

# Climates of the Past

## AOSC / CHEM 433 & AOSC 633

### Ross Salawitch & Walt Tribett

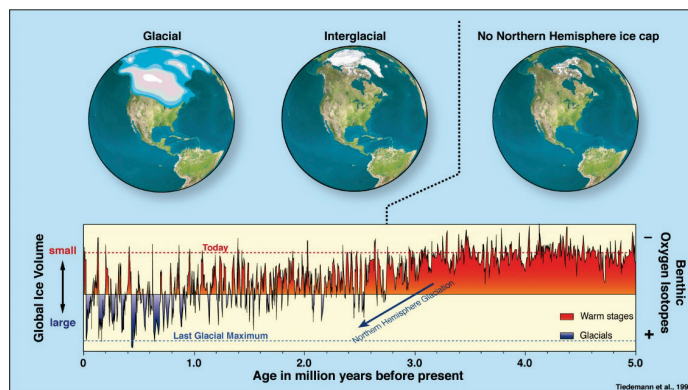
Class Web Sites:

<http://www.atmos.umd.edu/~rjs/class/spr2019>

<https://myelms.umd.edu/courses/1256337>

Email:

Ross: [rsalawit@umd.edu](mailto:rsalawit@umd.edu) or [rjs@atmos.umd.edu](mailto:rjs@atmos.umd.edu); Walt: [wtribett@umd.edu](mailto:wtribett@umd.edu)



Originally from [http://www.awi.de/en/research/research\\_divisions/geosciences/marine\\_geology\\_and\\_palaeontology](http://www.awi.de/en/research/research_divisions/geosciences/marine_geology_and_palaeontology)

Now at [https://silentwitness.files.wordpress.com/2012/08/klimakurve\\_webpage.jpg](https://silentwitness.files.wordpress.com/2012/08/klimakurve_webpage.jpg)

## Lecture 4

### 12 February 2019

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1

- a) textbook does indeed state 15°C
- b) STP (standard temperature and pressure) is actually 0°C (273.15 K)
- c) definition that 1 DU is 0.01 millimeter thickness of a gas is based upon evaluation at 1 atmosphere and 0°C

$$\begin{aligned}
 \text{Earth: } & \left\{ (1-0.3) \times \frac{1370 \text{ W m}^{-2}}{4} \times \frac{1}{5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}} \right\}^{1/4} = 255 \text{ K} \\
 \text{Venus: } & \left\{ (1-0.75) \times \frac{1370 \text{ W m}^{-2} \times \left( \frac{1}{0.72^2} \right)}{4} \times \frac{1}{5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}} \right\}^{1/4} = 232 \text{ K} \\
 \text{Mars: } & \left\{ (1-0.17) \times \frac{1370 \text{ W m}^{-2} \times \left( \frac{1}{1.5^2} \right)}{4} \times \frac{1}{5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}} \right\}^{1/4} = 217 \text{ K}
 \end{aligned}$$

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# Problem Set #1

AOSC / CHEM 433 & AOSC 633

Problem Set #1

Due: Tuesday, 19 February 2019 (at start of class)

*Late penalty: 10 points per day late, unless there is a legitimate medical or extra-curricular (i.e., band, athletics, GREs, etc) circumstance brought to our attention prior to the due date!*

## 1. Effective Temperature (433: 35 points; 633: 45 points).

Calculate the effective temperature of the Earth for the following cases. *Please show your work.*

- (5 points) Current solar conditions (i.e., use the value of  $S$  given in class) and an albedo = 0.3 that we'll assume applies at the start of the sea-ice data record.
- (10 points) As the Earth warms, snow and ice will melt. Indeed, scientists have reported a rather precipitous drop in so-called perennial sea ice in the Arctic ocean, as shown below:

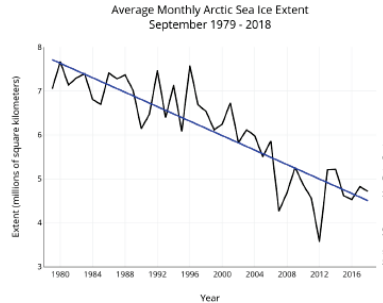


Figure from <http://nsidc.org/arcticseaicenews/2018/10/september-extent-ties-for-sixth-lowest>

The reflectivity (albedo) of the open ocean is 0.06, whereas the reflectivity of sea ice is about 0.6: <https://nsidc.org/cryosphere/seaice/processes/albedo.html>

Using the radius of Earth given in Lectures 1 and 3, find:

- the surface area of Earth in units of  $\text{km}^2$
- how much Earth's albedo would have changed, from the beginning of the Arctic Sea Ice Extent record (i.e., earliest part of blue line) to the end of this record (latest part of the blue line), based on loss of sea ice. In your reply, state whether this change is either a rise or a fall in albedo. Also, for this calculation, assume Earth's albedo was 0.30 at the start of the sea-ice data record shown above.

[http://www.atmos.umd.edu/~rjs/class/spr2019/problem\\_sets/ACC\\_2019\\_ps01.pdf](http://www.atmos.umd.edu/~rjs/class/spr2019/problem_sets/ACC_2019_ps01.pdf)

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## Climates of the Past

### Overview:

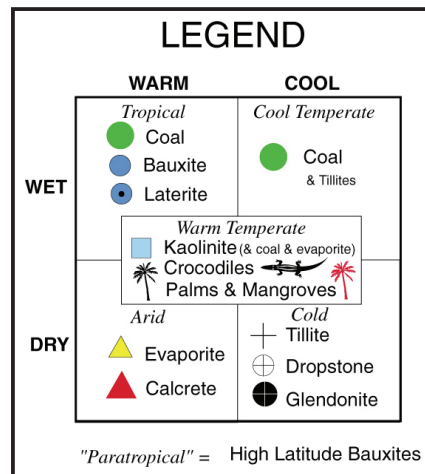
- 1) Techniques for quantifying past climate
- 2) Remarkable changes in past climate
- 3) Challenge in applying past climate sensitivity to future climate

The details of this "challenge" are quantitative and come at end of lecture.

I generally do not like to place quantitative material at the end of lecture;

please bear with me today as this arrangement seems best way to organize material.

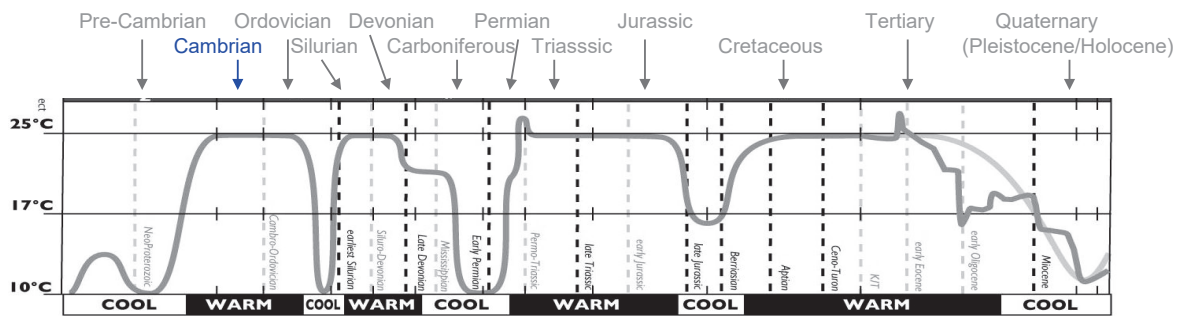
Legend for slides to follow →



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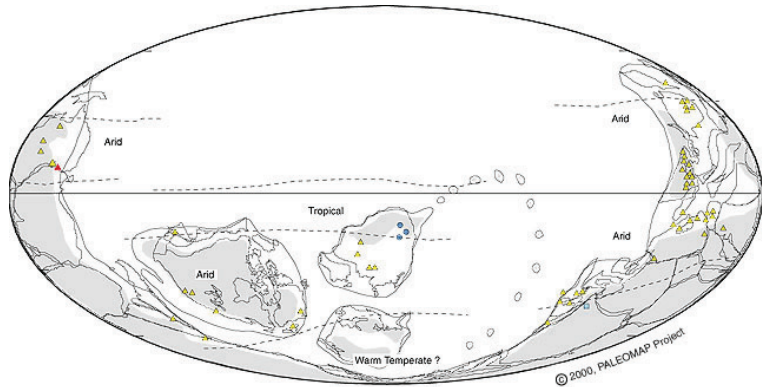


<http://www.scotese.com>

<http://pubs.usgs.gov/gip/fossils/fig15.gif>

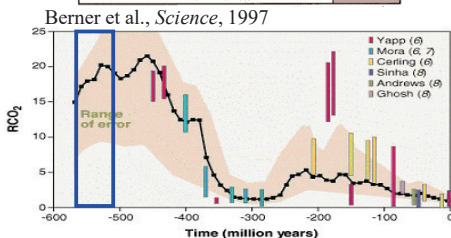
EON	ERA	PERIOD	MILLIONS OF YEARS AGO
Phanerozoic	Cenozoic	Quaternary	1.6
		Tertiary	66
	Mesozoic	Cretaceous	138
		Jurassic	205
		Triassic	240
		Permian	290
	Paleozoic	Pennsylvanian	330
		Mississippian	360
		Devonian	410
		Silurian	435
		Ordovician	500
Proterozoic	Paleoproterozoic	Cambrian	570
		Early Cambrian	570
		Early Proterozoic	2500
Archean	Archean	Early Archean	3800?
		Pre-Archean	3800?

## Early Cambrian Climate (540 million years ago)



The climate of the Cambrian is not well known. It was probably not very hot, nor very cold. There is no evidence of ice at the poles.

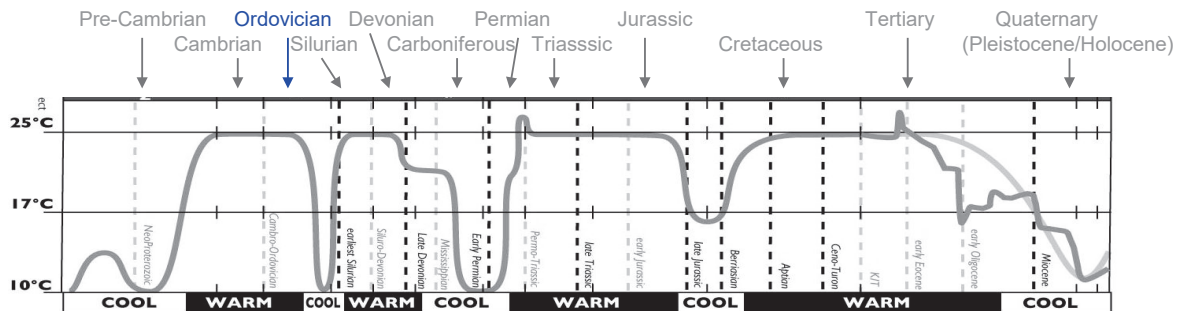
Source: <http://www.scotese.com/ecambcli.htm>



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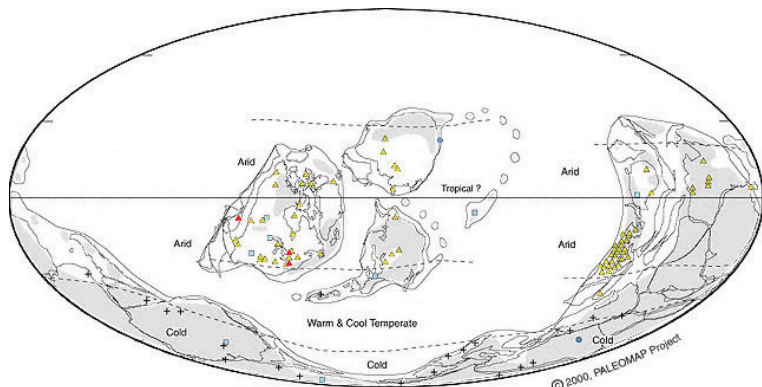


<http://www.scotese.com>

<http://pubs.usgs.gov/gip/fossils/fig15.gif>

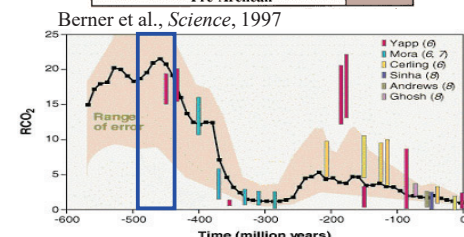
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		Early Cambrian	570
		Early Proterozoic	2500
Archean	Archean	Early Archean	3800?
		Pre-Archean	3800?

