

Geologic Evolution of Earth's Atmosphere

AOSC / CHEM 433 & AOSC / CHEM 633

Ross Salawitch

Class Web Sites:

<http://www2.atmos.umd.edu/~rjs/class/spr2022>

<https://myelms.umd.edu/courses/137772>



<https://www.videoblocks.com/video/earth-sunset-spacewalk-view-from-space-station-r7dydlcsgjd23vml0>

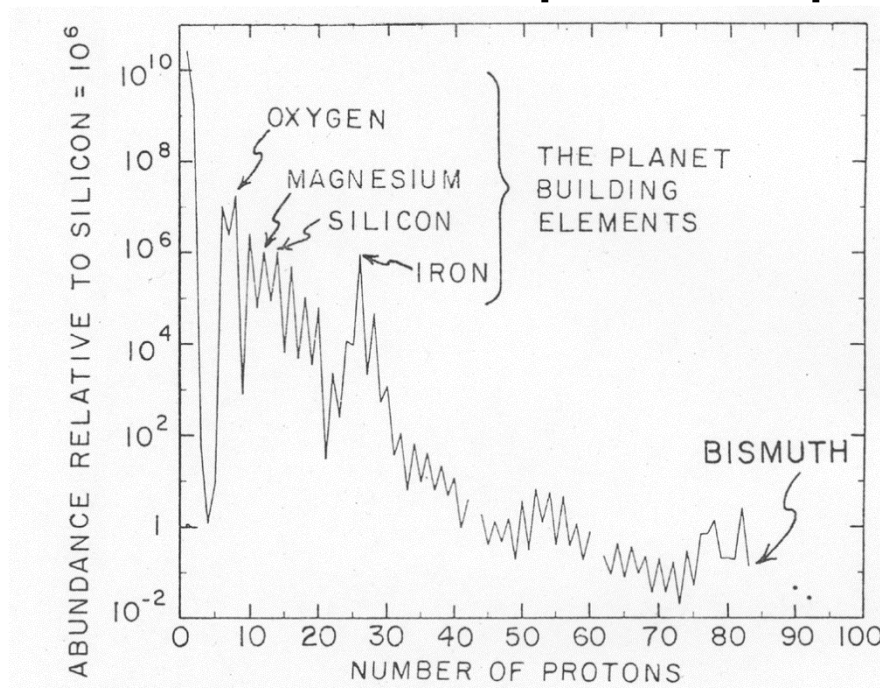
Lecture 1

27 January 2022

Geological Evolution of Earth's Atmosphere:

"In the Beginning"

- **Assemblage of 92 natural elements**
- **Elemental composition of Earth basically unchanged over 4.5 Gyr**
 - Gravitational escape restricted to a few gases (H, He)
 - Extra-terrestrial inputs (comets, meteorites) relatively unimportant
- **Biogeochemical cycling** of elements between reservoirs of Earth "system" determines atmospheric composition



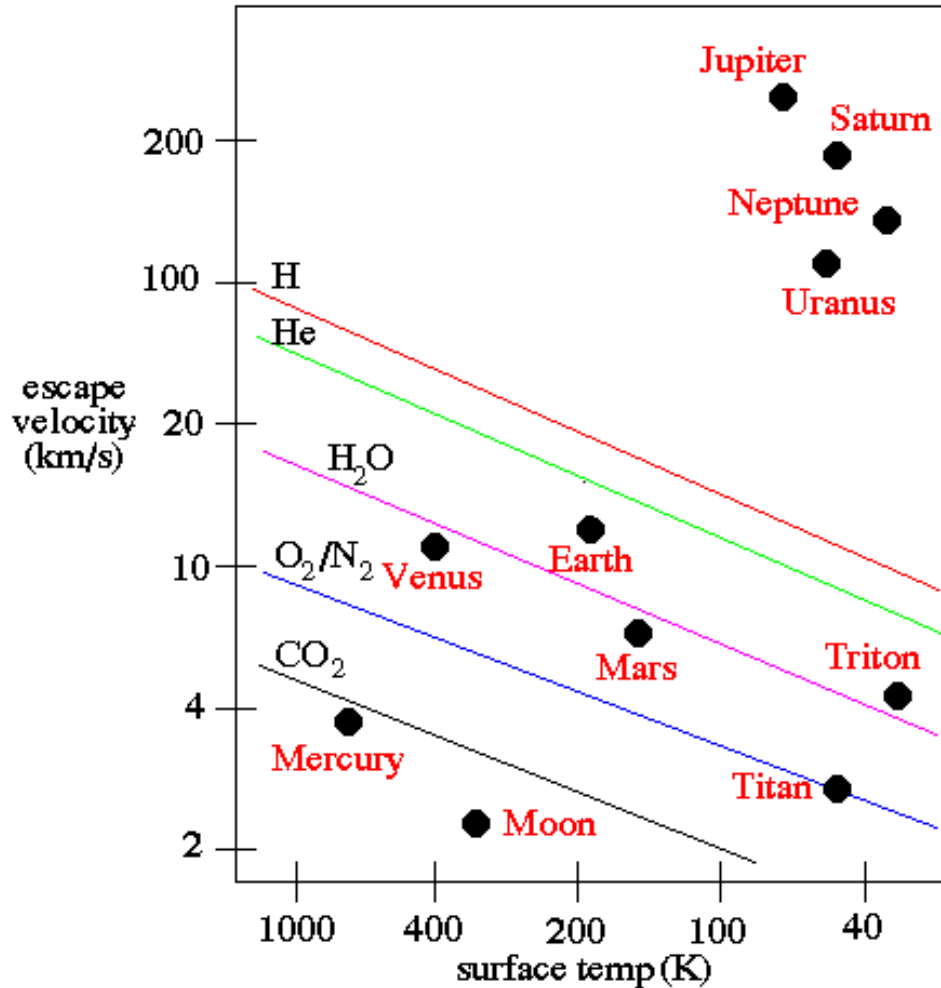
From "How to Build a Habitable Planet"
By W.S. Broecker, ELDIGIO Press, pg 57

