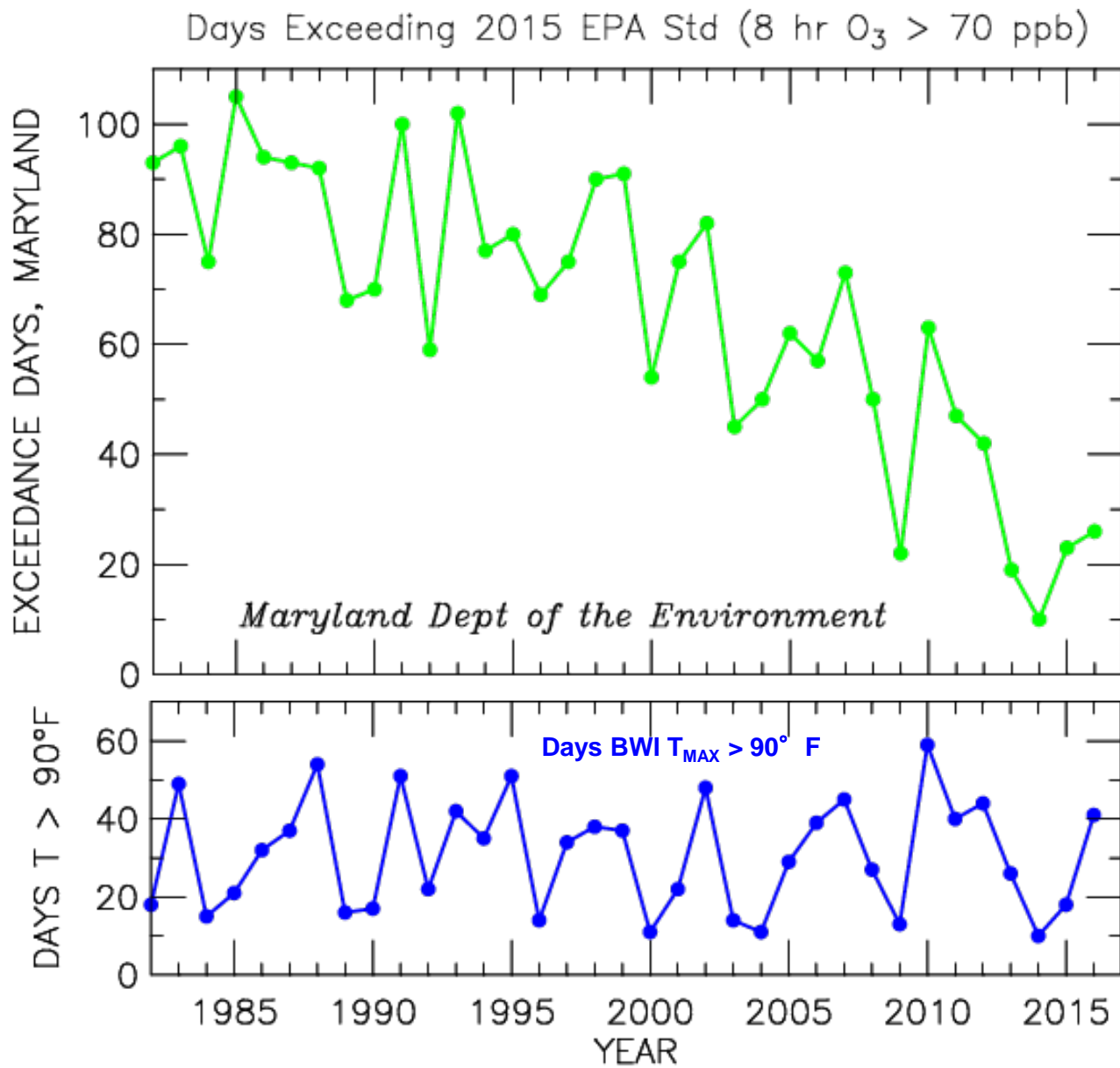


850mb GEOPOTENTIAL HEIGHTS (dam) 01-DAY MEAN FOR:  
Sun JUL 04 1999

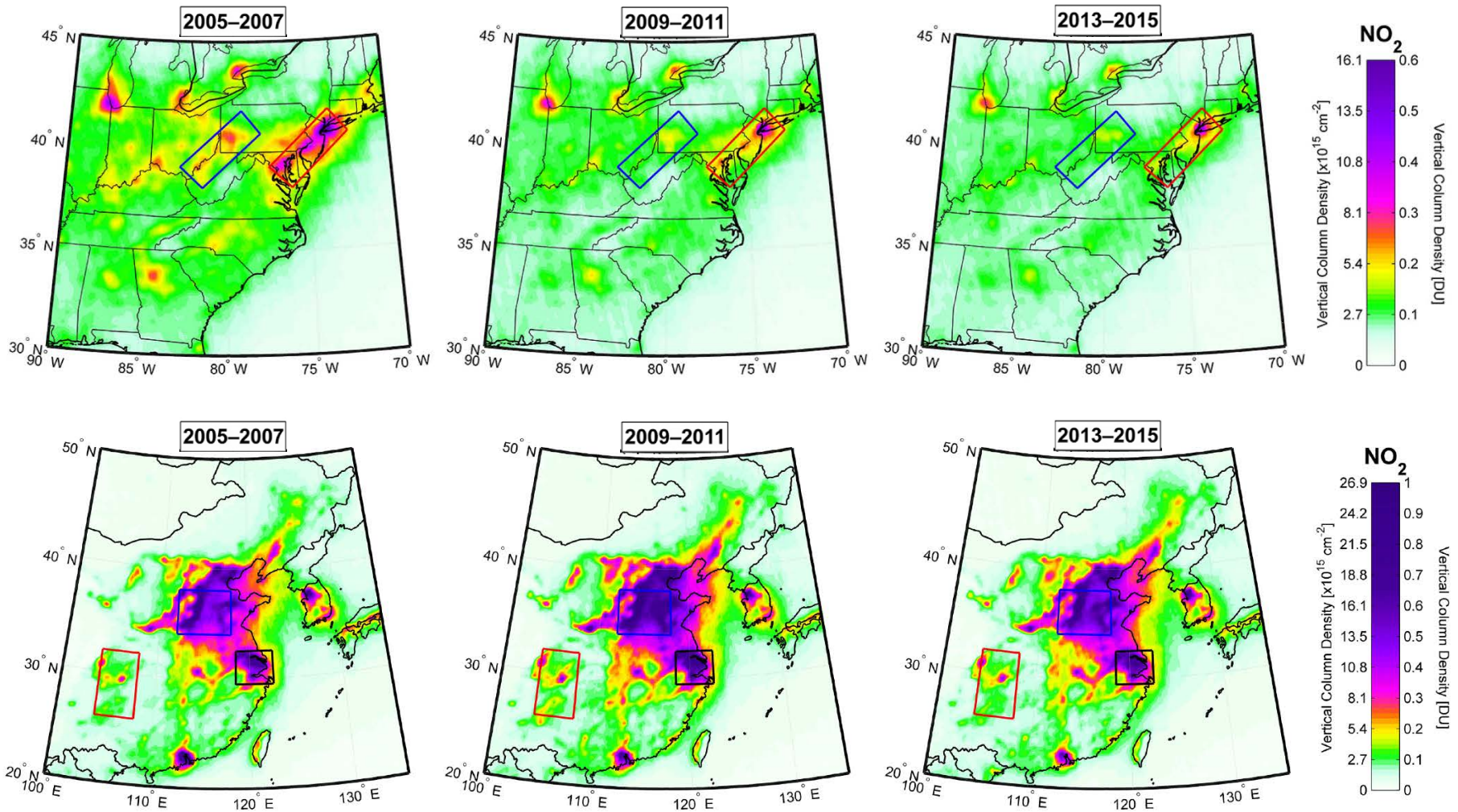
NCEP OPERATIONAL DATASET

<http://www.mde.state.md.us/assets/document/BJH%20-%20Basics%20on%20Ozone%20Transport.ppt>