## Early Ordovician Climate (480 million years ago)



Mild climates probably covered most of the globe. The continents were flooded by the oceans creating warm, broad tropical seaways.

Source: <http://www.scotese.com/eordclim.htm>

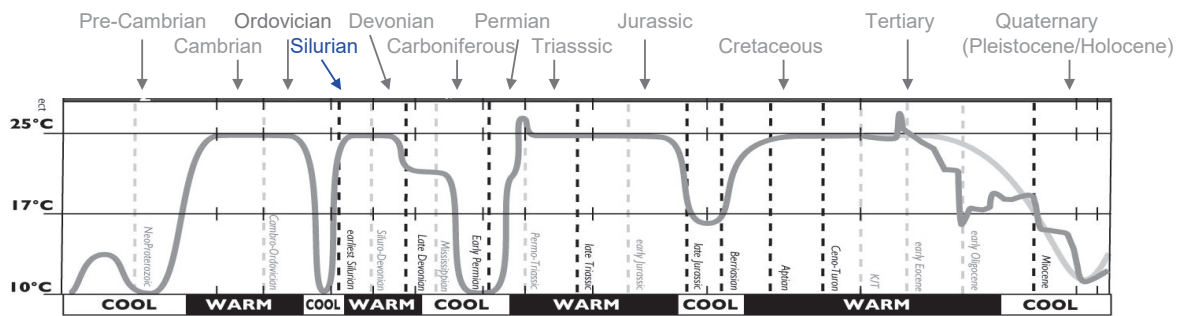


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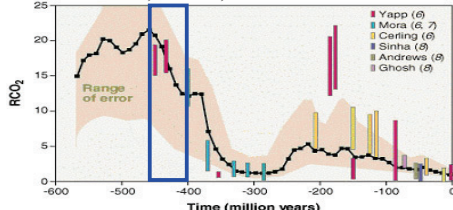


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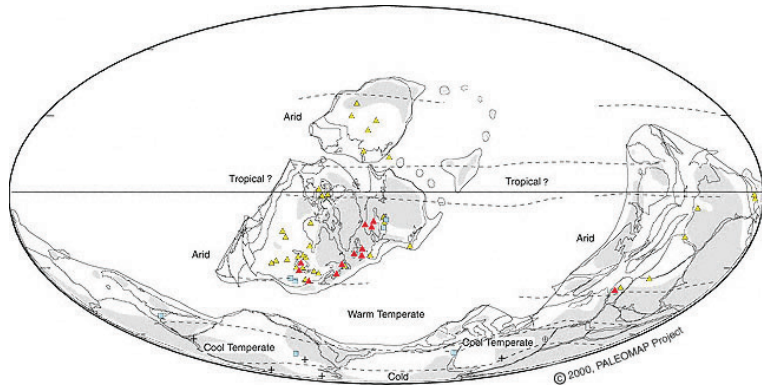
<http://pubs.usgs.gov/gip/fossils/fig15.gif>

EON	ERA	PERIOD	MILLIONS OF YEARS AGO
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		Ordovician	500
Proterozoic	Archean	Cambrian	570
		Late Proterozoic	2500
		Middle Proterozoic	2500
Pre-Archean		Early Proterozoic	3800?
		Pre-Archean	3800?

Berner et al., *Science*, 1997



## Silurian Climate (420 million years ago)



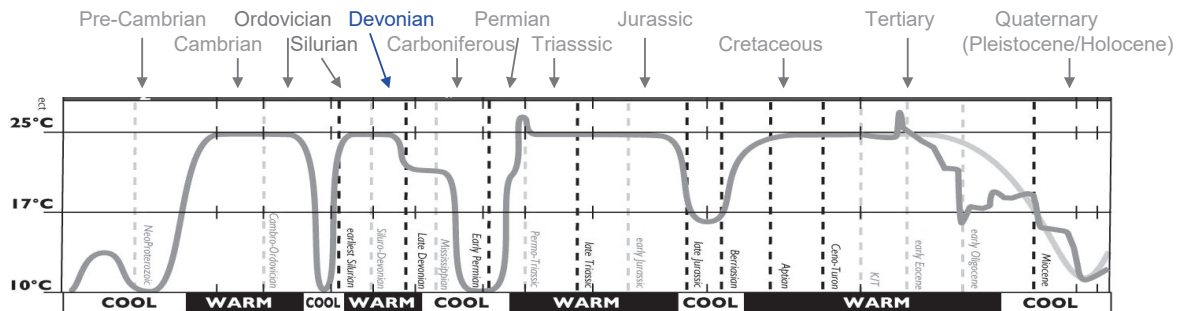
Coral reefs thrived in the clear sunny skies of the southern Arid Belt. Lingering glacial conditions prevailed near the South Pole.

Source: <http://www.scotese.com/silclim.htm>

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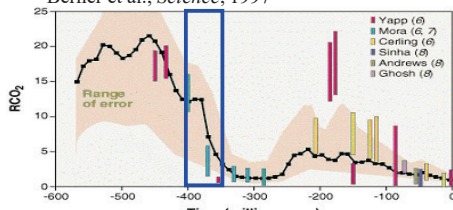


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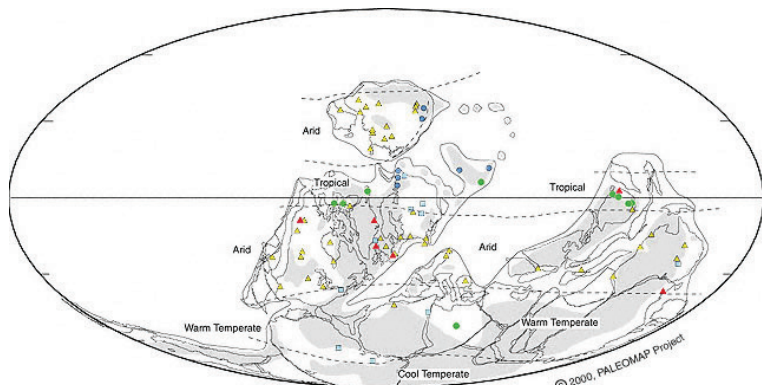
<http://pubs.usgs.gov/gip/fossils/fig15.gif>

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		Middle Proterozoic	2500
Pre-Archean		Early Proterozoic	3800?
		Pre-Archean	3800?

Berner et al., *Science*, 1997



## Middle Devonian Climate (380 million years ago)



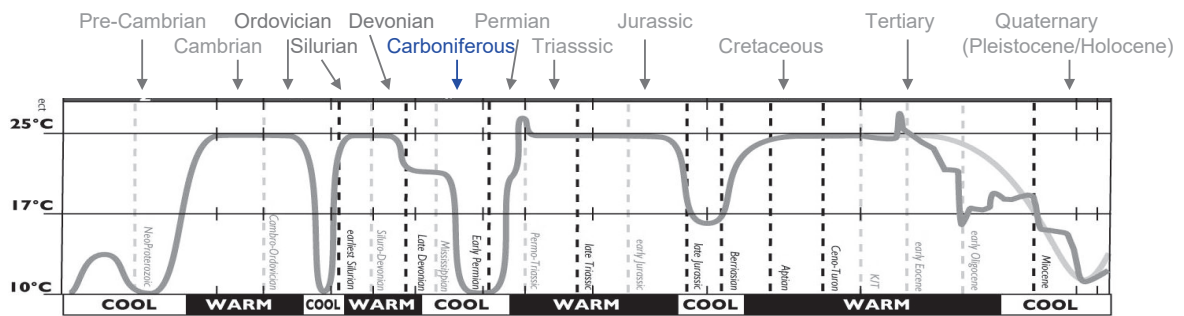
The Equator ran through today's Arctic Canada. Coal began to accumulate as land plants flourished in the equatorial rainy belt. Warm shallow seas covered much of today's North America & Siberia.

Source: <http://www.scotese.com/mdevclim.htm>

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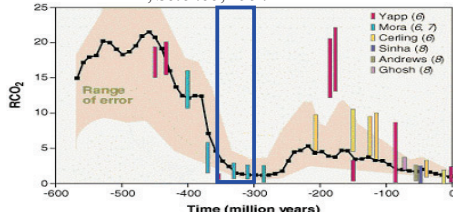


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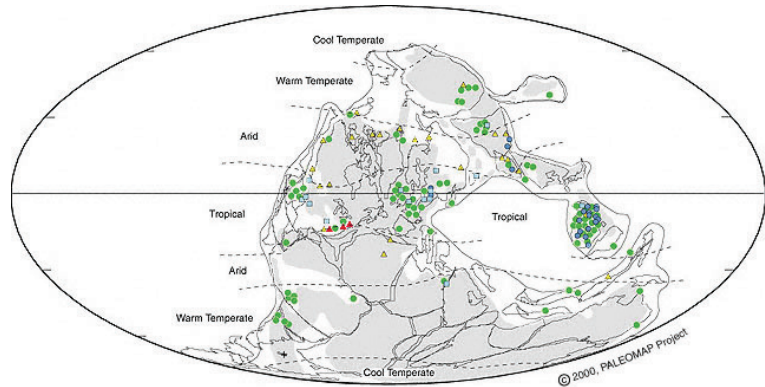
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	Proterozoic	Cambrian	570
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		Middle Proterozoic	2500
	Archean	Late Archean	3800?
		Early Archean	3800?
	Pre-Archean		

Berner et al., Science, 1997



## Carboniferous Climate (350 million years ago)



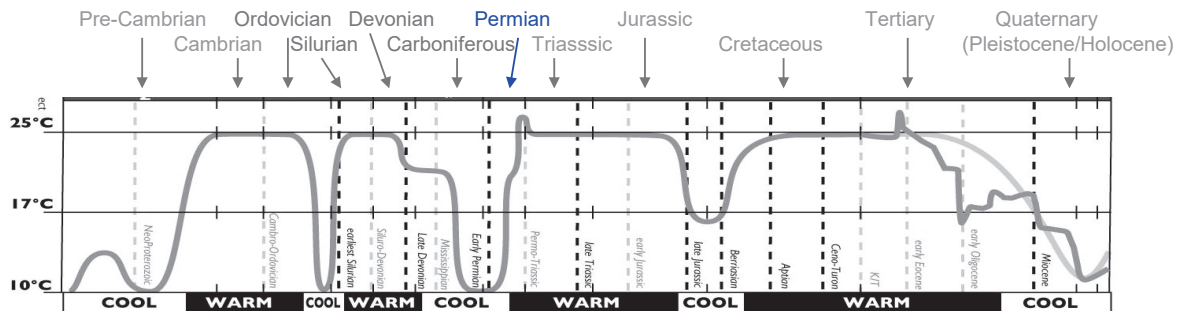
Rainforests covered the tropical regions of Pangea, which was bounded to the north and south by deserts.  
An **ice cap** began to form on the South Pole.