Geological Evolution of Earth's Atmosphere:

Earth, Mars, and Venus

	Earth	Venus	Mars
Radius (km)	6400	6100	3400
Mass (10^{24} kg)	6.0	4.9	0.6
Albedo	0.3	0.8	0.22
Distance from Sun (A.U.)	1	0.72	1.52
Surface Pressure (atm)	1	91	0.007
Surface Temperature (K)	$\sim 15^\circ\text{C}$	$\sim 460^\circ\text{C}$	-140°C to 20°C
N ₂ (mol/mol)	0.78	3.4×10^{-2}	2.7×10^{-2}
O ₂ (mol/mol)	0.21	6.9×10^{-5}	1.3×10^{-3}
CO ₂ (mol/mol)	3.7×10^{-4}	0.96	0.95
H ₂ O (mol/mol)	1×10^{-2}	3×10^{-3}	3×10^{-4}
SO ₂ (mol/mol)	1×10^{-9}	1.5×10^{-4}	Nil
Cloud Composition	H ₂ O	H ₂ SO ₄	Mineral Dust

Geological Evolution of Earth's Atmosphere: *Earth, Mars, and Venus*



<http://abyss.uoregon.edu/~js/ast121/lectures/lec14.html>

Geological Evolution of Earth's Atmosphere:

Earth is of course the water planet
The source of Earth's water has been unclear

A new study finds that Earth's water may have come from materials that were present in the inner solar system at the time the planet formed – instead of far-reaching comets or asteroids delivering such water. The findings published 28 Aug 2020 in Science suggest that Earth may have always been wet.

“Our discovery shows that the Earth's building blocks might have significantly contributed to the Earth's water,” said lead author Laurette Piani. “Hydrogen-bearing material was present in the inner solar system at the time of the rocky planet formation, even though the temperatures were too high for water to condense.”

The findings from this study are surprising because the Earth's building blocks are often presumed to be dry. They come from inner zones of the solar system where temperatures would have been too high for water to condense and come together with other solids during planet formation.

Enstatite chondrites have similar oxygen, titanium and calcium isotopes as Earth, and this study showed that their hydrogen and nitrogen isotopes are similar to Earth's, too. In the study of extraterrestrial materials, the abundances of an element's isotopes are used as a distinctive signature to identify where that element originated.

Press release: <https://source.wustl.edu/2020/08/meteorite-study-suggests-earth-may-have-always-been-wet>