# Dramatic Improvements Local Air Quality, Past 4 Decades



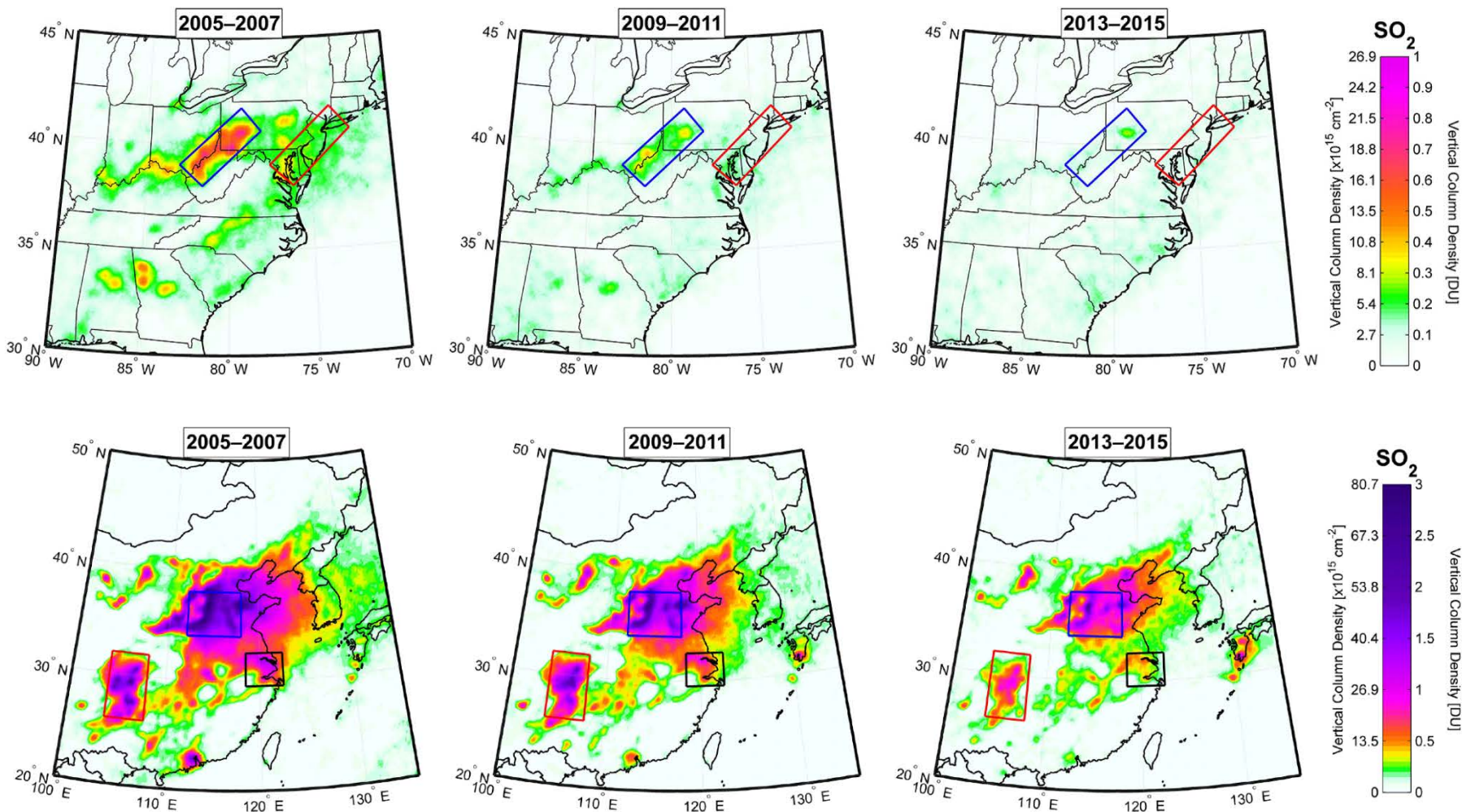
# NO<sub>2</sub> Trends from Space



Krotkov *et al.*, *ACP*, 2016



# SO<sub>2</sub> Trends from Space



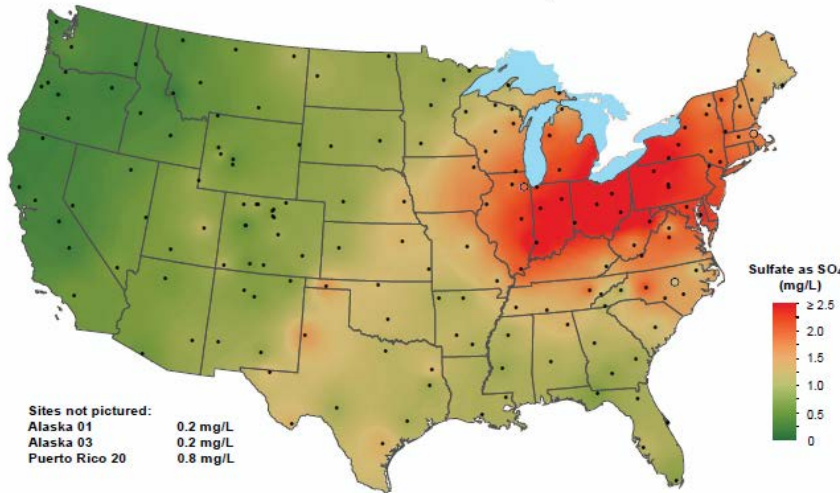
Krotkov *et al.*, ACP, 2016

# Sulfate Deposition

1994

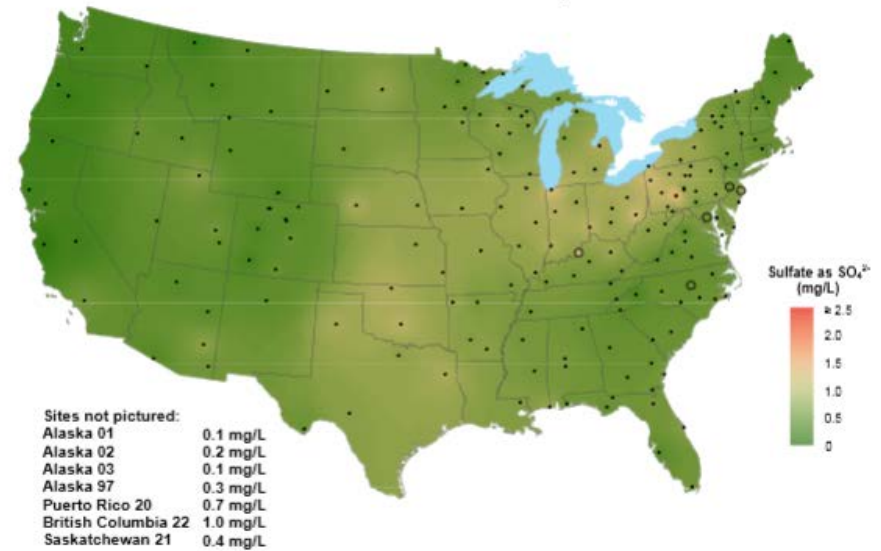
2013

Sulfate ion concentration, 1994



National Atmospheric Deposition Program/National Trends Network  
<http://nadp.isws.illinois.edu>

Sulfate ion concentration, 2013



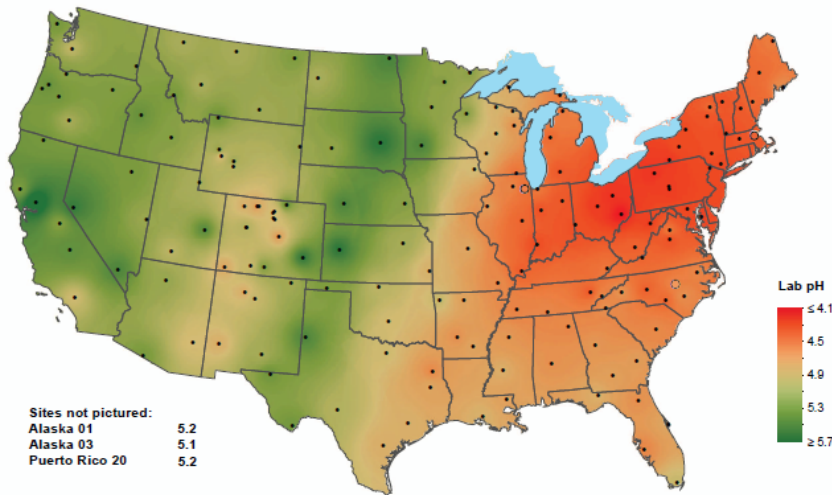
National Atmospheric Deposition Program/National Trends Network  
<http://nadp.isws.illinois.edu>

<http://nadp.sws.uiuc.edu/>

# pH

## 1994

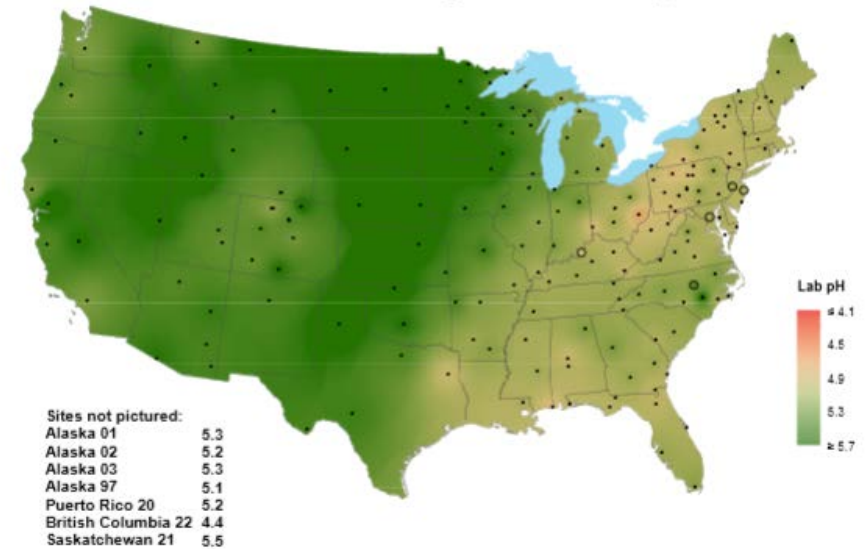
Hydrogen ion concentration as pH from measurements made at the Central Analytical Laboratory, 1994



National Atmospheric Deposition Program/National Trends Network  
<http://nadp.isws.illinois.edu>

## 2013

Hydrogen ion concentration as pH from measurements made at the Central Analytical Laboratory, 2013

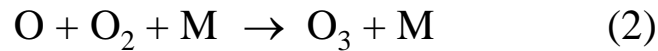
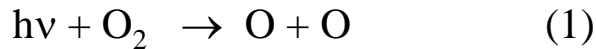


National Atmospheric Deposition Program/National Trends Network  
<http://nadp.isws.illinois.edu>

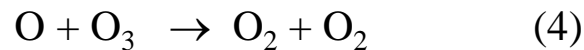
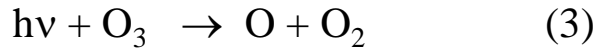
<http://nadp.sws.uiuc.edu/>

# Stratospheric Ozone: Chapman Chemistry

- Production of O<sub>3</sub> initiated when O<sub>2</sub> is photodissociated by UV sunlight
- O<sub>3</sub> formed when resulting O atom reacts with O<sub>2</sub> :



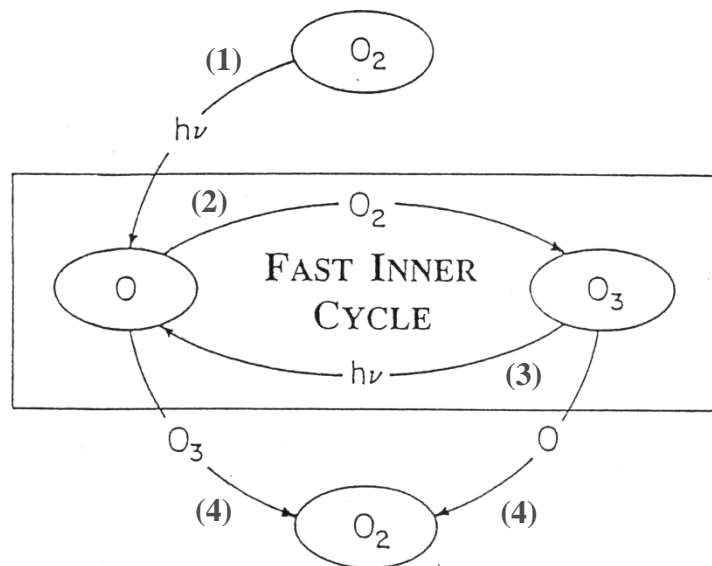
- O<sub>3</sub> removed by photodissociation (UV sunlight) or by reaction with O :



This reaction sequence was first worked out in the 1930s by Sidney Chapman, an English mathematician and geophysicist

# Chapman Chemistry

- The cycling between O and O<sub>2</sub> (rxns 2 and 3) occurs *much* more rapidly than leakage into (rxn 1) or out of the system (rxn 4)
- The sum O + O<sub>3</sub> is commonly called “*odd oxygen*”