Source: <http://www.scotese.com/serpukcl.htm>

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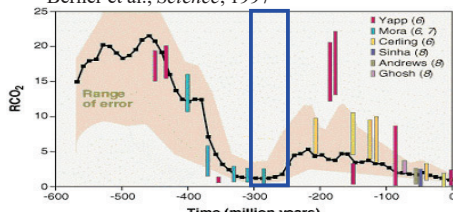


<http://www.scotese.com>

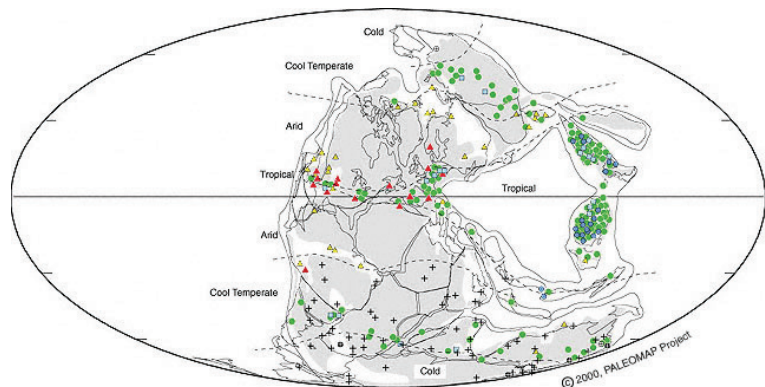
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	Archean	Late Archean	3800?
		Early Archean	3800?
	Pre-Archean		

Berner et al., Science, 1997



## Early Permian Climate (280 million years ago)



Much of the SH was covered by ice as glaciers pushed equator ward.  
Coal was produced in Equatorial & Temperate rainforests during warmer "Interglacial" periods.

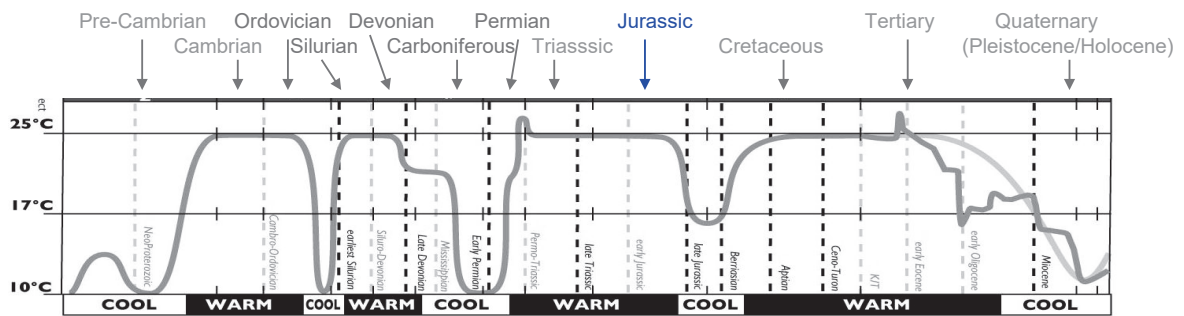
Source: <http://www.scotese.com/epermcli.htm>

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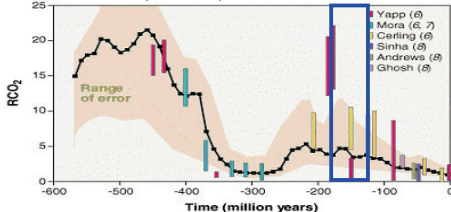


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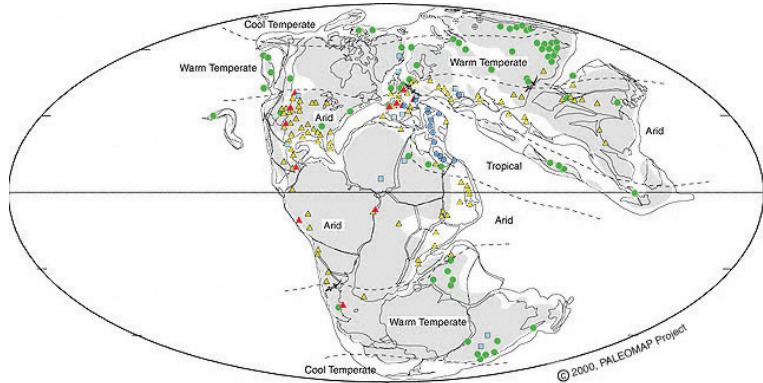
<http://pubs.usgs.gov/gip/fossils/fig15.gif>

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		Early Proterozoic	
Archean	Late Archean	3800?	
	Middle Archean		
	Early Archean		
Pre-Archean			

Berner et al., *Science*, 1997



## Late Jurassic Climate (150 million years ago)



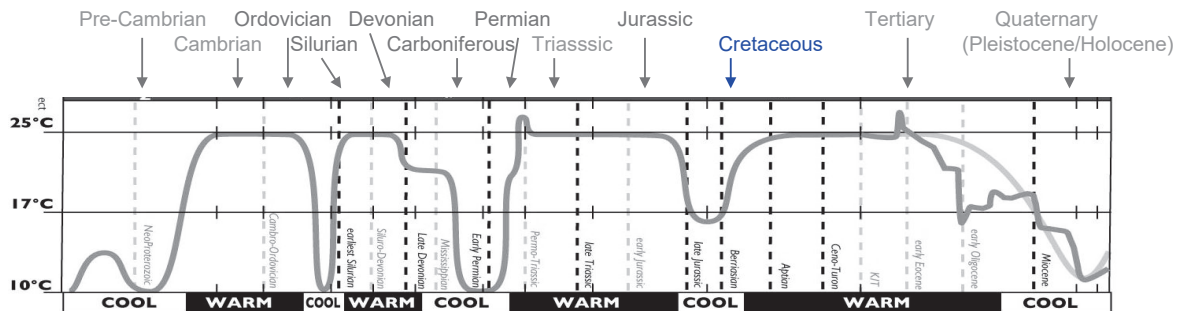
Global climate began to change due to breakup of Pangea.  
The interior of Pangea became moister and seasonal snow & ice frosted the polar regions

Source: <http://www.scotese.com/ljurclim.htm>

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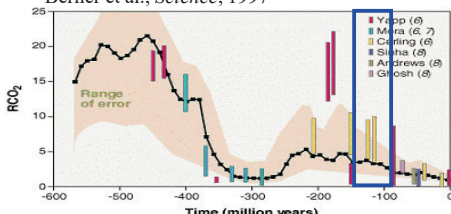


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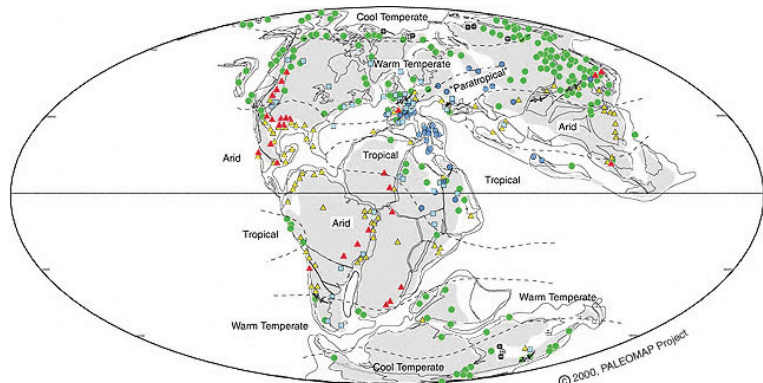
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	Proterozoic	Late Proterozoic	2500
	Middle Proterozoic		
	Early Proterozoic		
Archean	Late Archean	3800?	
	Middle Archean		
	Early Archean		
Pre-Archean			

Berner et al., *Science*, 1997



## Early Cretaceous Climate (120 million years ago)



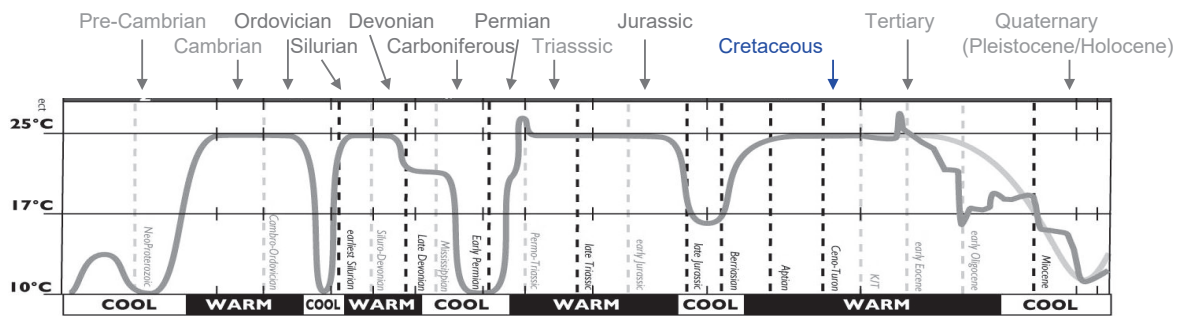
Climate was a mild "Ice House" world. Snow and ice were present during winter and cool temperate forests covered polar regions.

Source: <http://www.scotese.com/ecretcli.htm>

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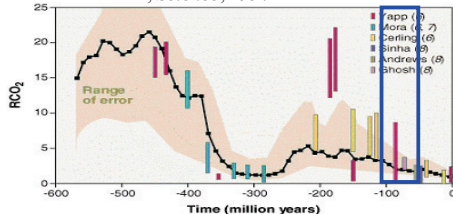
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<http://pubs.usgs.gov/gip/fossils/fig15.gif>

EON	ERA	PERIOD	MILLIONS OF YEARS AGO
Phanerozoic	Cenozoic	Quaternary	1.6
		Tertiary	66
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	Proterozoic	Late Proterozoic	2500
	Archean	Middle Archean	2500
		Early Archean	3800?
		Pre-Archean	3800?