Paper: <https://www.science.org/doi/10.1126/science.aba1948>

Geological Evolution of Earth's Atmosphere: *Outgassing*

H_2

N_2

CO_2

H_2O

oceans
form

CO_2
dissolves



Outgassing

Life forms

4.5 Gy
before present

4 Gy

3.5 Gy

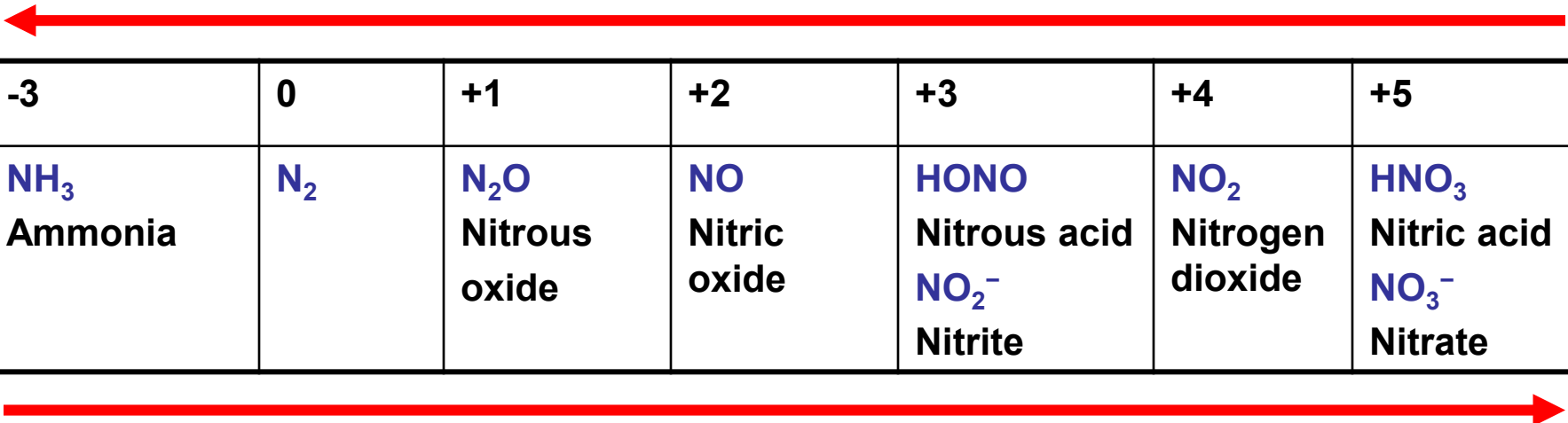
0.4 Gy

present

Geological Evolution of Earth's Atmosphere:

Early Atmosphere: Reducing Environment

Decreasing oxidation number (reduction reactions)



-3	0	+1	+2	+3	+4	+5
NH₃ Ammonia	N₂	N₂O Nitrous oxide	NO Nitric oxide	HONO Nitrous acid NO₂⁻ Nitrite	NO₂ Nitrogen dioxide	HNO₃ Nitric acid NO₃⁻ Nitrate

Increasing oxidation number (oxidation reactions)

Oxidation state represents number of electrons:
added to an element (– oxidation state) or
removed from an element (+ oxidation state)

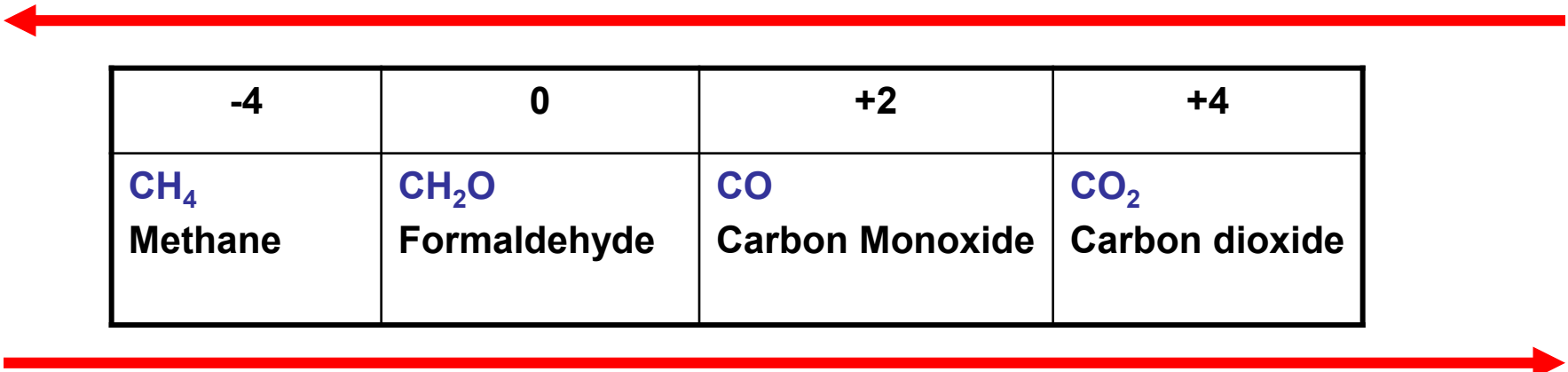
Oxidation state of a compound: $\sum = -2 \times \# \text{ O atoms} + 1 \times \# \text{ H atoms}$;
Oxidation of element = Electrical Charge – \sum