Rxn (1) produces two *odd oxygen* molecules

Rxn (4) consumes two *odd oxygen* molecules

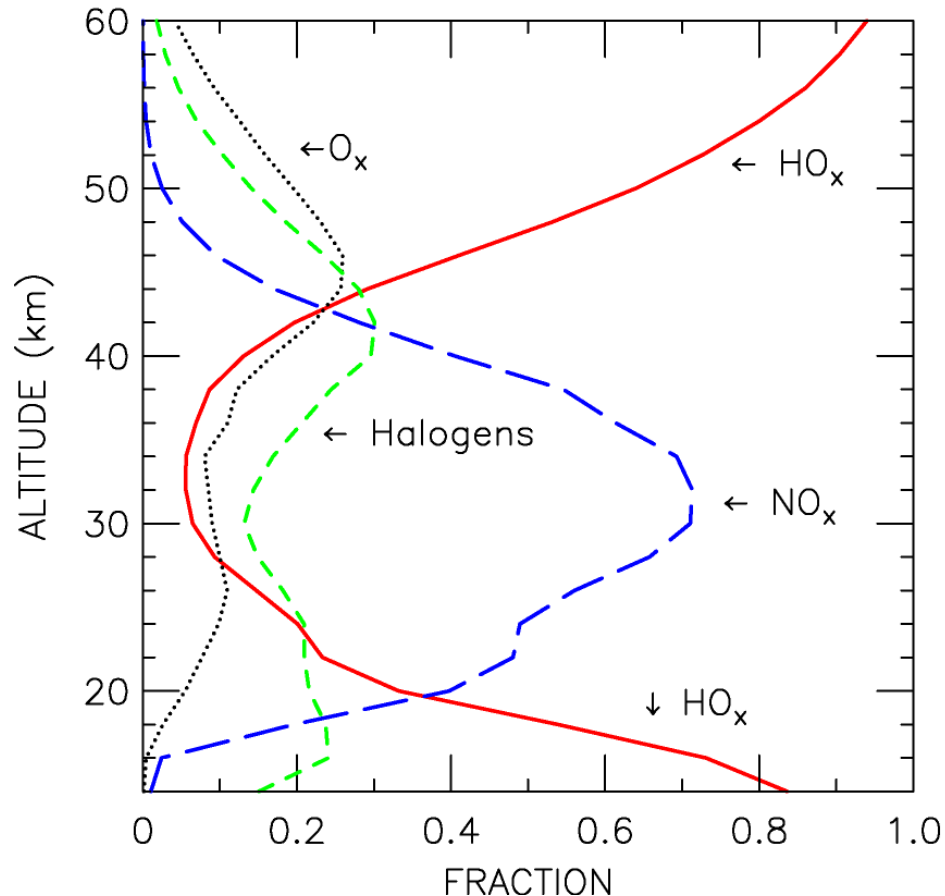
and reactions 2 and 3 recycle *odd oxygen* molecules

# Stratospheric Photochemistry: Odd Oxygen Loss By Families

Fraction of  $O_x$  Loss Due to Each Catalytic Family

JPL 2002 Kinetics

35°N, Sept



Calculated fraction of odd oxygen loss due to various families of radicals

After Osterman et al., GRL, 24, 1107, 1997;

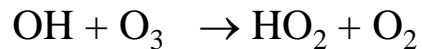
Sen et al., JGR, 103, 3571, 1998;

Sen et al., JGR, 104, 26653, 1999.

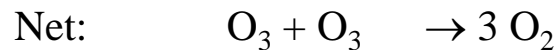
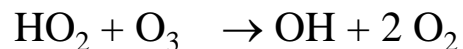
# One Atmosphere – One Photochemistry

## Stratosphere

HO<sub>2</sub> formation:

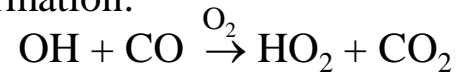


HO<sub>2</sub> loss:

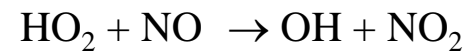


## Troposphere

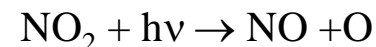
HO<sub>2</sub> formation:



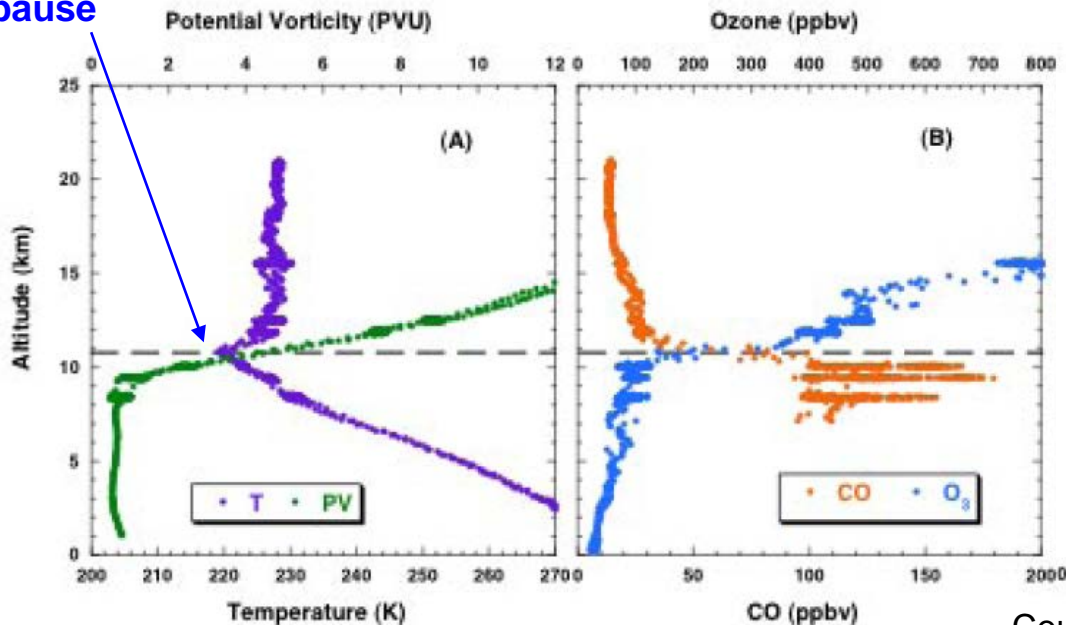
HO<sub>2</sub> loss:



Followed by:



Tropopause



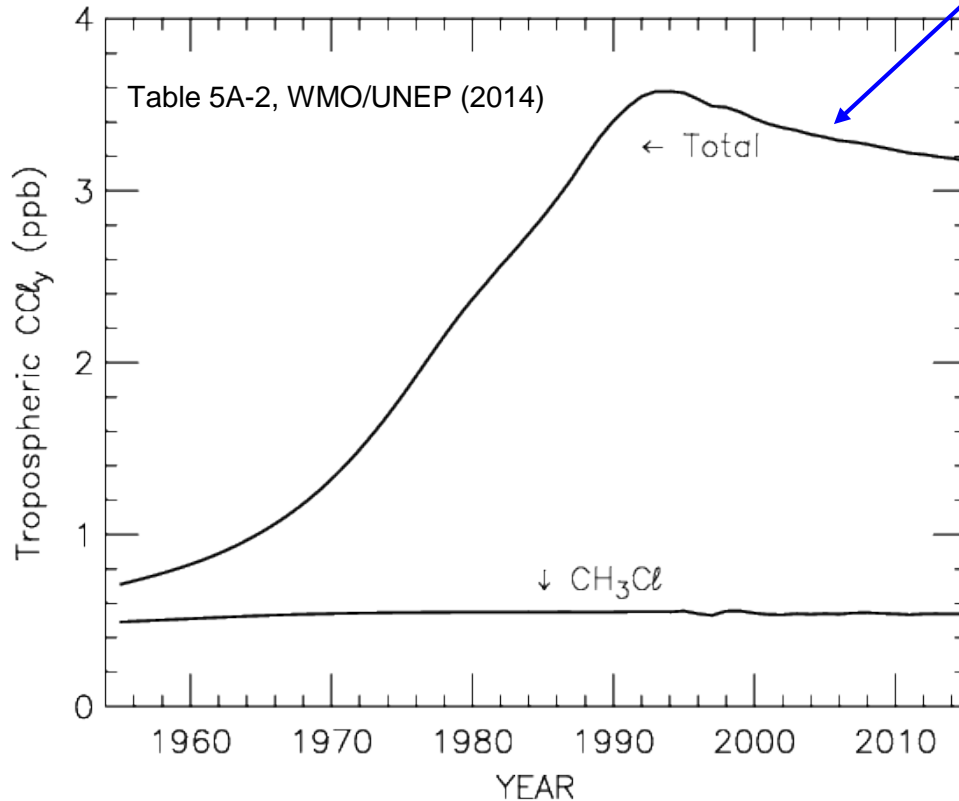
**Above Tropopause:**  
 Lots of O<sub>3</sub>, little CO  
**Below Tropopause:**  
 Lots of CO, little O<sub>3</sub>

Courtesy of Laura Pan, NCAR

# Tropospheric Chlorine Loading

## Total Organic Chlorine ( $\text{CCl}_y$ ):

- Peaked at ~3.6 ppb around 1993
- Slowly declining
- Montreal Protocol and Amendments have banned production of CFCs

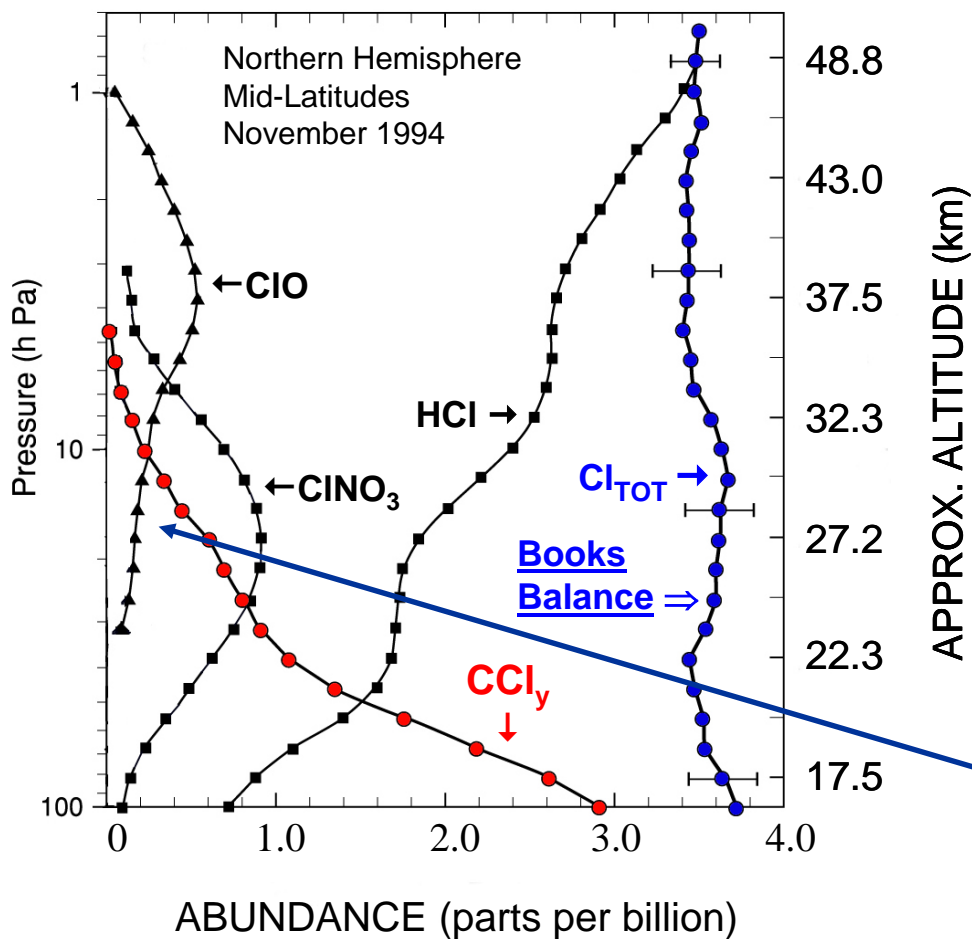


## CFCs:

- long lived (50 to 100 yr lifetime)
- decompose in the stratosphere
- lose memory of emission location when enter stratosphere