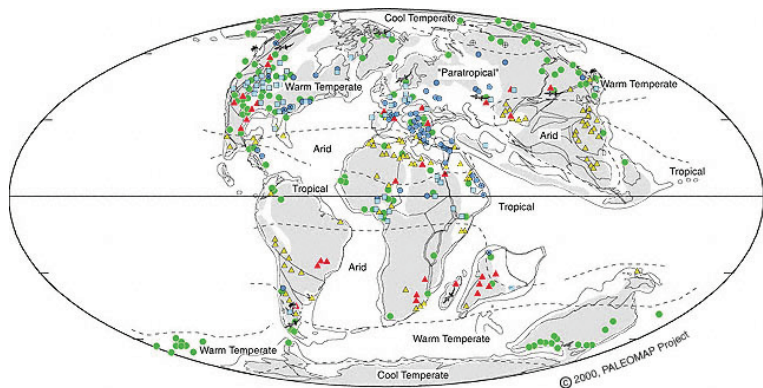
Berner et al., *Science*, 1997



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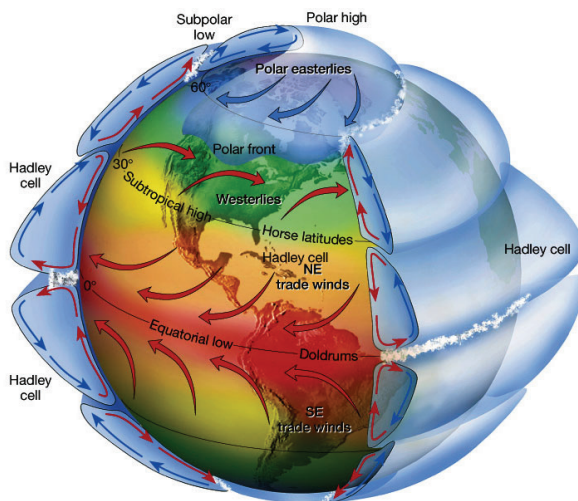
## Late Cretaceous Climate (70 million years ago)



Global climate was much warmer than today. No ice existed at the Poles. Dinosaurs migrated between Temperate Zones as the seasons changed.

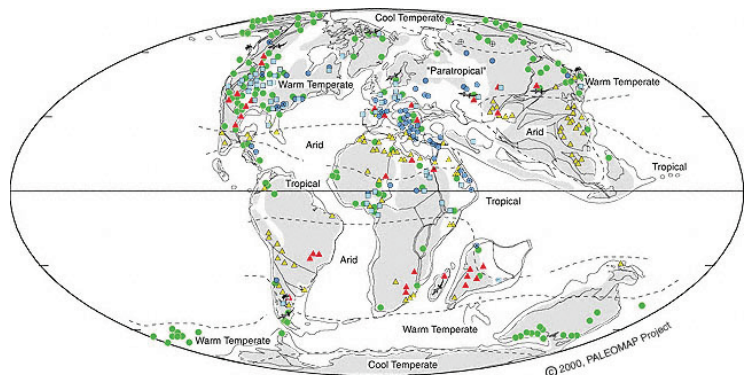
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Lecture #3, Last Slide

## Late Cretaceous Climate (70 million years ago)



Global climate was much warmer than today. No ice existed at the Poles. Dinosaurs migrated between Temperate Zones as the seasons changed.

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Source: <http://www.scotese.com/ecrctcli.htm>

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# Earth's Climate History

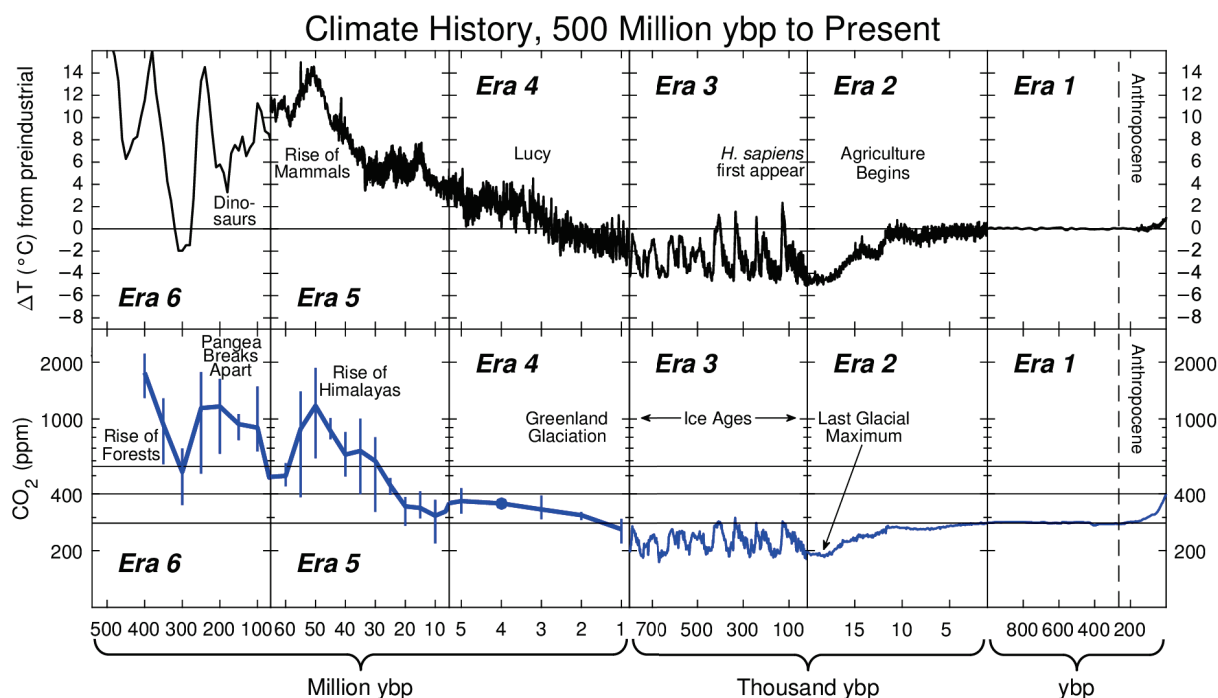


Fig 1.1, *Paris Beacon of Hope*

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## Oxygen Isotopes and the Quaternary Climate Record

Oxygen has three stable isotopes  $^{16}\text{O}$ ,  $^{17}\text{O}$ , and  $^{18}\text{O}$

	Electrons	Protons	Neutrons	Abundance
$^{16}\text{O}$	8	8	8	99.76 %
$^{17}\text{O}$	8	8	9	00.04 %
$^{18}\text{O}$	8	8	10	00.20 %

$^{17}\text{O}$  has such a low abundance that we shall focus on  $^{16}\text{O}$  and  $^{18}\text{O}$

Chemical and biological reactions involving  $^{18}\text{O}$  require more energy than reactions involving  $^{16}\text{O}$  due to increased atomic mass

This “isotope effect” can be used as a proxy to infer past temperature!

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# Oxygen Isotopes and the Quaternary Climate Record

Scientists measured the ratio of  $^{18}\text{O}$  to  $^{16}\text{O}$  in a sample (sea water, shells, etc.) and compare to a “standard value”

$$\delta^{18}\text{O} \text{ (per mil)} = \left( \frac{\left( \frac{^{18}\text{O}}{^{16}\text{O}} \right)_{\text{Sample}} - \left( \frac{^{18}\text{O}}{^{16}\text{O}} \right)_{\text{Standard}}}{\left( \frac{^{18}\text{O}}{^{16}\text{O}} \right)_{\text{Standard}}} \right) \times 10^3$$

Standard often referred to as SMOW: Standard Mean Ocean Water

If  $\delta^{18}\text{O}$  is negative, the sample is “depleted” with respect to current conditions.

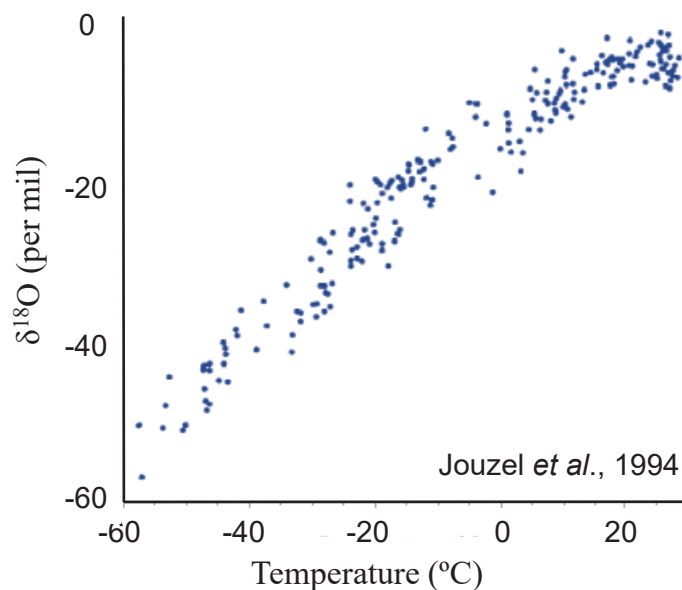
If positive, the sample is “enriched”.

**How might  $\delta^{18}\text{O}$  become enriched or depleted?**

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As temperatures drops, the  $\delta^{18}\text{O}$  of precipitation decreases.

**Why does this occur?**

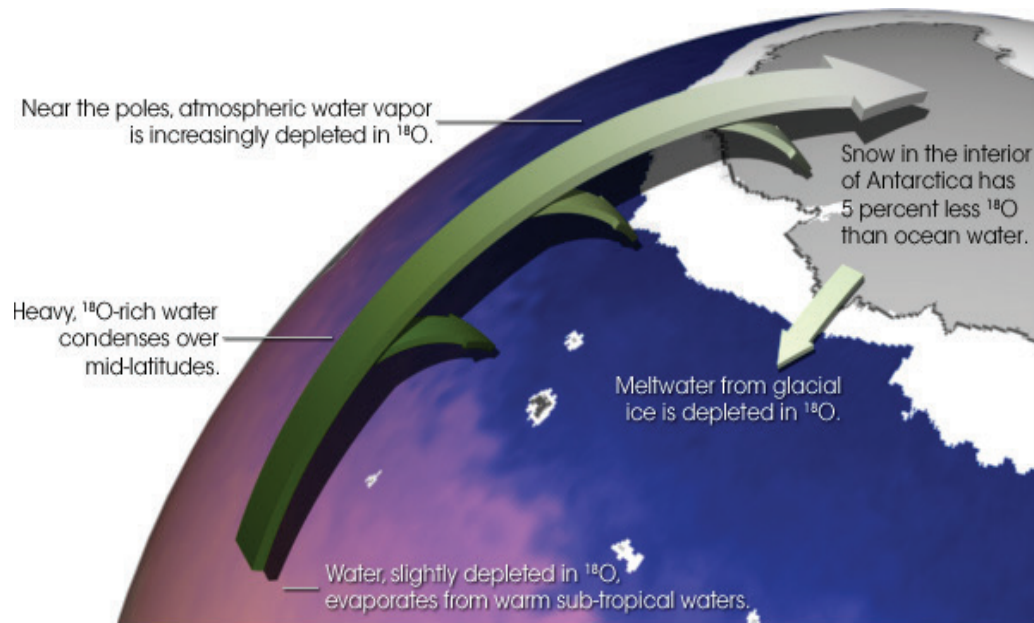
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As an air mass travels poleward,  $\text{H}_2^{18}\text{O}$  rains out more readily than  $\text{H}_2^{16}\text{O}$

When the air mass reaches the pole, its water can have up to ~5% less  $^{18}\text{O}$  than SMOW.