Note: there are some exceptions to this rule, such as oxygen in peroxides

Geological Evolution of Earth's Atmosphere:

Early Atmosphere: Reducing Environment

Decreasing oxidation number (reduction reactions)



-4	0	+2	+4
CH₄ Methane	CH₂O Formaldehyde	CO Carbon Monoxide	CO₂ Carbon dioxide

Increasing oxidation number (oxidation reactions)

Oxidation state represents number of electrons:
added to an element (– oxidation state) or
removed from an element (+ oxidation state)

Oxidation state of a compound: $\sum = -2 \times \# \text{ O atoms} + 1 \times \# \text{ H atoms}$;
Oxidation of element = Electrical Charge – \sum

Note: there are some exceptions to this rule, such as oxygen in peroxides

Geological Evolution of Earth's Atmosphere:

Early Atmosphere: Reducing Environment

How do we know early atmosphere was reducing ?

Why was a reducing environment important ?

Geological Evolution of Earth's Atmosphere: *Onset of Photosynthesis*

H_2

N_2

CO_2

H_2O

*Incipient rise of O_2 in the ancient atmosphere
signaled by first appearance of continental red-beds
of ferric iron (about 2 Gy B.P.)*

oceans
form

CO_2
dissolves

O_2

Outgassing

Life forms

Onset of
photosynthesis

4.5 Gy
before present

4 Gy

3.5 Gy

0.4 Gy

present



Geological Evolution of Earth's Atmosphere:

Atmospheric O₂ on Geological Time Scales

- Rise of atmospheric O₂ linked to evolution of life:

The rise of atmospheric O₂ that occurred ~2.4 billion years ago was the greatest environmental crisis the Earth has endured. [O₂] rose from one part in a million to one part in five: from 0.00001 to 21% ! Earth's original biosphere was like an alien planet. Photosynthetic bacteria, frantic for hydrogen, discovered water and its use led to the build up of atomic O, a toxic waste product.

Many kinds of microbes were wiped out. O and light together were lethal. The resulting O-rich environment tested the ingenuity of microbes, especially those non-mobile microorganisms unable to escape the newly abundant reactive atmospheric gas. The microbes that survived invented various intracellular mechanisms to protect themselves from and eventually exploit this most dangerous pollutant.

Lynn Margulis and Dorion Sagan, *Microcosmos: Four Billion Years of Microbial Evolution*, 1986

The rise of atmospheric oxygen led to something else critical to
“life as we know it” – what did rising [O₂] lead to ?!?

Geological Evolution of Earth's Atmosphere: *Atmospheric O₂ on Geological Time Scales*

- Rise of atmospheric O₂ linked to evolution of life:

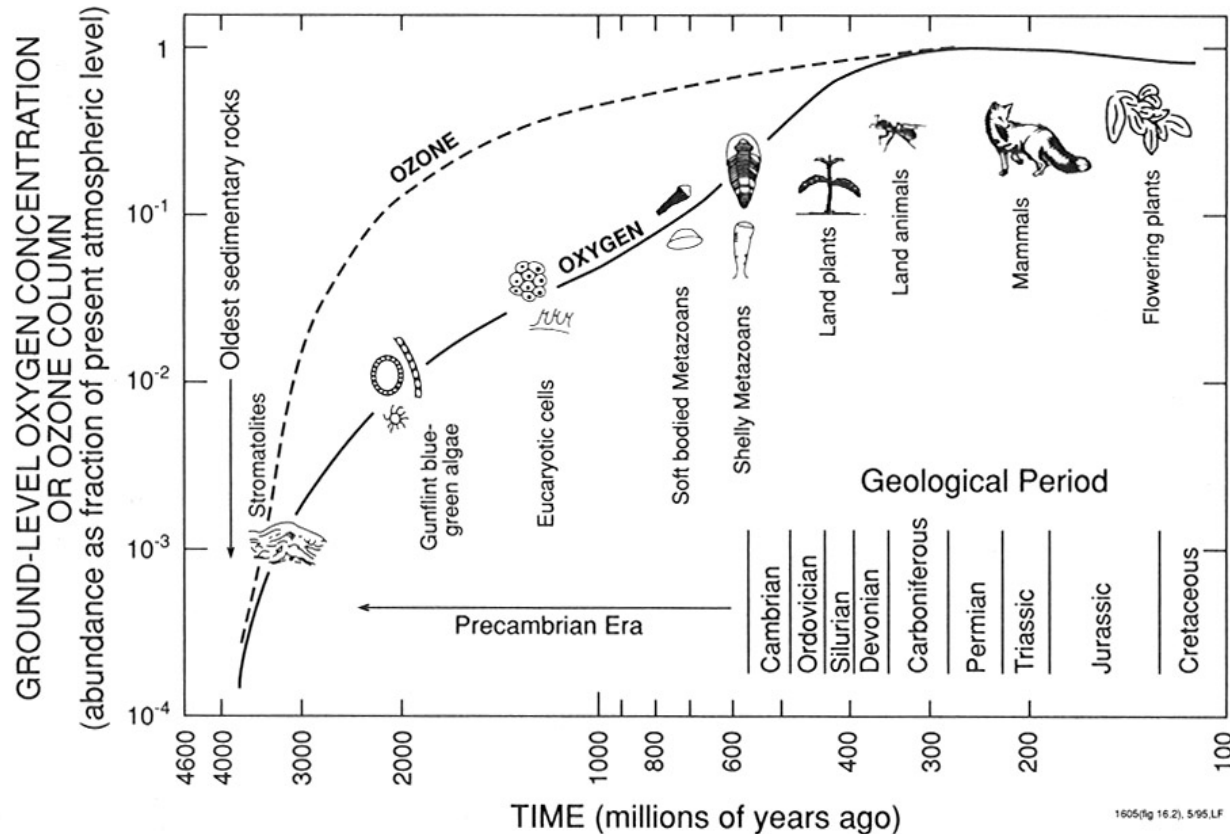


Figure 16.3. Probable evolution of the oxygen and ozone abundance in the atmosphere (fraction of present levels) during the different geological periods of the Earth's history (Wayne, 1991; reprinted by permission of Oxford University Press).

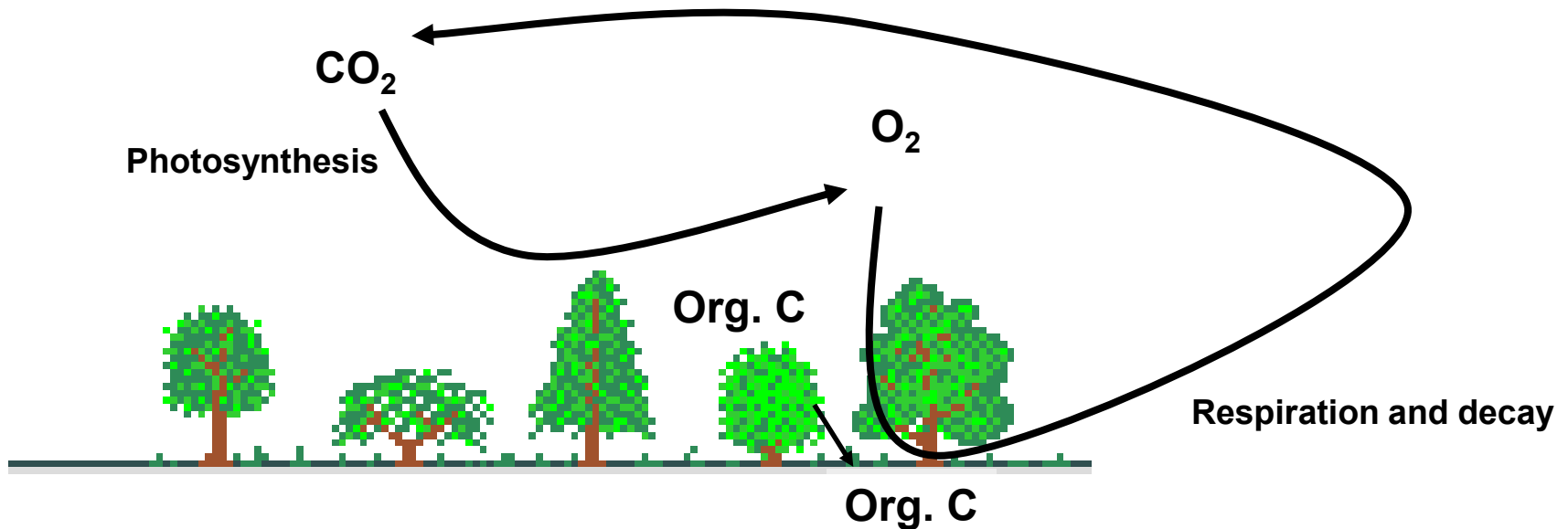
Geological Evolution of Earth's Atmosphere:

Early Atmosphere: Photosynthesis

- **Photosynthesis: Source of O₂**



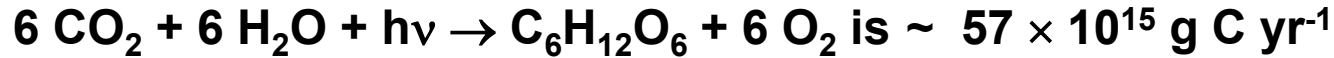
- **Respiration and Decay: Sink of O₂**



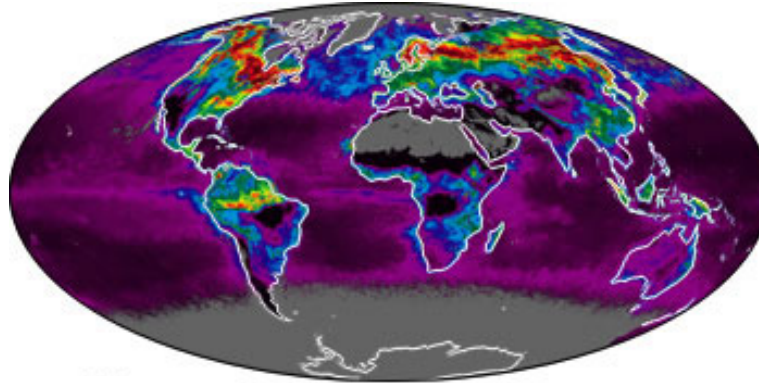
Geological Evolution of Earth's Atmosphere:

Early Atmosphere: Photosynthesis

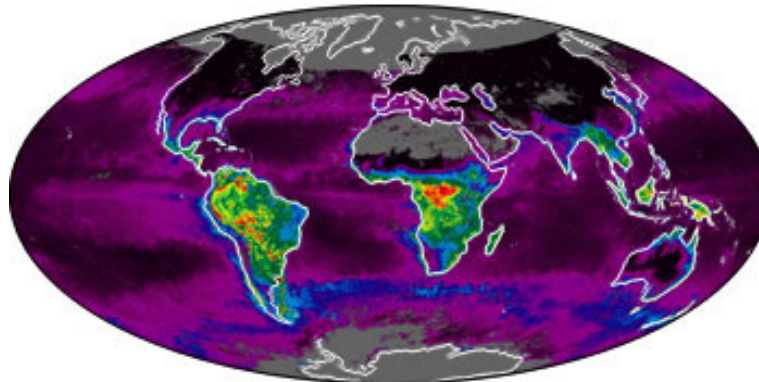
- Net primary productivity of organic matter:



Imhoff *et al.*, *Nature*, 2004



<http://www.globalcarbonproject.org/science/figures/FIGURE9.htm>



Geological Evolution of Earth's Atmosphere:

Early Atmosphere: Photosynthesis

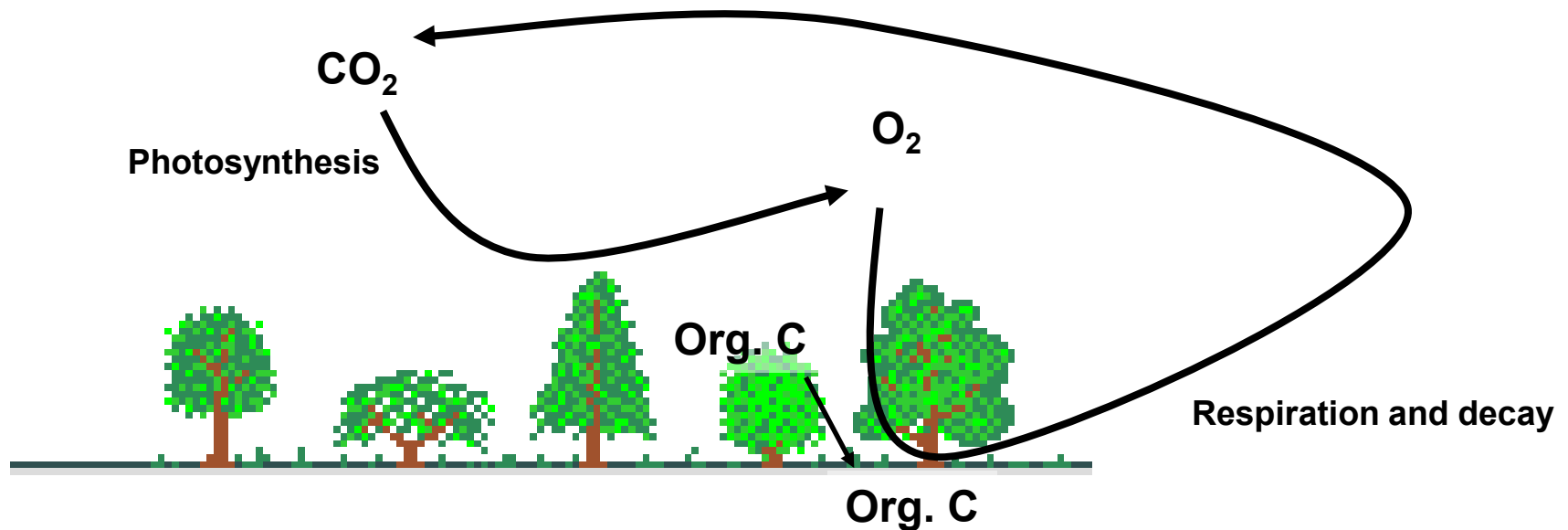
- Net primary productivity of organic matter:



Production of atmospheric O_2 is therefore $\sim 152 \times 10^{15} \text{ g O}_2 \text{ yr}^{-1}$ Flux

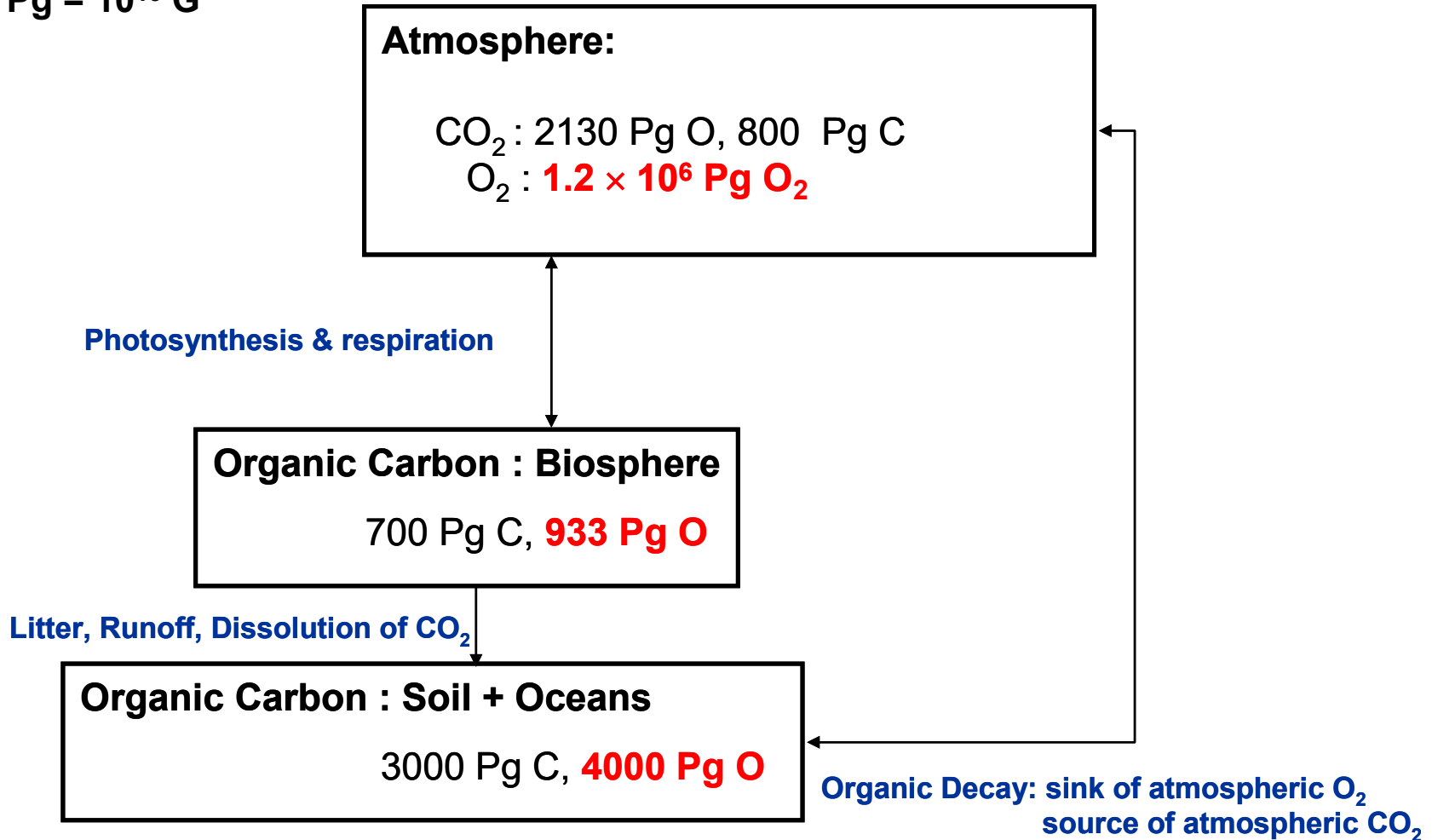
- Mass O_2 in atmosphere = $0.21 \times (5.2 \times 10^{21} \text{ g}) \times (32 / 28.8) \approx 1.2 \times 10^{21} \text{ g}$ Amount