# Chlorine Abundance, Mid-Latitude Stratosphere

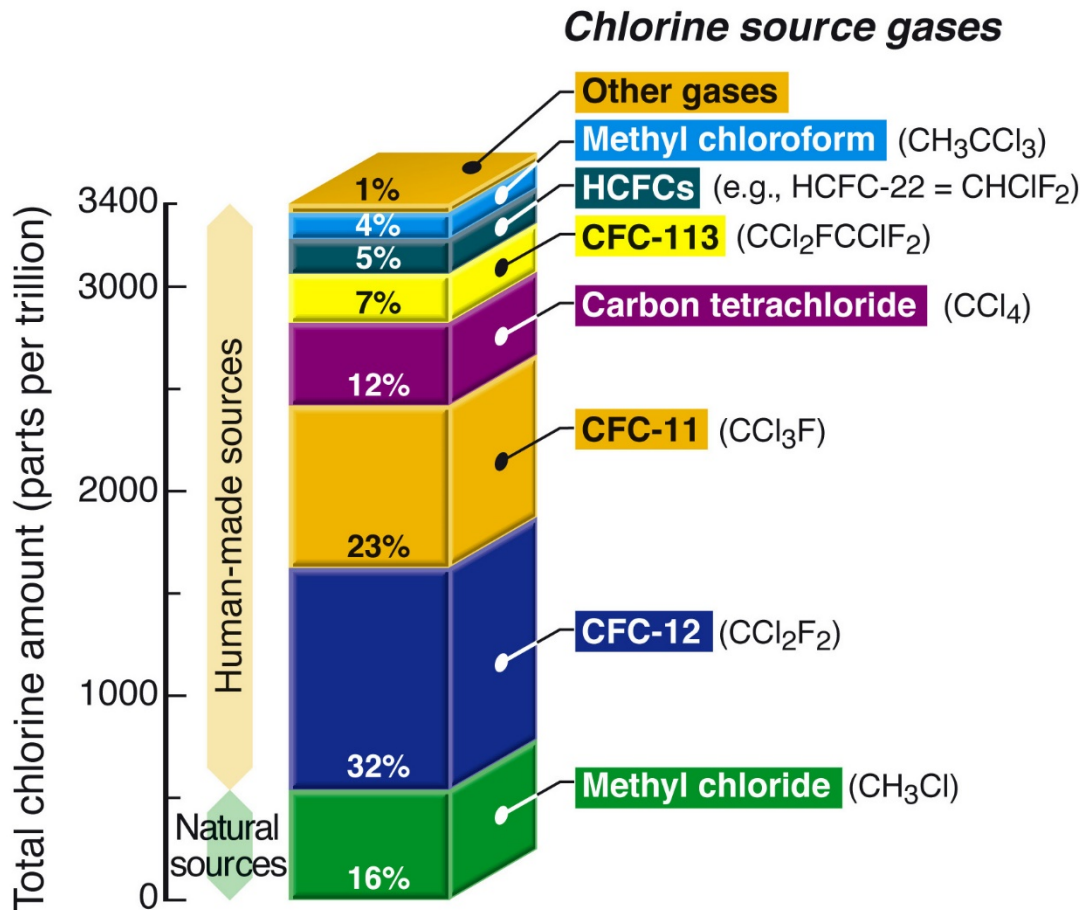


Note: Below ~30 km,  
 $\text{ClO} \ll \text{ClNO}_3$  and  $\text{HCl}$

Zander *et al.*, *GRL*, 1996

# Chlorine Source Gases

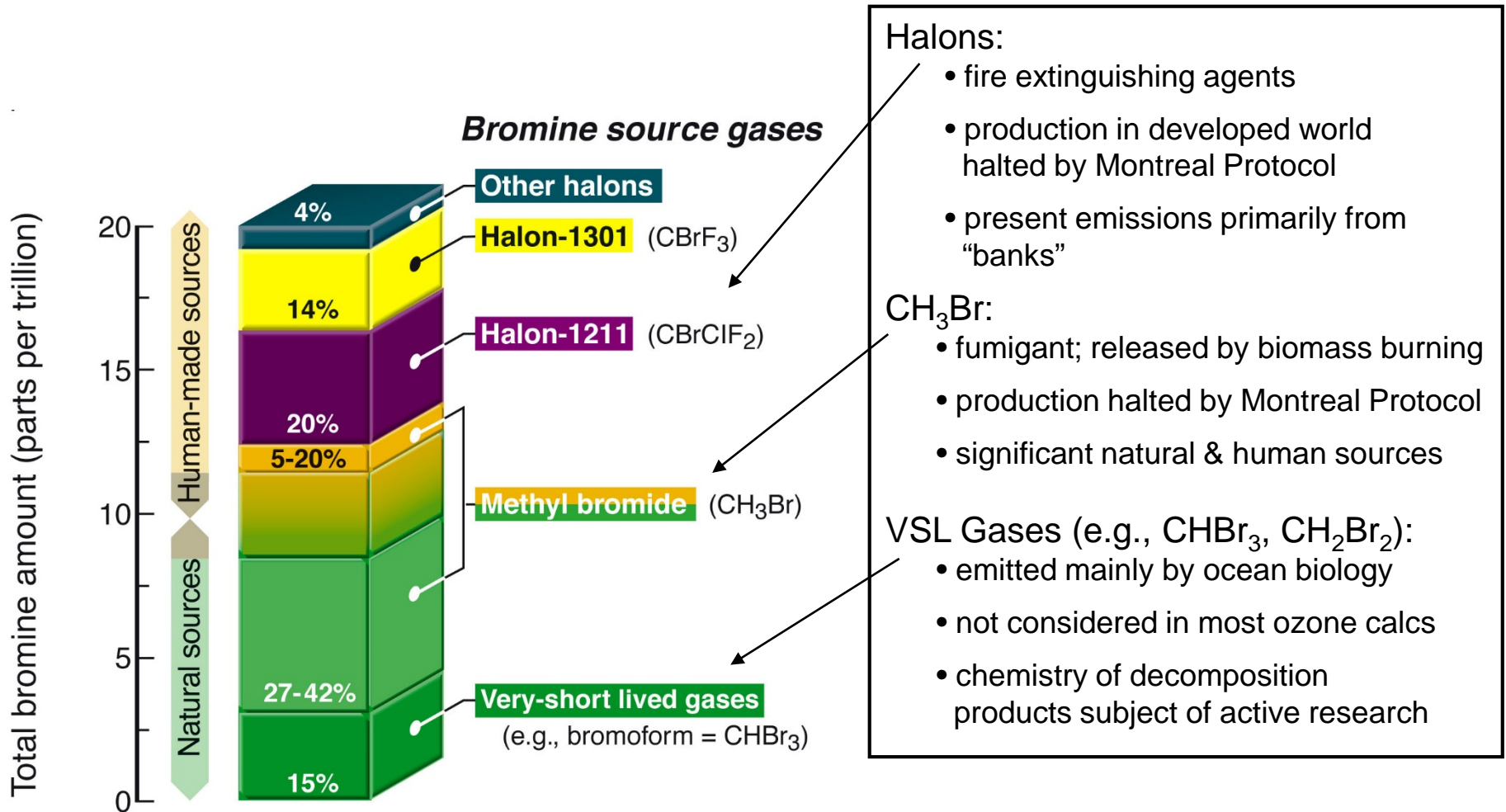
## Primary Sources of Chlorine for the Stratosphere in 1999



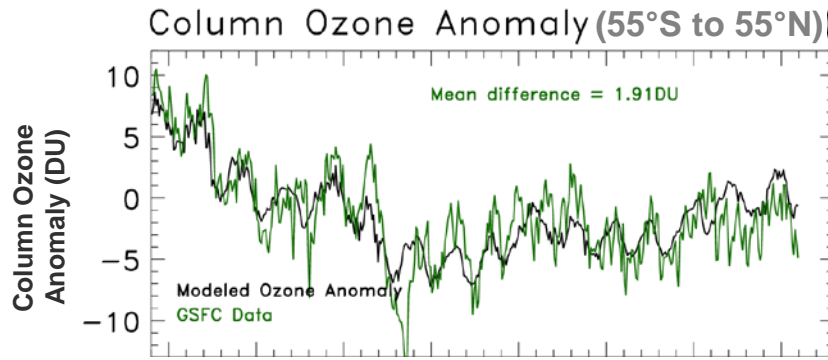
**Table Q7-1. Atmospheric Lifetimes and Ozone Depletion Potentials of some halogen source & HFC substitute gases.**

Gas	Atmospheric Lifetime (years)	Ozone Depletion Potential (ODP) <sup>c</sup>
<b>Halogen source gases</b>		
<i>Chlorine gases</i>		
CFC-11	45	1
CFC-12	100	0.82
CFC-113	85	0.85
Carbon tetrachloride ( $\text{CCl}_4$ )	26	0.82
HCFCs	1–17	0.01–0.12
Methyl chloroform ( $\text{CH}_3\text{CCl}_3$ )	5	0.16
Methyl chloride ( $\text{CH}_3\text{Cl}$ )	1	0.02
<i>Bromine gases</i>		
Halon-1301	65	15.9
Halon-1211	16	7.9
Methyl bromide ( $\text{CH}_3\text{Br}$ )	0.8	0.66
Very short-lived gases (e.g., $\text{CHBr}_3$ )	Less than 0.5	<sup>b</sup> very low
<b>Hydrofluorocarbons (HFCs)</b>		
HFC-134a	13.4	0
HFC-23	222	0