[http://earthobservatory.nasa.gov/Study/Paleoclimatology\\_OxygenBalance/oxygen\\_balance.html](http://earthobservatory.nasa.gov/Study/Paleoclimatology_OxygenBalance/oxygen_balance.html)

**Deuterium (heavy hydrogen) behaves in a way quite similar to  $^{18}\text{O}$  (heavy oxygen) !**

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## Isotopes in Ice Cores: Late Quaternary

- As the air reaches the pole, ambient water precipitate (*i.e.*, it snows!)
- Over many years, layers of snow accumulate, forming an ice sheet. The water in this ice sheet contains a record of climate **at the time the snow was deposited**
- By drilling, extracting, and measuring the  $\delta^{18}\text{O}$  &  $\delta\text{D}$  (deuterium/hydrogen ratio) of ice, scientists are able to estimate past **global temperature & ice volume**
- In reconstructing climate during the quaternary (last 1.6 million years), scientists also look at:
  - $\text{CO}_2$ ,  $\text{CH}_4$ , and  $\text{N}_2\text{O}$  of trapped air
  - $\delta^{18}\text{O}$  of trapped  $\text{O}_2$  in trapped air
  - $\delta^{13}\text{C}$  of  $\text{CO}_2$  in trapped air
  - Particulate matter and a wide range of ions



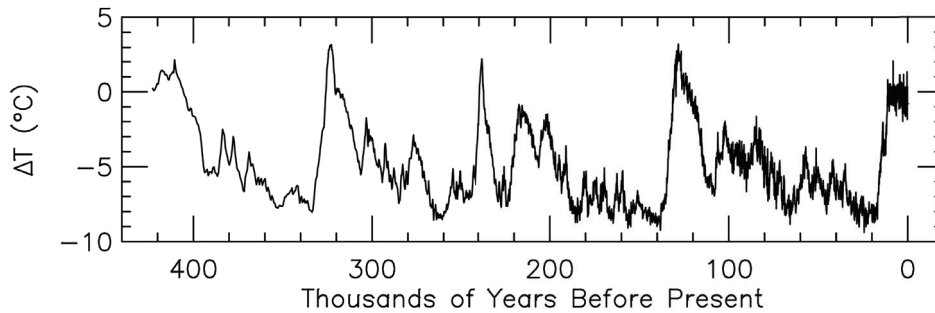
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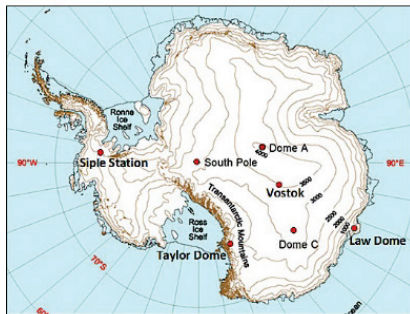
20

# Vostok Ice Core

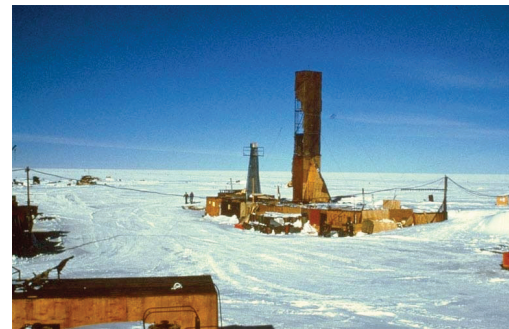
- January 1998: ice core with depth of 3.6 km extracted at Russian Vostok Station, Antarctica
- Vostok ice-core record extends back 400,000 years in time (Petit *et al.*, *Nature*, 1999)
- **Reconstructed temperature based on measurement of the deuterium content of ice**
- $\delta^{18}\text{O}$  shows tremendous variations in global ice volume (not shown)
- Ice core data show last four ice ages, punctuated by relatively brief interglacials



Vostok Drill



[https://cdiac.ess-dive.lbl.gov/trends/co2/ice\\_core\\_co2.html](https://cdiac.ess-dive.lbl.gov/trends/co2/ice_core_co2.html)



<http://www.astrosurf.com/luxorion/Sciences/vostok-drill.jpg>

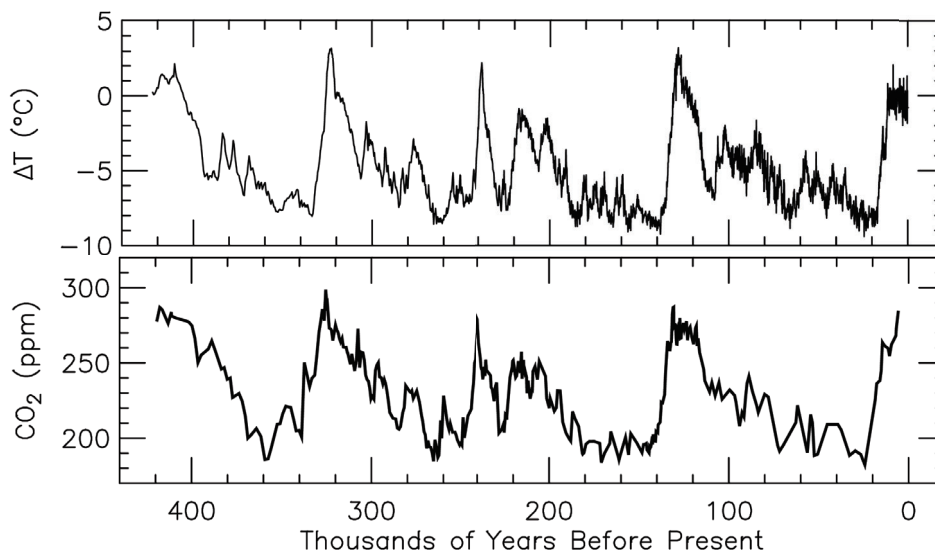
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# Vostok Ice Core

- January 1998: ice core with depth of 3.6 km extracted at Russian Vostok Station, Antarctica
- Vostok ice-core record extends back 400,000 years in time (Petit *et al.*, *Nature*, 1999)
- Reconstructed temperature based on measurement of the deuterium content of ice
- $\delta^{18}\text{O}$  shows tremendous variations in global ice volume (not shown)
- Ice core data show last four ice ages, punctuated by relatively brief interglacials



- **CO<sub>2</sub> (air trapped in ice bubbles) and inferred temperature very highly correlated**
- **Variations in ΔT & CO<sub>2</sub> synchronous upon correction of movement of air bubbles (CO<sub>2</sub>) relative to ice (ΔT) (Parrenin *et al.*, *Science*, 2013: <http://science.sciencemag.org/content/339/6123/1060>)**

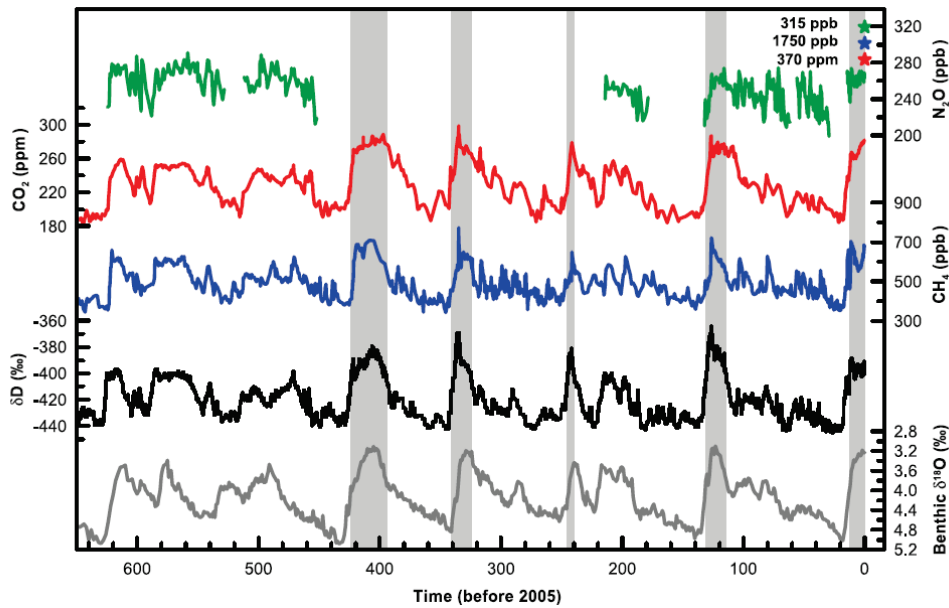
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## Going Back 600,000 years



**Figure 6.3.** Variations of deuterium ( $\delta D$ ; black), a proxy for local temperature, and the atmospheric concentrations of the greenhouse gases  $CO_2$  (red),  $CH_4$  (blue), and nitrous oxide ( $N_2O$ ; green) derived from air trapped within ice cores from Antarctica and from recent atmospheric measurements (Petit et al., 1999; Indermühle et al., 2000; EPICA community members, 2004; Spahni et al., 2005; Siegenthaler et al., 2005a,b). The shading indicates the last interglacial warm periods. Interglacial periods also existed prior to 450 ka, but these were apparently colder than the typical interglacials of the latest Quaternary. The length of the current interglacial is not unusual in the context of the last 650 kyr. The stack of 57 globally distributed benthic  $\delta^{18}O$  marine records (dark grey), a proxy for global ice volume fluctuations (Lisiecki and Raymo, 2005), is displayed for comparison with the ice core data. Downward trends in the benthic  $\delta^{18}O$  curve reflect increasing ice volumes on land. Note that the shaded vertical bars are based on the ice core age model (EPICA community members, 2004), and that the marine record is plotted on its original time scale based on tuning to the orbital parameters (Lisiecki and Raymo, 2005). The stars and labels indicate atmospheric concentrations at year 2000.