# Bromine Source Gases

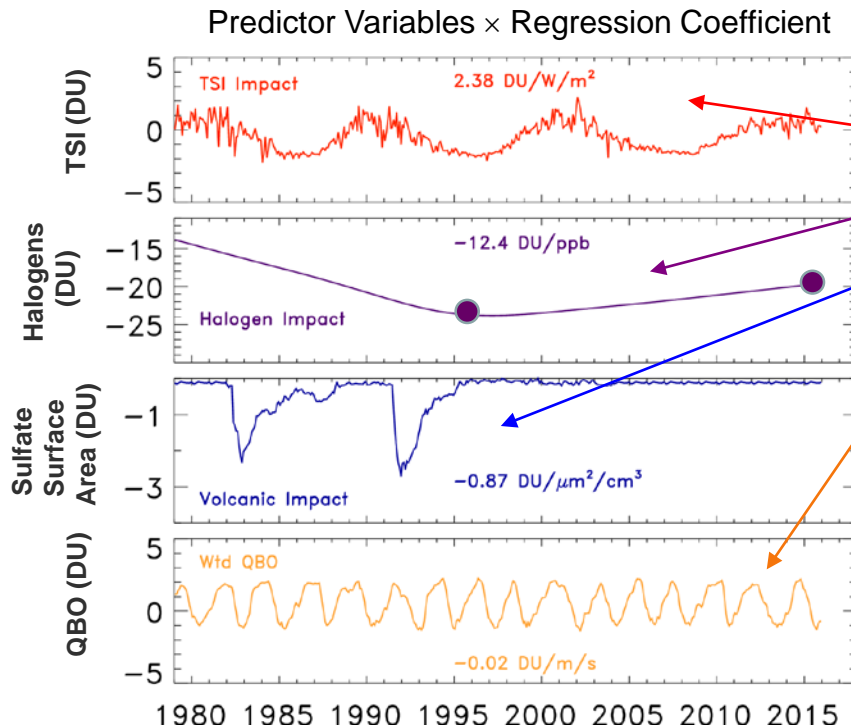


# Ozone Depletion at Mid-Latitudes



Ozone data from

[http://acdb-ext.gsfc.nasa.gov/Data\\_services/merged](http://acdb-ext.gsfc.nasa.gov/Data_services/merged)



Column Ozone Anomaly (DU) =

$$19.6 \text{ DU} + 2.38 \text{ DU} / \text{W m}^{-2} \times \text{TSI} + -12.4 \text{ DU} / \text{ppb} \times \text{Halogens} + -0.87 \text{ DU} \times \ln(\text{SSA}) + -0.02 \text{ DU} / \text{m s}^{-1} \times \text{QBO}$$

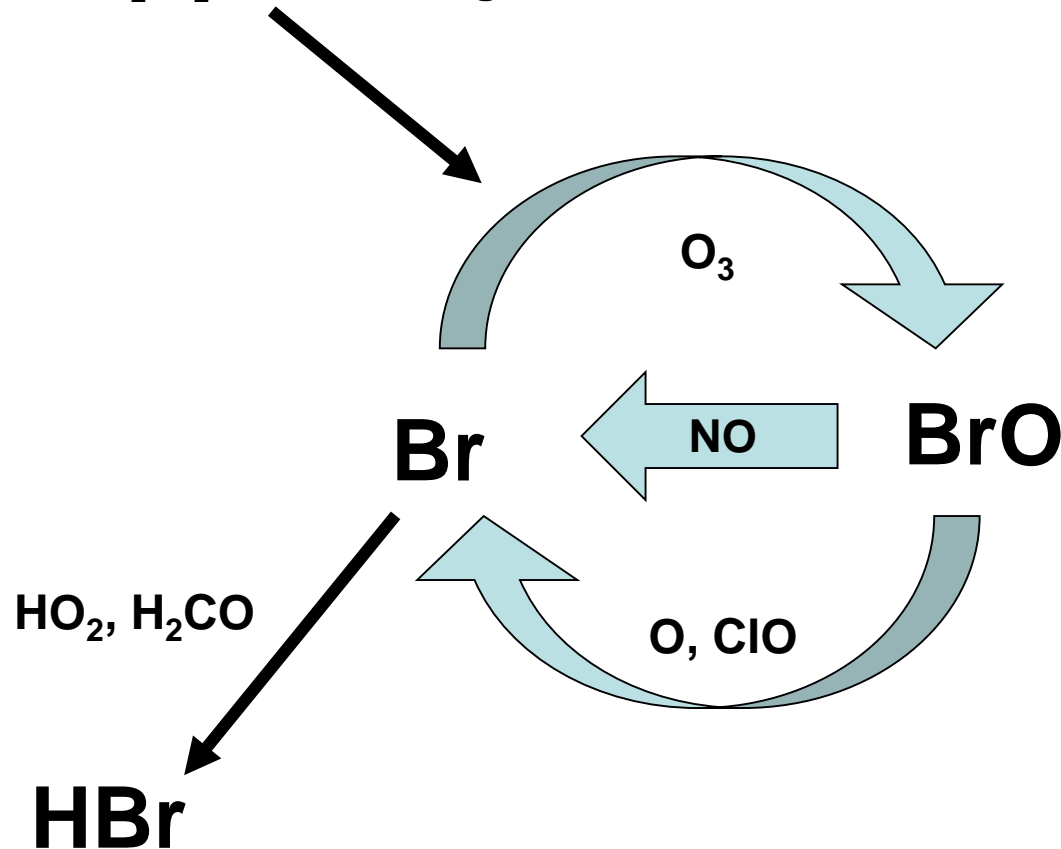
where

- TSI = total solar irradiance
- Halogens = stratospheric chlorine & bromine loading
- SSA = Sulfate Surface Area
- QBO = Quasi-biennial oscillation of the direction of winds in the tropical lower strat

# BrO<sub>x</sub> : BrO and Br

BrO is central to stratospheric photochemistry, at mid-latitudes and polar regions

**Production :** CH<sub>3</sub>Br + hν → Inorganic bromine } Considered in global models  
Halons + hν → Inorganic bromine }  
CHBr<sub>3</sub> + hν → Inorganic bromine } Generally neglected in global models  
CH<sub>2</sub>Br<sub>2</sub> + hν → Inorganic bromine }



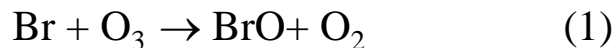
**Final sink : HBr solubility & rainout (lowermost stratosphere)**

# BrO<sub>x</sub> : BrO and Br

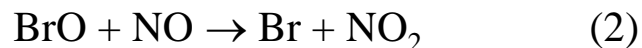
BrO is central to stratospheric photochemistry, at mid-latitudes and polar regions:

Rapid inner cycle:

BrO formation:



BrO loss:

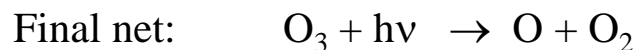
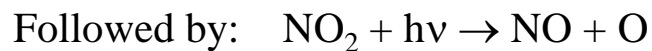
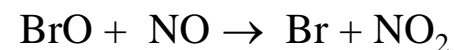
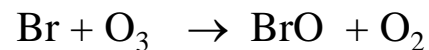


Can show:

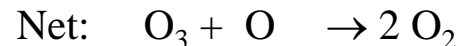
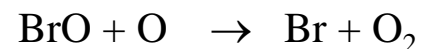
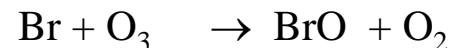
$$\frac{d\text{O}_3}{dt} + \frac{d\text{O}}{dt} = \frac{d(\text{Odd Oxygen})}{dt} = -2 k_3 [\text{BrO}][\text{O}]$$

As a convenient short hand, **we consider BrO to be odd oxygen**

BrO loss step (2):



BrO loss step (3):



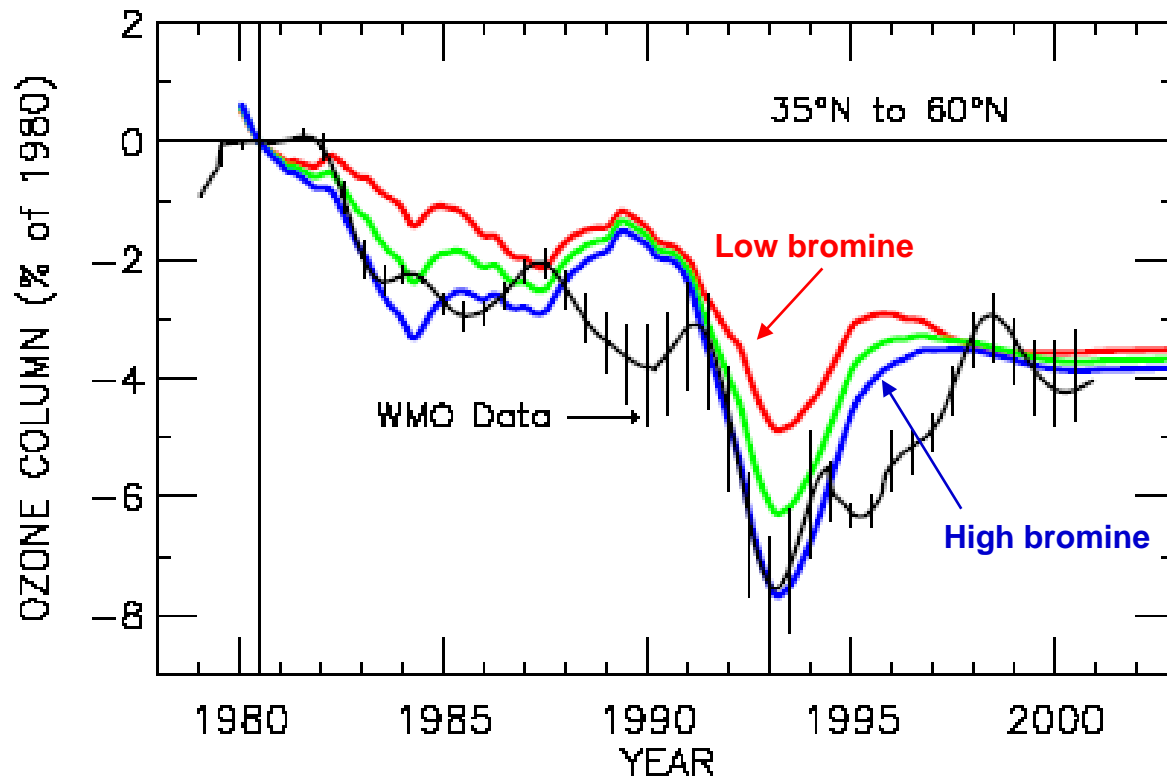
# Bromine Overview

Bromine is more efficient, per molecule, than chlorine at removing ozone because:

- Organic bromine compounds degrade faster than CFCs
  - lifetime of  $\text{CH}_3\text{Br}$  much shorter than lifetime of CFCs
  - Therefore  $\text{Br}_y$  liberated more easily from  $\text{CBr}_y$  than  $\text{Cl}_y$  liberated from  $\text{CCl}_y$
- $\text{BrO}/\text{Br}_y \gg \text{ClO}/\text{Cl}_y$ 
  - $\text{Br} + \text{CH}_4 \rightarrow \text{HBr} + \text{CH}_3$  is endothermic and has rate constant of zero
  - $\text{Cl} + \text{CH}_4 \rightarrow \text{HCl} + \text{CH}_3$  is exothermic and proceeds at reasonable rate
  - Also,  $J_{\text{BrONO}_2} > J_{\text{ClONO}_2}$
- Loss due to  $\text{BrO} + \text{ClO}$  reaction generally counted as “bromine loss”, even though loss of ozone by this cycle depends on  $\text{Br}_y$  and  $\text{Cl}_y$

## Ozone responds to:

- rise and fall of chlorine
- volcanic perturbations to aerosol loading
- amount of bromine in lowermost stratosphere

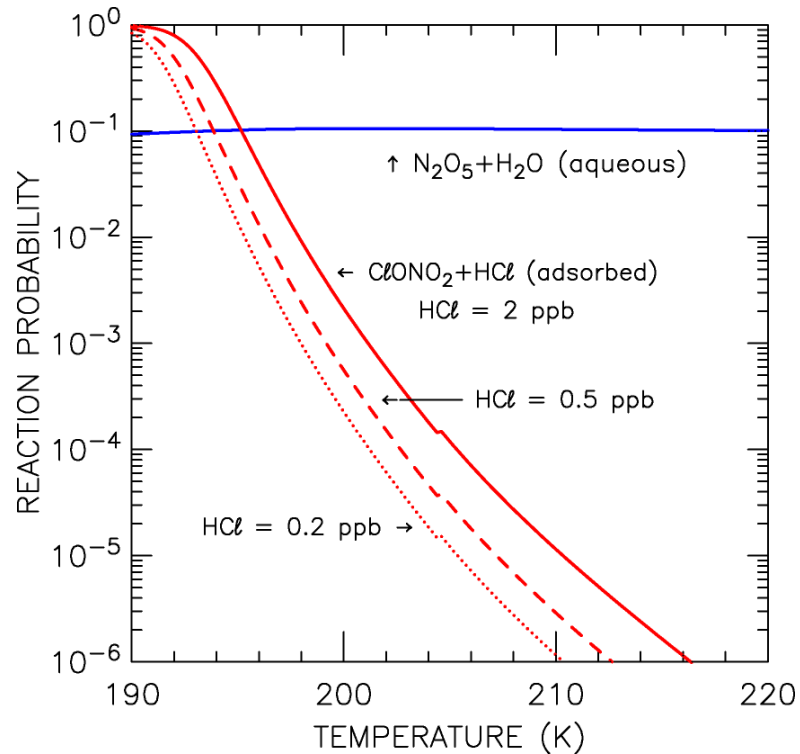


Salawitch *et al.*, *GRL*, 2004



# Heterogeneous Chemistry, Mid-Latitude vs Polar Regions

In all cases,  $\gamma$  must be measured in the laboratory



Reaction probabilities given for various surface types, with formulations of various degrees of complexity, in **Section 5** of the JPL Data Evaluation.

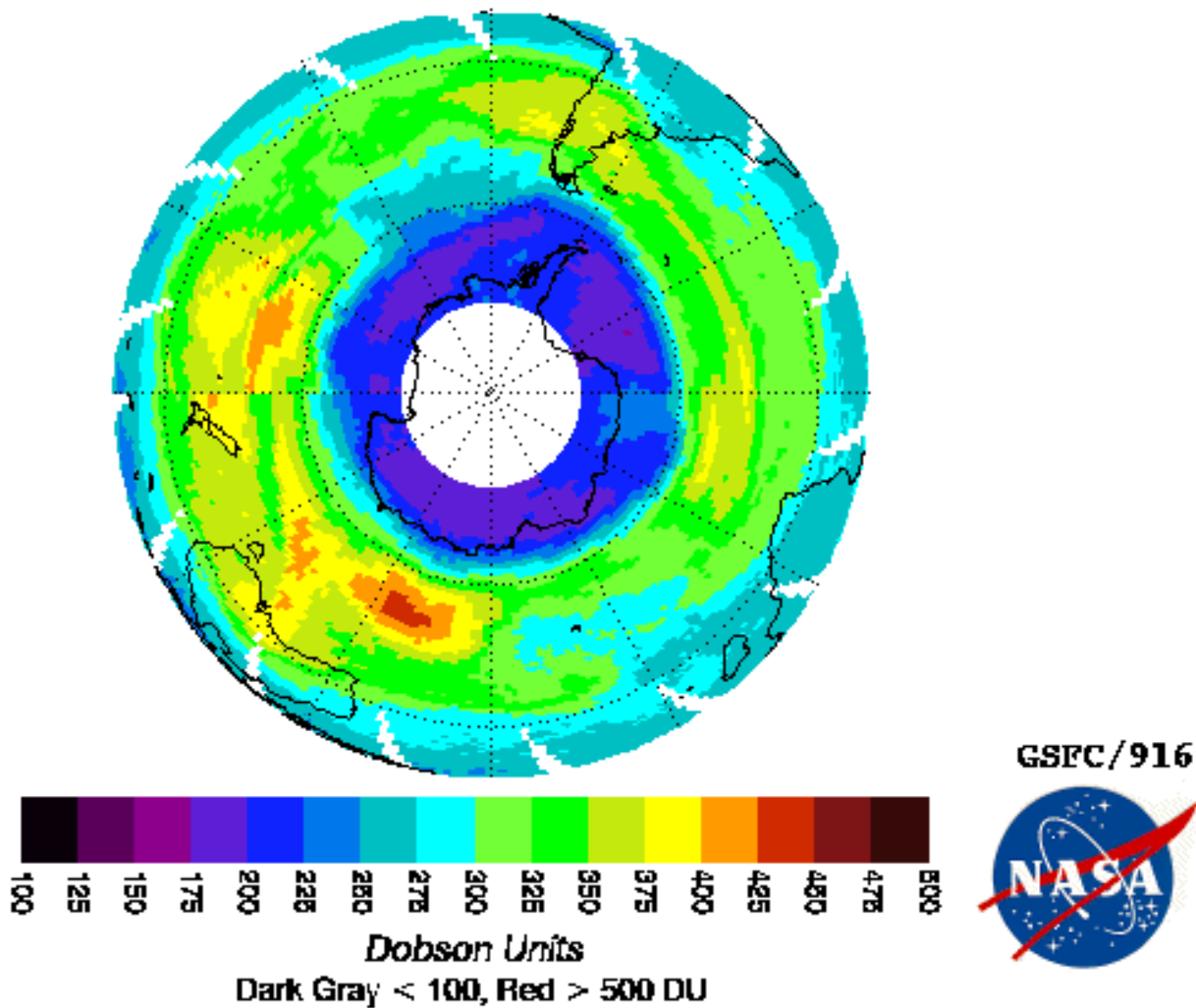
*Atmospheric Chemistry and Physics* by Seinfeld and Pandis provides extensive treatment of aqueous phase chemistry, properties of atmospheric aerosol, organic aerosols, etc.

# POLAR OZONE LOSS

- COLD TEMPERATURES → POLAR STRATOSPHERIC CLOUDS (PSCs)
- REACTIONS ON PSC SURFACES LEAD TO ELEVATED ClO
  - $\text{HCl} + \text{ClONO}_2 \rightarrow \text{Cl}_2 \text{ (gas)} + \text{HNO}_3 \text{ (solid)}$
  - $\text{ClONO}_2 + \text{H}_2\text{O} \rightarrow \text{HOCl} + \text{HNO}_3$
  - $\text{Cl}_2 + \text{SUNLIGHT} + \text{O}_3 \rightarrow \text{ClO}$
  - $\text{HOCl} + \text{SUNLIGHT} + \text{O}_3 \rightarrow \text{ClO}$
  - $\text{HNO}_3$  SEDIMENTS (PSCs fall due to gravity)
- ELEVATED **ClO** + SUNLIGHT DESTROYS  $\text{O}_3$
- **BrO** : REACTION PARTNER FOR **ClO** ⇒ ADDITIONAL  $\text{O}_3$  LOSS