Figure 6.3, IPCC 2007

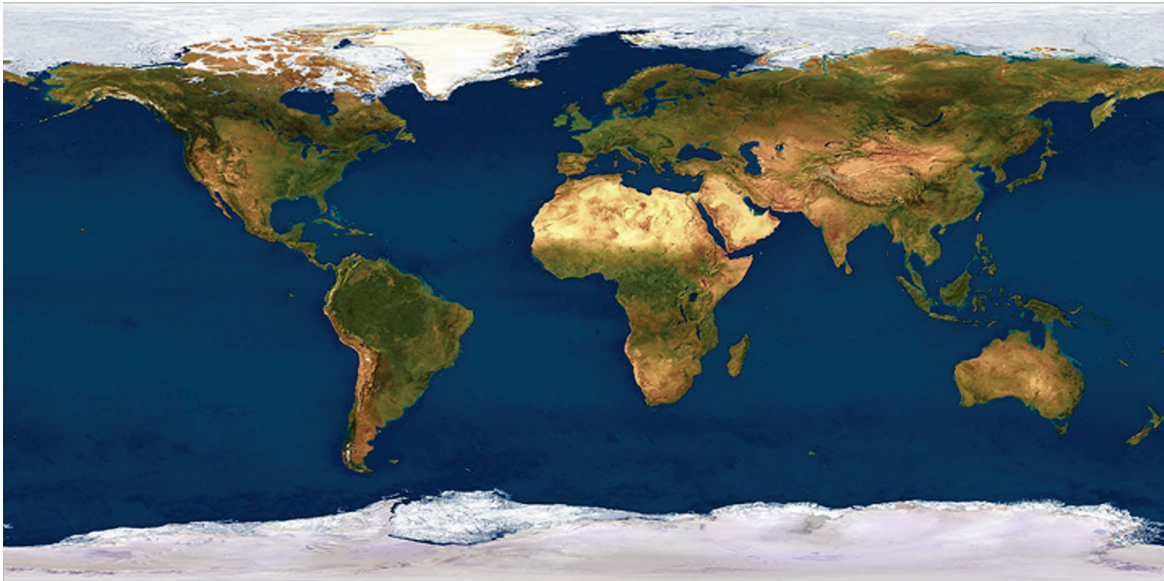
See <https://epic.awi.de/id/eprint/18400/1/Oer2008a.pdf> for description of EPICA, European Project for Ice Coring in Antarctica

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## Present Day



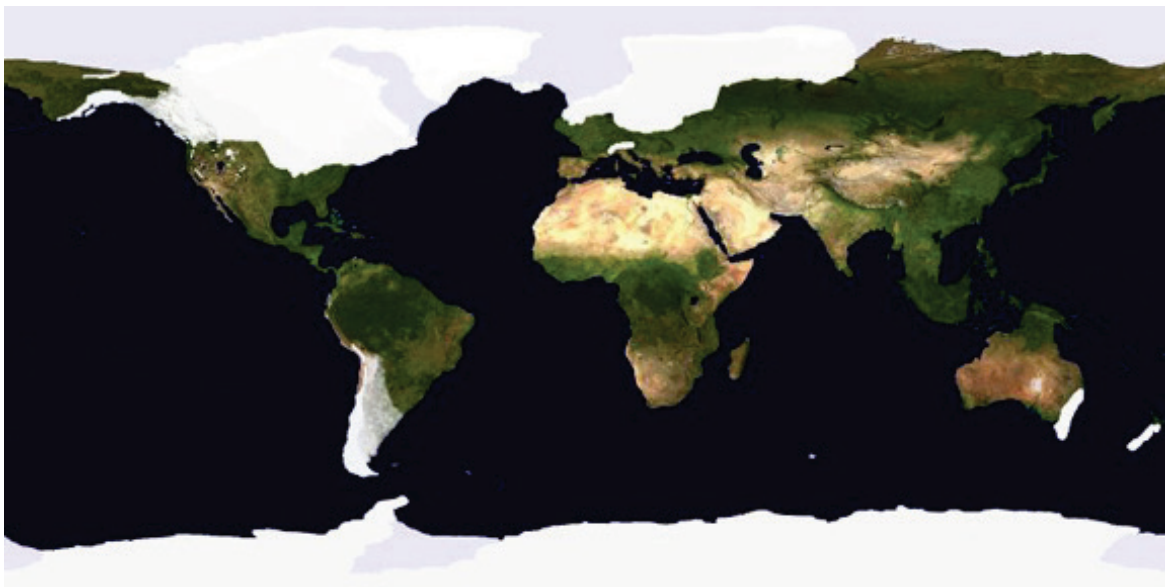
<http://www.planetaryvisions.com/release2007-1/satmap.jpg>

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# Glacial Maximum



<http://www.johnstonsarchive.net/spaceart/cylmaps.html>

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# No Polar Ice



<http://www.johnstonsarchive.net/spaceart/cylmaps.html>

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# Fairly Late Appreciation that Earth Undergoes Ice Ages

On 24 July 1837, at the annual meeting of the Swiss Society of Natural Sciences, Louis Agassiz (1807–1873) startled his learned associates by presenting a paper dealing not, as expected, with the fossil fishes found in far-off Brazil, but with the scratched and faceted boulders that dotted the Jura mountains around Neuchâtel itself. Agassiz argues that these erratic boulders ... chunks of rock appearing in locations far removed from their areas of origin ... could only be interpreted as evidence of past glaciation.

This began a dispute – one of the most violent in the history of geology – that was to rage for more than a quarter century and would end with the universal acceptance of the ice-age theory.

Although this concept did not begin with Agassiz, he served to bring the glacial theory out of scientific obscurity and into the public eye.



Portrait of Louis Agassiz at the Unteraar Glacier

<http://www.museum-neuchatel.ch/new/images/dynamic/pages/12/agassiz.jpg>

*Ice Ages*, Imbrie and Imbrie, Harvard Univ Press, 1979.

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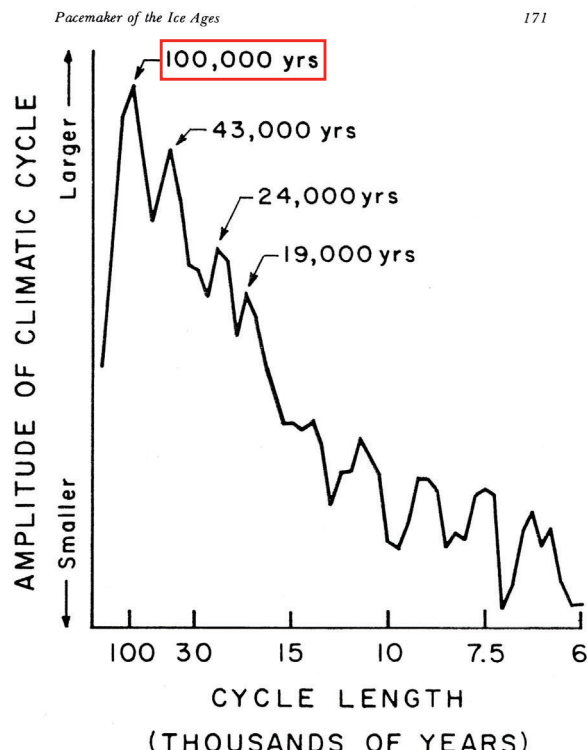
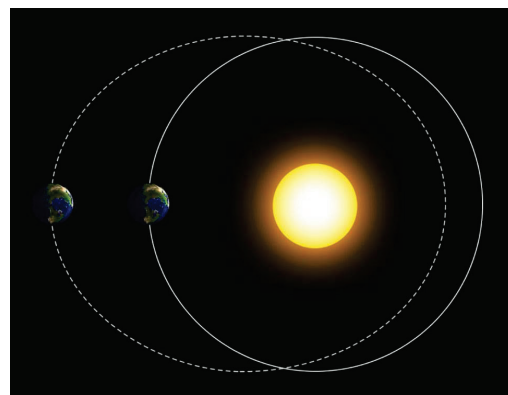


Figure 42. Spectrum of climatic variation over the past half-million years. This graph—showing the relative importance of different climatic cycles in the isotopic record of two Indian Ocean cores—confirmed many predictions of the Milankovitch theory. (Data from J.D. Hays et al., 1976.)

*Ice Ages*, Imbrie and Imbrie, Harvard Univ Pres, 1979

Fourier analysis reveals Earth's climate is changing in a periodic fashion

**100,000 year cycle** due to changes in the eccentricity of Earth's orbit, mainly due to gravitational pull of Jupiter and Saturn.



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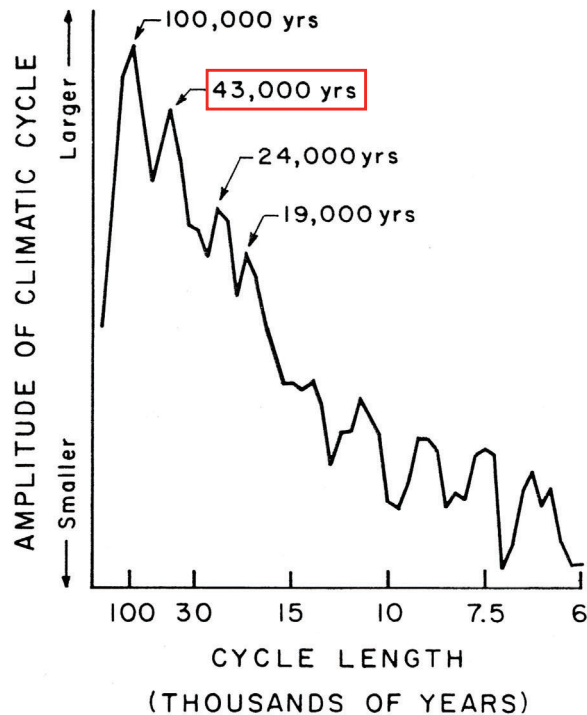


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*Ice Ages*, Imbrie and Imbrie, Harvard Univ Pres, 1979

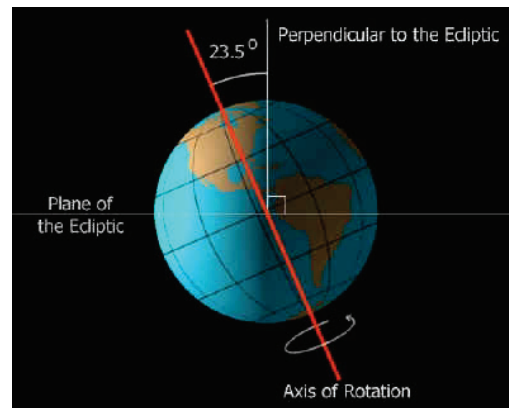
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Fourier analysis reveals Earth's climate is changing in a periodic fashion

**43,000 year cycle** due to changes in tilt of Earth's axis (obliquity).



Fourier analysis reveals Earth's climate is changing in a periodic fashion

**24,000 and 19,000 year cycles** due to Earth "wobbling" on its axis.

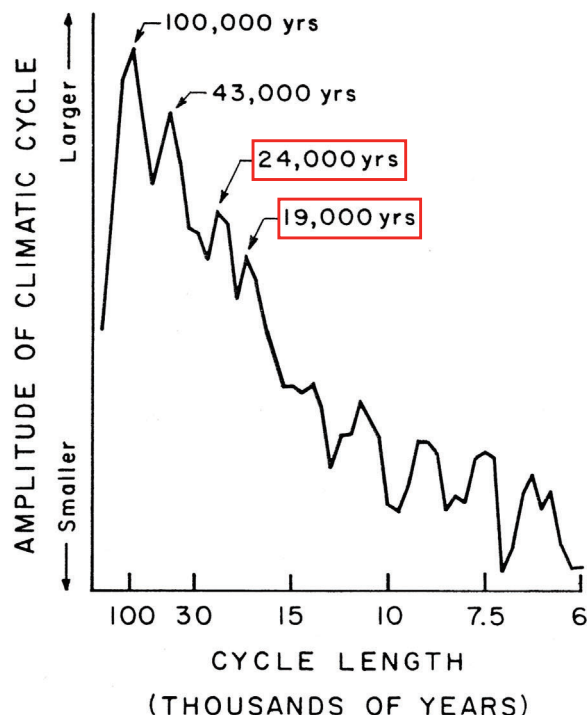
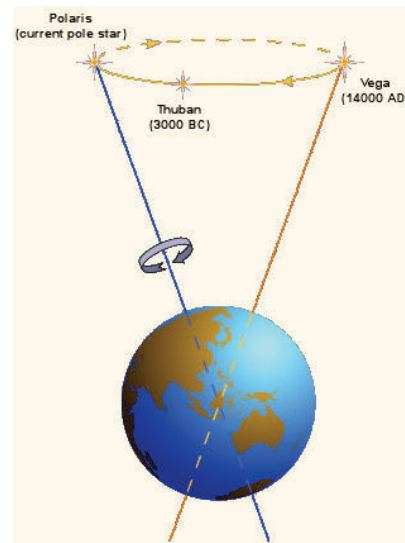