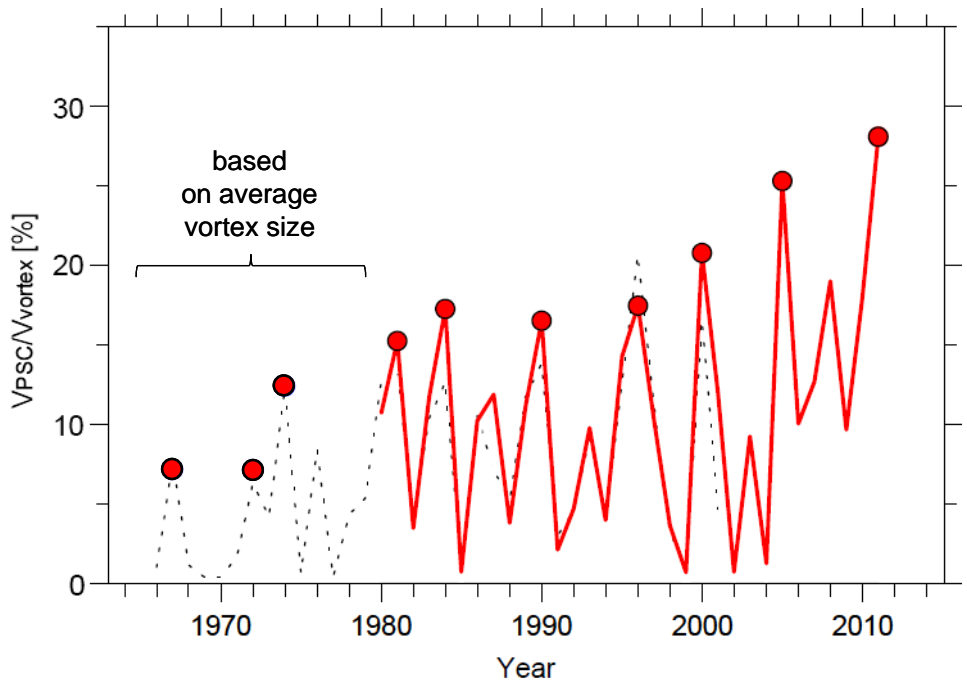
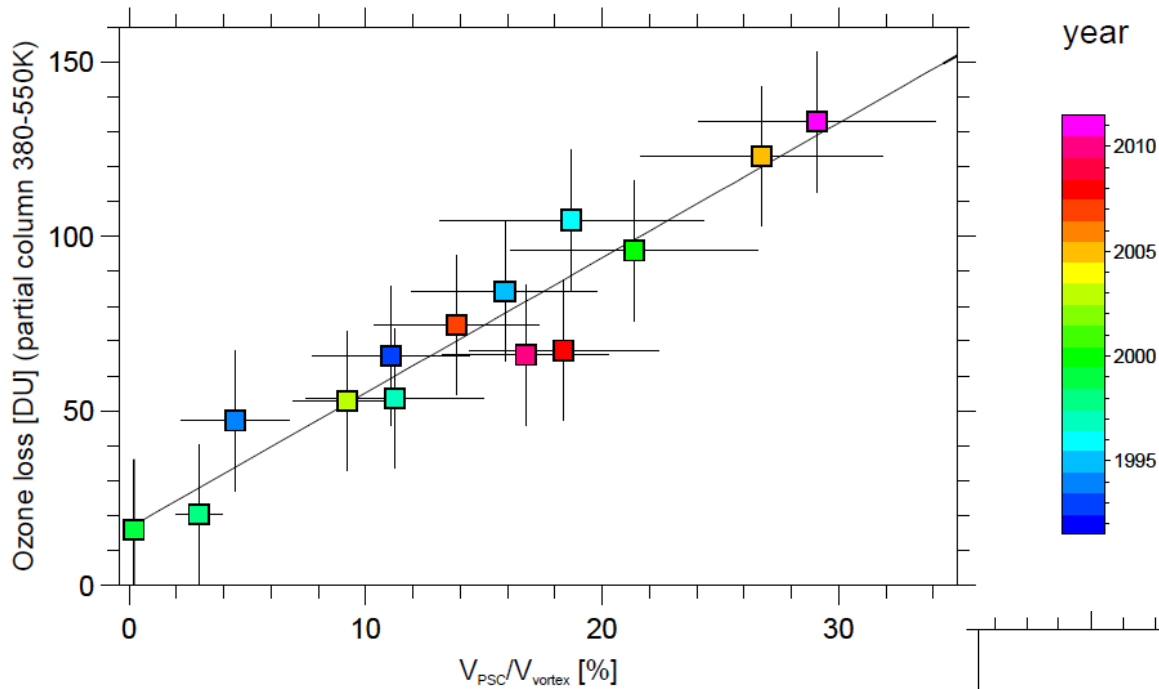


# EP/TOMS Total Ozone for Sep 1, 2001

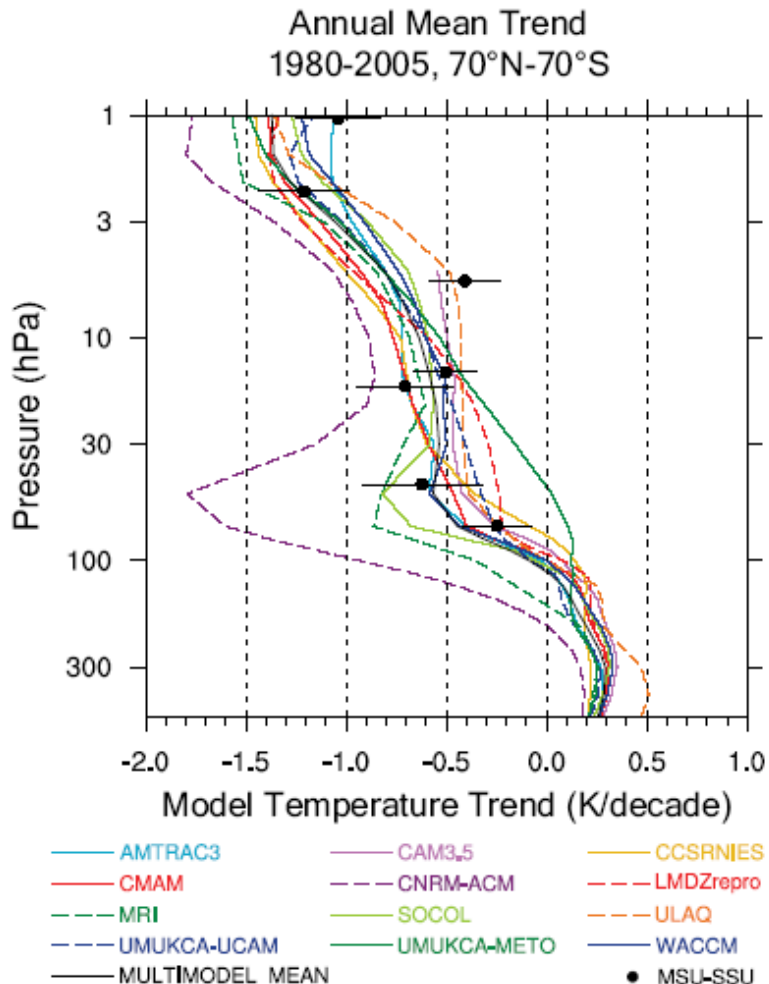


GEN:271:2001

# Arctic Ozone 2011 in Context of Prior Years



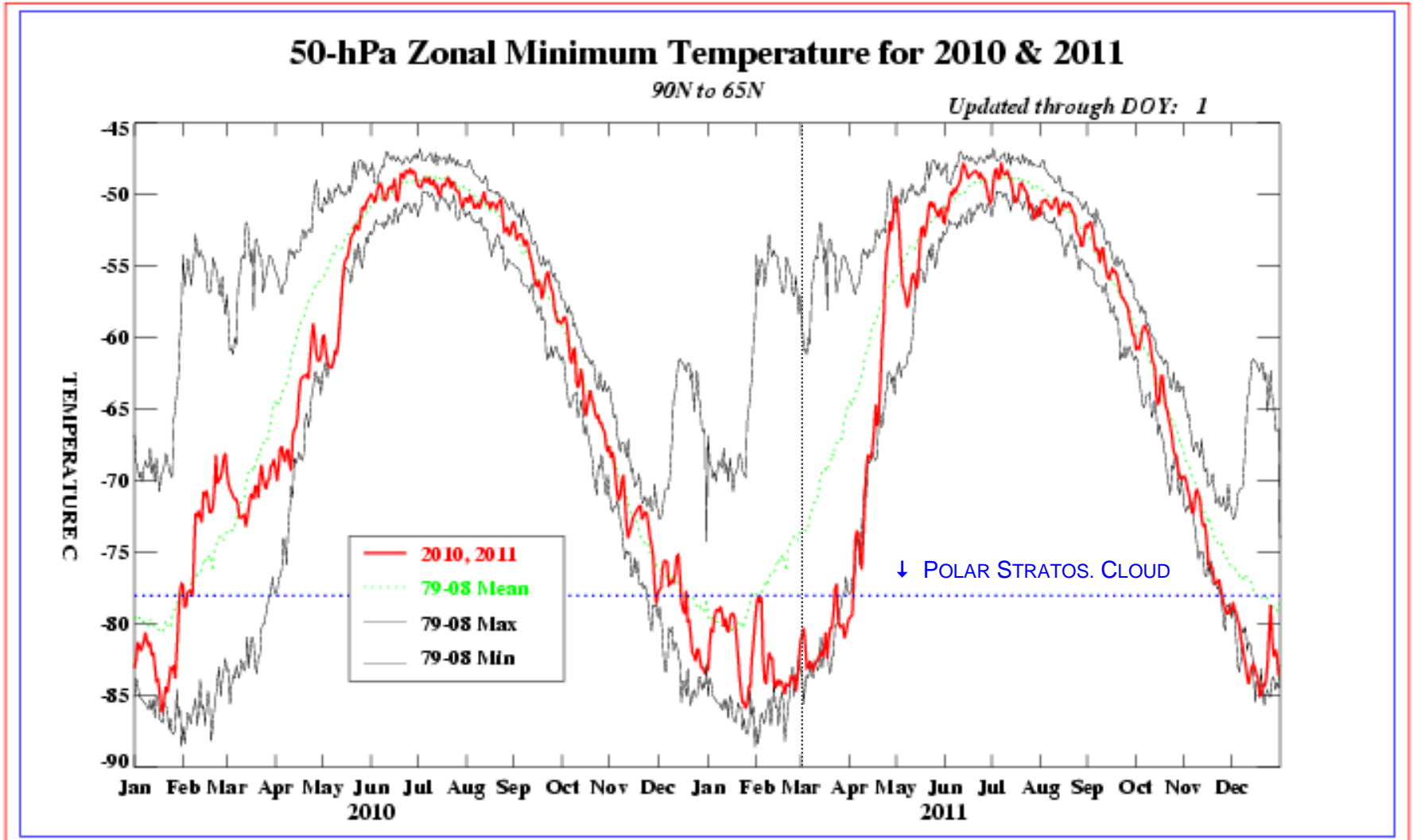
# The Stratosphere Cools as the Surface Warms



If the stratosphere continues to cool, for which region of the stratosphere will ozone be most vulnerable to future decline ?

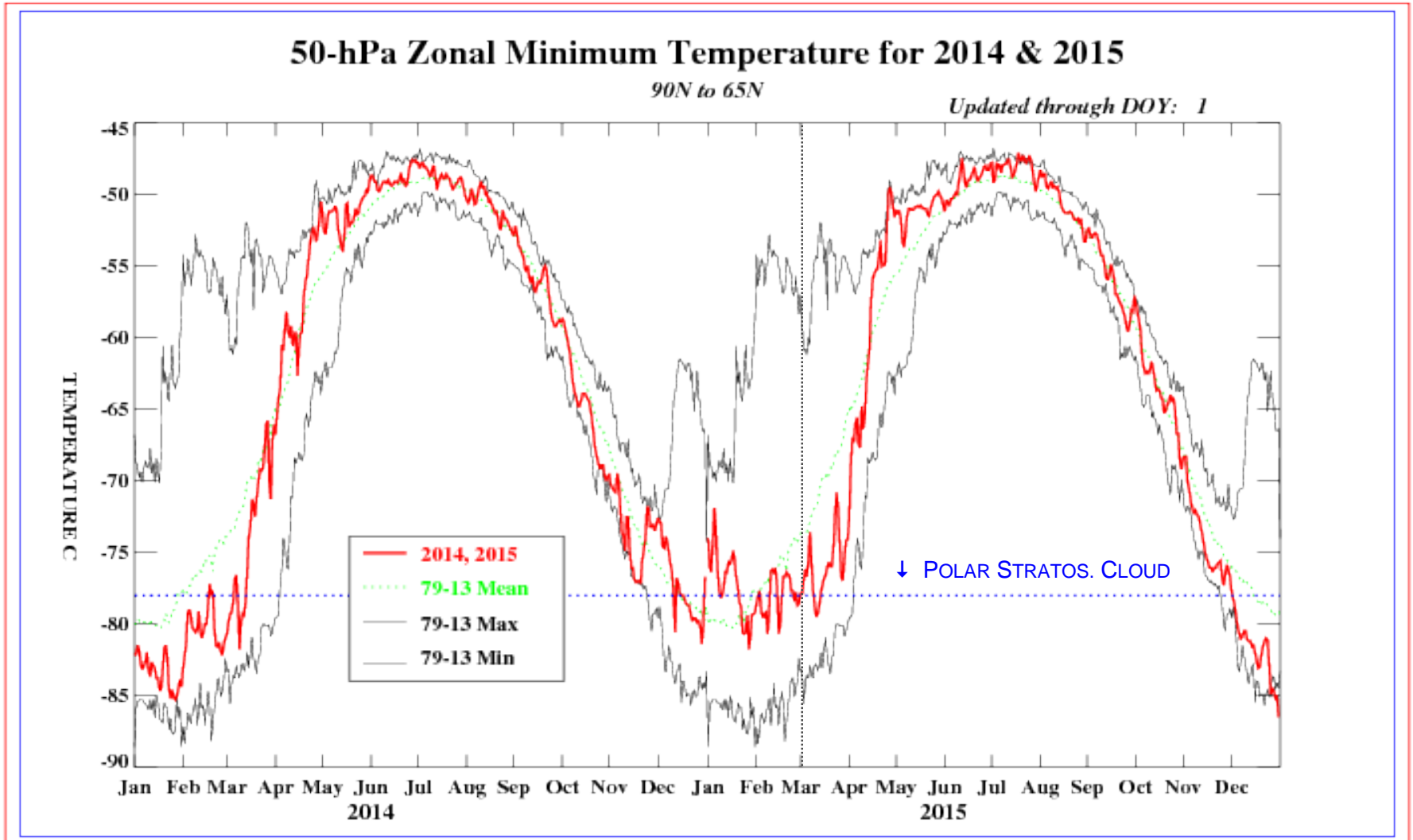
Figure 4-11, WMO/UNEP (2011)

# Arctic Temperature: Mar 2011



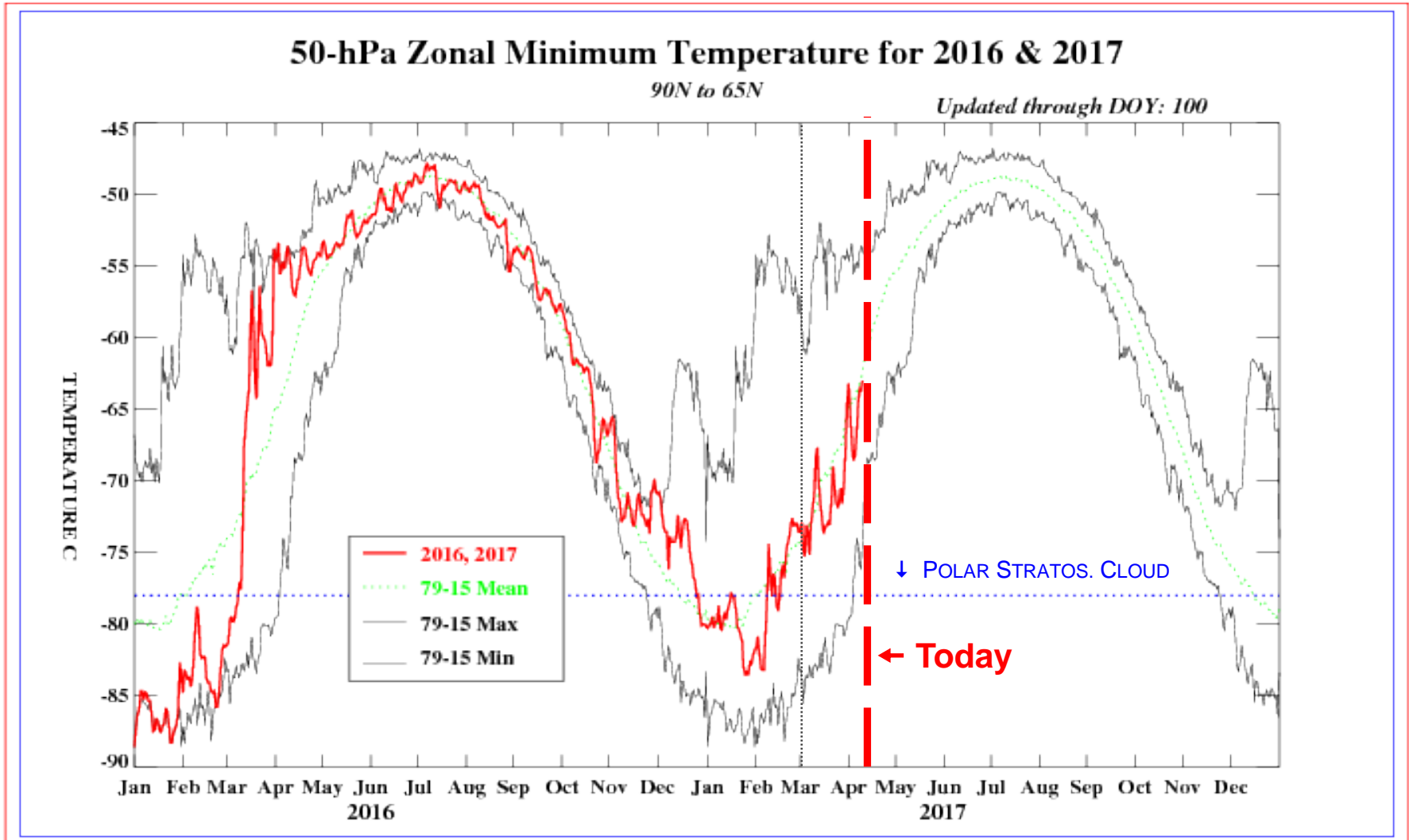
[http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/archive/50mbnhlo\\_2011.gif](http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/archive/50mbnhlo_2011.gif)

# Arctic Temperature: Mar 2015



[http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/archive/50mbnhlo\\_2015.gif](http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/archive/50mbnhlo_2015.gif)

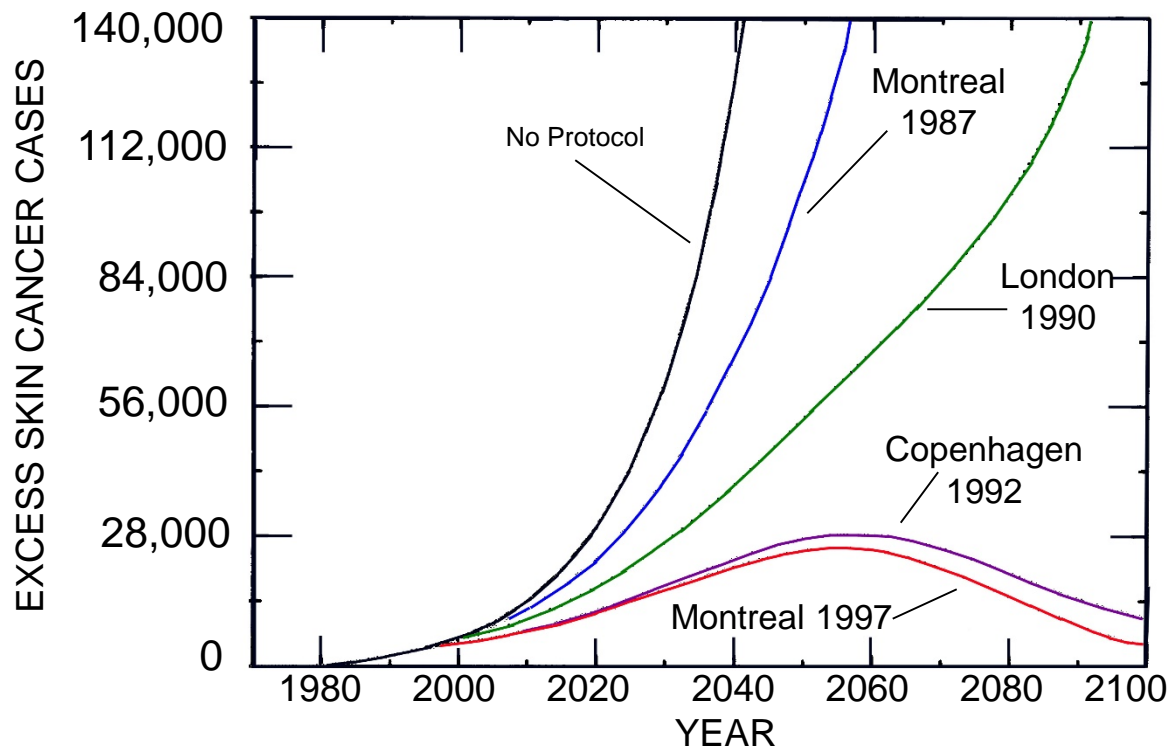
# Arctic Temperature: Mar 2017



<http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/50mbnhlo.png>



EXCESS SKIN CANCER CASES IN THE UNITED STATES,  
PER YEAR, DUE TO OZONE DEPLETION  
FOR VARIOUS CFC EMISSION SCENARIOS



Longstreth *et al.*, *J. of Photochemistry and Photobiology B*, 46, 20–39, 1998.

See also Slaper *et al.*, Estimates of ozone depletion and skin cancer incidence to examine the Vienna Convention achievements, *Nature*, 384, 256–258, 1996, who state:

The no-restrictions and Montreal Protocol scenarios produce a runaway increase in skin cancer incidence, up to a quadrupling and doubling, respectively, by year 2100.

# Second Exam

- Thursday, 13 April, 2:00 pm to 3:15 pm
- CSS 2416
- Closed book, no notes
  
- Focus mainly on Lectures 9 to 16
- Conceptual questions only: no calculator
- Backbone of course is the lectures and material from readings highlighted in class
- We'll be present: *please let us know if a question requires clarification*
- **Exam for 633 will differ somewhat from exam for 433**