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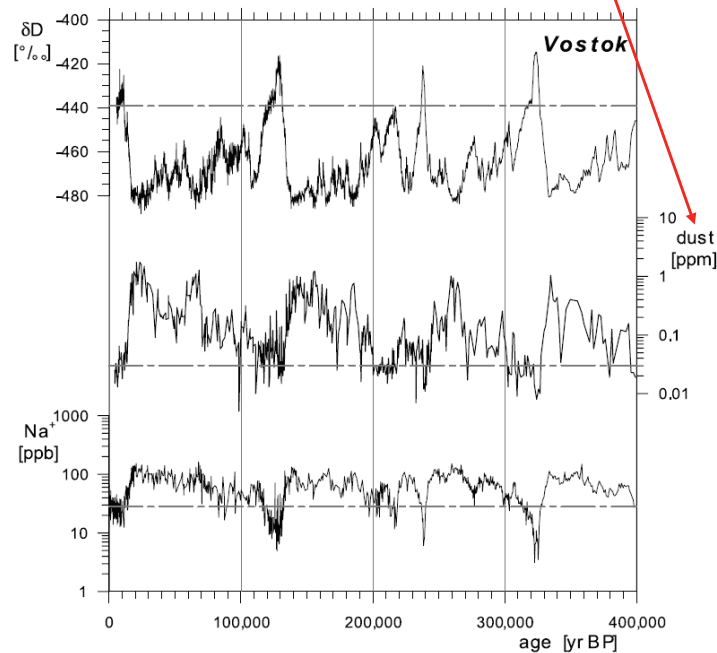
*Ice Ages*, Imbrie and Imbrie, Harvard Univ Pres, 1979

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# Glacial Periods MUCH Dustier than Interglacials



**Figure 3.** Temporal evolution of  $\delta D$  representing changes in the average local condensation temperature during snow formation, the particulate dust, and the sea-salt component  $Na^+$  over the last four glacial cycles as recorded in the East Antarctic Vostok ice core [Petit *et al.*, 1999]. Dashed-dotted lines indicate the mean Holocene level from 0 to 10,000 years B.P.

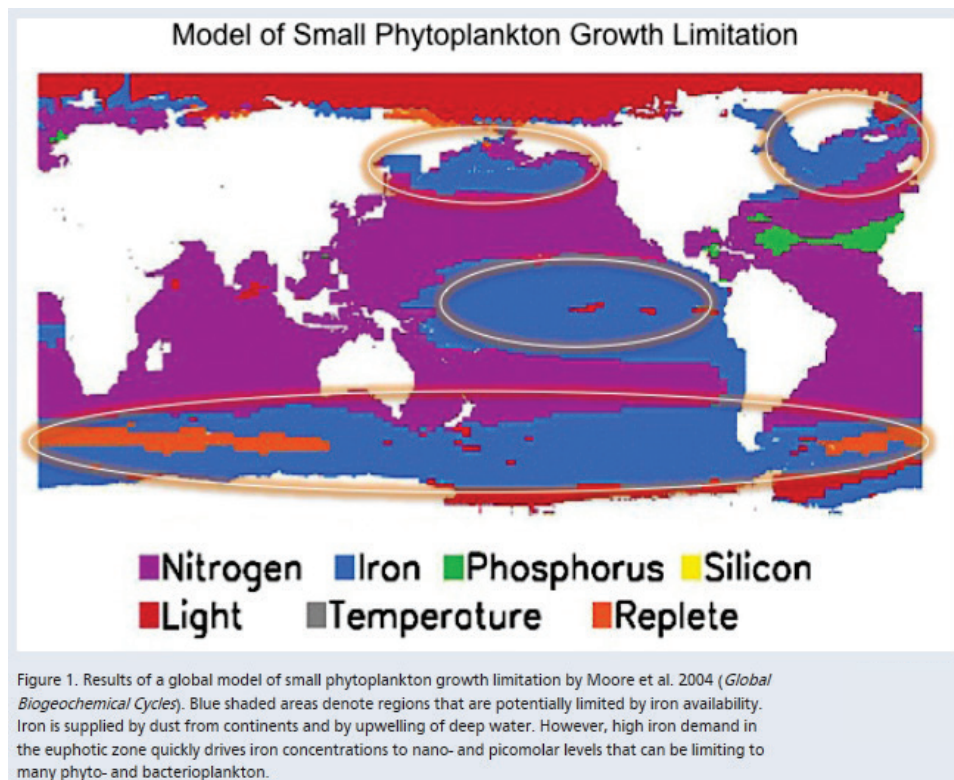
Fischer *et al.*, Reviews of Geophysics, 2007

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## Biology in Today's Ocean



<http://www.whoi.edu/page.do?pid=130796>

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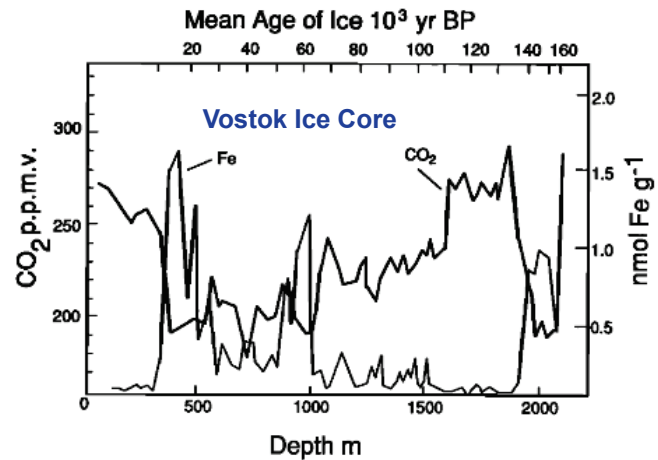
# Connection to Glacial CO<sub>2</sub>

GLACIAL-INTERGLACIAL CO<sub>2</sub> CHANGE:  
THE IRON HYPOTHESIS

PALEOCEANOGRAPHY, VOL.5,  
NO.1, PAGES 1-13 1990

John H. Martin

In contrast, atmospheric dust Fe supplies were 50 times higher during the last glacial maximum (LGM). Because of this Fe enrichment, phytoplankton growth may have been greatly enhanced, larger amounts of upwelled nutrients may have been used, and the resulting stimulation of new productivity may have contributed to the LGM drawdown of atmospheric CO<sub>2</sub> to levels of less than 200 ppm. Background information and arguments in support of this hypothesis are presented.



<http://onlinelibrary.wiley.com/doi/10.1029/PA005i001p00001/abstract>

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## Time to get quantitative:

### how do changes in radiative forcing affect temperature?

Let's relate a change in temperature to a change in radiative forcing:

$$\Delta T = \lambda \Delta F$$

$\lambda$  is the climate sensitivity factor in units of  $\frac{K}{W/m^2}$

For an ideal blackbody:  $F = \sigma T^4$

$$\frac{dF}{dT} = 4 \sigma T^3$$

Above equation can be re-arranged to yield:

$$\Delta T \approx \frac{1}{4 \sigma T^3} \Delta F$$

So: 
$$\lambda = \frac{1}{4 \sigma T^3}$$

If we plug in value of Boltzmann's constant and global mean T at which Earth radiates to space, we find  $\lambda_{BB} \approx 0.3 K / (W m^{-2})$

Here: BB refers to Black Body

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## Time to get quantitative:

### how do changes in radiative forcing affect temperature?

Let's relate a change in temperature to a change in radiative forcing:

$$\Delta T = \lambda \Delta F$$

$\lambda$  is the climate sensitivity factor in units of  $\frac{\text{K}}{\text{W/m}^2}$

For an ideal blackbody:  $F = \sigma T^4$

$$\frac{dF}{dT} = 4 \sigma T^3$$

**We write:**

$$\lambda_{\text{ACTUAL}} = \lambda_{\text{BB}} (1 + f_{\text{H}_2\text{O}})$$

where  $f_{\text{H}_2\text{O}}$  is the H<sub>2</sub>O feedback

Here,  $f_{\text{H}_2\text{O}} \approx 1.08$

Above equation can be re-arranged to yield:

$$\Delta T \approx \frac{1}{4 \sigma T^3} \Delta F$$

So: 
$$\lambda = \frac{1}{4 \sigma T^3}$$

**Another estimate of the response of T to  $\Delta F$  can be found using a climate model representing that as the atmosphere warms, it can hold more H<sub>2</sub>O:**

$$\lambda_{\text{ACTUAL}} \approx 0.63 \pm 0.13 \text{ K / (W m}^{-2}\text{)}$$

Table 9.5, IPCC (2013)

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## Time to get quantitative:

### how do changes in radiative forcing affect temperature?

Hence: 
$$\Delta T \approx 0.63 \frac{\text{K}}{\text{W/m}^2} \Delta F$$

**How much does  $\Delta F$  change when CO<sub>2</sub> changes?**

As we will explore in more detail later in class (16 Feb 2017):

$$\Delta F \approx 5.35 \text{ W/m}^2 \ln \left( \frac{\text{CO}_2^{\text{Final}}}{\text{CO}_2^{\text{Initial}}} \right)$$

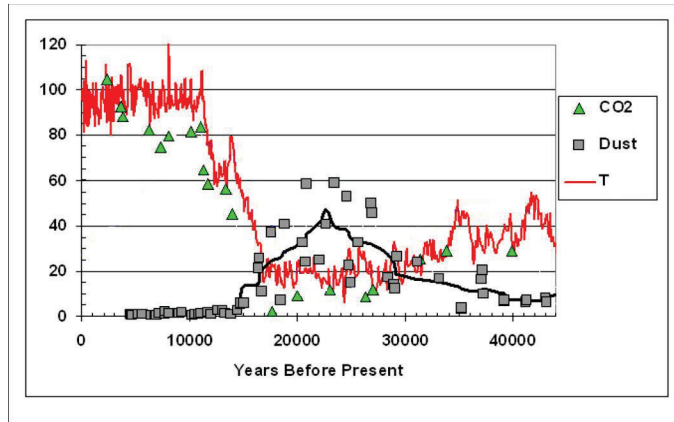
Changes in  $\Delta F$  can be caused by changes in chemical composition (GHGs), albedo, aerosol loading, as well as solar output

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## Glacial to interglacial changes in T, CO<sub>2</sub> and dust



Vostok ice core data for changes in temperature (units of 0.1 K), CO<sub>2</sub> (ppmv), and dust aerosols (linear scale normalized to unity for Holocene). Black line shows 5 point running mean of dust.

Chylek and Lohmann, GRL, 2008

Chylek and Lohmann (2008) assume:

a) global avg  $\Delta T$ , glacial to interglacial, was 4.65 K \*

b)  $\Delta F_{\text{CO}_2} = 2.4 \text{ W m}^{-2}$ ,  $\Delta F_{\text{CH}_4+\text{N}_2\text{O}} = 0.27 \text{ W m}^{-2}$ ,  $\Delta F_{\text{ALBEDO}} = 3.5 \text{ W m}^{-2}$ , &  $\Delta F_{\text{AEROSOLS}} = 3.3 \text{ W m}^{-2}$

From this they deduce  $\lambda_{\text{ACTUAL}} = 0.49 \text{ K / W m}^{-2}$

Since  $0.49 \text{ K / W m}^{-2} < 0.63 \text{ K / W m}^{-2}$ , one would conclude that either the H<sub>2</sub>O feedback is smaller than found in IPCC climate models and/or changes in clouds serve as a negative feedback

\* Global  $\Delta T$  is about half that recorded at Vostok, as stated in the caption of Fig 4.9a of Houghton

## Glacial to interglacial changes in T, CO<sub>2</sub> and dust

Chylek and Lohmann (2008) are trying to calculate the sensitivity of climate to various forcings, with and without the consideration of aerosols

$\Delta F$ with aerosols(W/m <sup>2</sup> )	
CO <sub>2</sub>	2.40
CH <sub>4</sub> +N <sub>2</sub> O	0.27
Albedo	3.50
Aerosols	3.30



$$\Delta T = \lambda_{\text{Considering Aerosols}} (\Delta F_{\text{CO}_2} + \Delta F_{\text{CH}_4+\text{N}_2\text{O}} + \Delta F_{\text{ALBEDO}} + \Delta F_{\text{AEROSOLS}})$$

$$\lambda_{\text{Considering Aerosols}} = \frac{\Delta T}{\Delta F_{\text{CO}_2} + \Delta F_{\text{CH}_4+\text{N}_2\text{O}} + \Delta F_{\text{ALBEDO}} + \Delta F_{\text{AEROSOLS}}} = \frac{4.65 \text{ K}}{9.47 \text{ W m}^{-2}} = 0.49 \text{ K / W m}^{-2}$$

$$\text{If } \lambda_{\text{Considering Aerosols}} = \lambda_{\text{BB}} (1 + f) \text{ and } \lambda_{\text{BB}} = 0.3 \text{ K / W m}^{-2}, \\ \text{then } f = 0.63$$

# Glacial to interglacial changes in T, CO<sub>2</sub> and dust

Chylek and Lohmann (2008) are trying to calculate the sensitivity of climate to various forcings, with and without the consideration of aerosols

	$\Delta F$ with aerosols(W/m <sup>2</sup> )	$\Delta F$ without aerosols (W/m <sup>2</sup> )
CO <sub>2</sub>	2.40	2.40
CH <sub>4</sub> +N <sub>2</sub> O	0.27	0.27
Albedo	3.50	3.50
Aerosols	3.30	0.

$$\Delta T = \lambda_{\text{No Aerosols}} (\Delta F_{\text{CO}_2} + \Delta F_{\text{CH}_4+\text{N}_2\text{O}} + \Delta F_{\text{ALBEDO}})$$

$$\lambda_{\text{No Aerosols}} = \frac{\Delta T}{\Delta F_{\text{CO}_2} + \Delta F_{\text{CH}_4+\text{N}_2\text{O}} + \Delta F_{\text{ALBEDO}}} = \frac{4.65 \text{ K}}{6.17 \text{ W m}^{-2}} = 0.75 \text{ K / W m}^{-2}$$

If  $\lambda_{\text{No Aerosols}} = \lambda_{\text{BB}} (1 + f)$  and  $\lambda_{\text{BB}} = 0.3 \text{ K / W m}^{-2}$ ,  
then  $f = 1.5$

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## Let's apply these two climate sensitivities to *future* temperature

Both future scenarios assume:

- CO<sub>2</sub> doubles: i.e.,  $\Delta F_{\text{CO}_2} = 5.35 \ln(2) \text{ W/m}^2$  or  $= 3.7 \text{ W/m}^2$
- surface radiative forcing of CH<sub>4</sub> + N<sub>2</sub>O will be 40% of CO<sub>2</sub> (future mimics past)

Scenario #1: Weak **Feedback** found considering aerosol radiative forcing in paleo data & no future change in Earth's albedo

Scenario #2: Strong **Feedback** found assuming no aerosol radiative forcing in paleo data & additional surface radiative forcing of 3.4 W/m<sup>2</sup> due to decline in Earth's albedo (i.e., the positive ice-albedo feedback will occur)

	Scenario #1 $\Delta F$ (W m <sup>-2</sup> )	Scenario #2 $\Delta F$ (W m <sup>-2</sup> )
CO <sub>2</sub>	3.7	3.7
CH <sub>4</sub> + N <sub>2</sub> O	1.5	1.5
Albedo	0.0	3.4
Total $\Delta F$	5.2	8.6

$\Delta T \Rightarrow$

or

### Take away messages:

- Climate sensitivity inferred from ice core record depends on how aerosols are handled
- Future climate will be quite sensitive to:
  - the efficacy of atmospheric feedbacks (H<sub>2</sub>O, clouds)
  - the radiative forcing of aerosols (not considered in our simple future scenario)
  - how surface albedo changes

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## Final Thought

There is much more “recent climate history”, such as:

- a) Younger Dryas cooling event at end of last ice age
- b) Medieval climate maximum
- c) the Little Ice Age (1650 to 1850)

that is deserving of our attention. A few slides on these topics are included in the Extra Material that follows (you will not be tested on the material in these 3 slides)

**Problem Set #1** is due at start of class on Tuesday, February 19 (one week from today) and covers material presented in Lectures 1 to 5

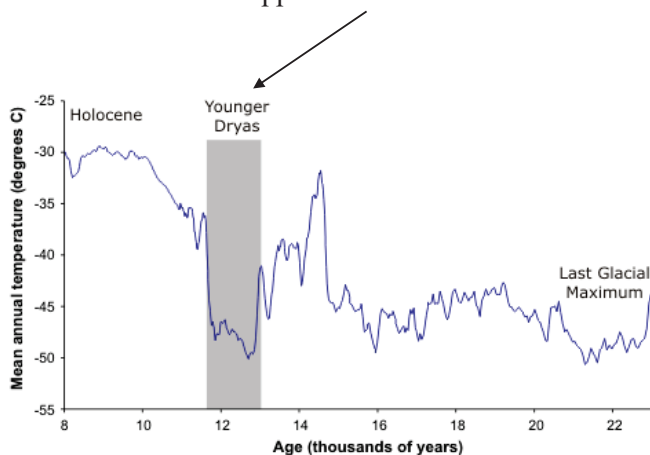
If you have questions, please stop by our offices (Ross: Atlantic 2403; Walt: Atlantic 4100) during either our office hours or normal working hours.

You're also welcome to email us to set up a time to meet

### Extra Slide 1

#### Younger Dryas (about 12,000 years ago)

Around 12,000 years ago, mean annual temperatures abruptly dropped to levels similar to those during the last glacial maximum



Most scientists believe the cool conditions of the Younger Dryas resulted from a flood of fresh water into the North Atlantic that shut down ocean's thermohaline circulation.

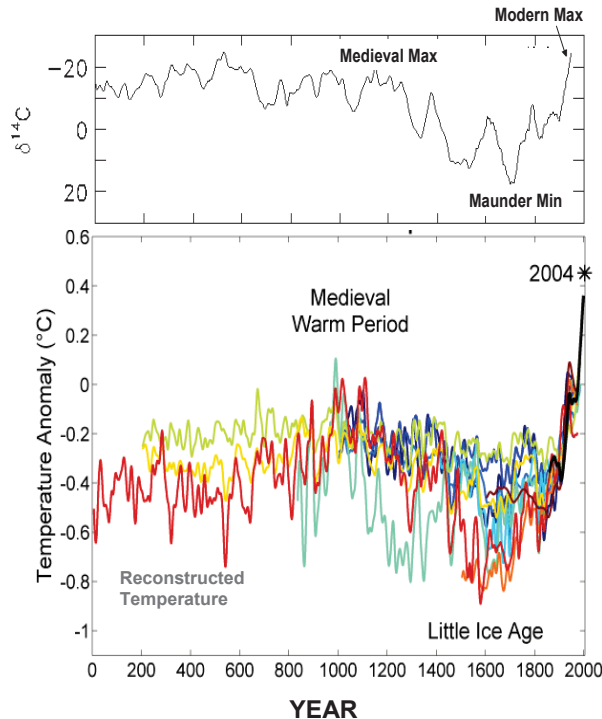
The flood of fresh water was due to discharge from glacial lakes, formed by the melt water of retreating glaciers.

Some geologists (Firestone *et al.*, *PNAS*, 2007) believe that the Younger Dryas was compounded by a terrestrial impact.

<http://www.ncdc.noaa.gov/paleo/abrupt/data4.html>

# Medieval Warm Period (MWP) ~800 to 1300 AD

## Extra Slide 2



$\delta^{14}\text{C}$  (radiocarbon) is a proxy that can be used to estimate past solar activity.

Carbon-14 is produced when cosmic rays hit nitrogen ( $^{14}\text{N}$ ), inducing a decay that transforms this molecule to Carbon-14 (half life of ~5,730 yrs).

Increased solar activity results in a reduction of cosmic rays reaching Earth's atmosphere, reducing production of carbon-14, because cosmic rays are blocked by the outward sweep of magnetic fields of the solar wind.

Measurements of  $^{14}\text{C}$  suggest primary cause of warm conditions during MWP was rise in solar activity

[http://en.wikipedia.org/wiki/Solar\\_variation](http://en.wikipedia.org/wiki/Solar_variation)

[http://en.wikipedia.org/wiki/Medieval\\_Warm\\_Period](http://en.wikipedia.org/wiki/Medieval_Warm_Period)

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## Extra Slide 3

### Little Ice Age (~1350 to 1900)

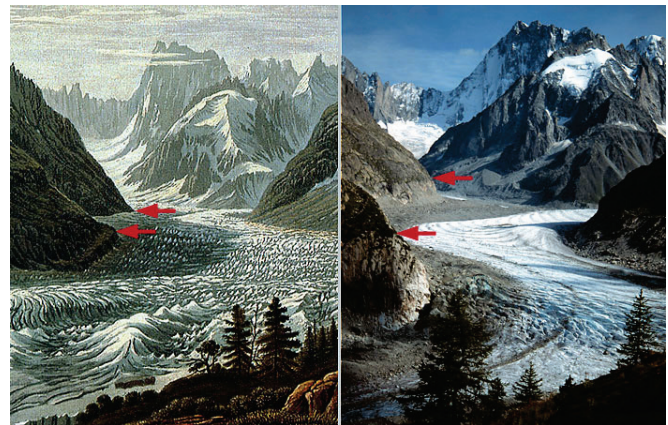
Major rivers (Thames) & waterways (NY harbor) frequently froze.

Crops and livestock failed.

Cities flooded.

Glaciers expanded.

Why did this happen?



<http://www.swisseduc.ch/glaciers/glossary/little-ice-age-two-en.html>

1. Little ice age was an extended period of quiet solar activity:  
coldest time period is associated with the Maunder Minimum  
(time of very low sunspot activity  $\Rightarrow$  reduced solar irradiance).
2. Several large volcanic eruptions during this period; resulting aerosol loading led to a reduction in amount solar radiation reaching the surface.
3. Increase in albedo associated with the colder temperatures (colder T results in more ice) led to even more cooling.

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