- Lifetime of atmospheric O_2 due to biology = Amount / Flux



Geological Evolution of Earth's Atmosphere: *Oxygen and Carbon Reservoirs*

1 Pg = 10^{15} G



Atmospheric O₂ reservoir much larger than O₂ content of biosphere, soils, and ocean; therefore, some *other process* must control atmospheric O₂

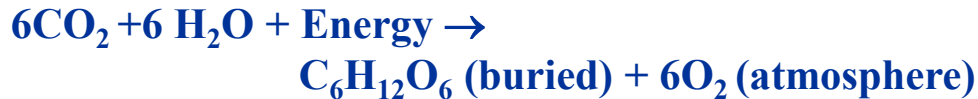
Geological Evolution of Earth's Atmosphere: *Oxygen Reservoirs & Pathways*

1 Pg = 10^{15} G

Atmosphere:

$O_2 : 1.2 \times 10^6 \text{ Pg } O_2$

Burial of organic matter is source of atmospheric O_2 :



Sediments: Buried Organic Carbon

$O_2 : \sim 32 \times 10^6 \text{ Pg } O$

O_2 Lifetime \approx 4 million years

Weathering of mantle is sink of atmospheric O_2 :

For example:



**Crust and Mantle: Oxides of Fe, Si, S, Mg, etc:
 FeO , Fe_2O_3 , $FeSiO_3$, SiO_4 , MgO , etc**

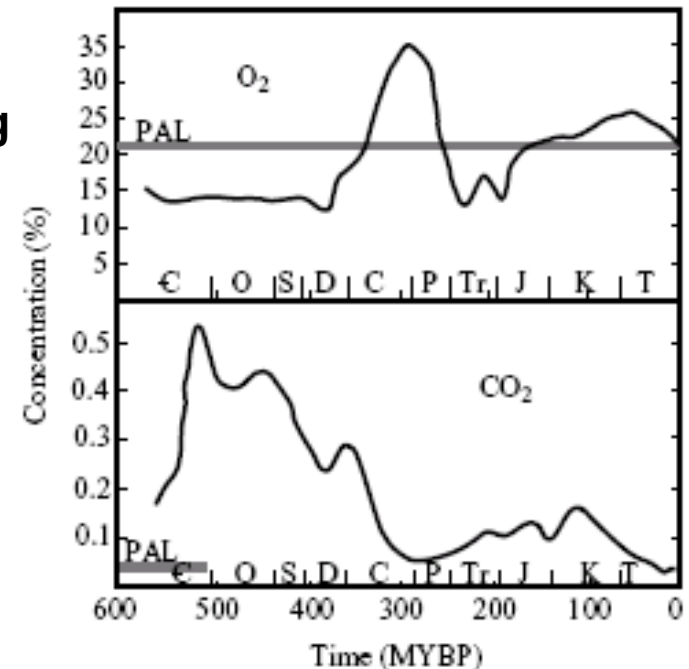
This is where the bulk of the oxygen resides!

Geological Evolution of Earth's Atmosphere: *Atmospheric O₂ on Geological Time Scales*

- **Rise of atmospheric O₂ linked to evolution of life:**
 - 400 My B.P. O₂ high enough to form an ozone layer
 - 400 to 300 My B.P.: first air breathing lung fish & primitive amphibians
- **On geological timescales, level of O₂ represents balance between burial of organic C & weathering of sedimentary material:**

(see Chapter 12, “Evolution of the Atmosphere” in *Chemistry of the Natural Atmosphere* by P. Warneck (2nd ed) for an excellent discussion)
- **Present atmosphere is oxidizing:**

CH₄ ⇒ CO₂ with time scale of ~9 years



From R. Dudley, Atmospheric O₂, Giant Paleozoic Insects, and the Evolution of Aerial Locomotor Performance, *J. Exper. Biol.*, 201, 1043, 1998.

Geological Evolution of Earth's Atmosphere:

Atmospheric CO₂ on Geological Time Scales

~500 to 300 My B.P.

- **Development of vascular land plants**
- **Plants became bigger and bigger and less reliant on water**
- **Once buried, lignin in woody material resists decay**
- **Burial rate of terrestrial plant matter increases dramatically:
(evidence : $\delta^{13}\text{C}$ analysis)**
- **Past burial rate of vascular plant material may have been much higher than present, due to the lack (way back when) of abundant bacteria, fungi, and small soil animals that now recycle plant matter**

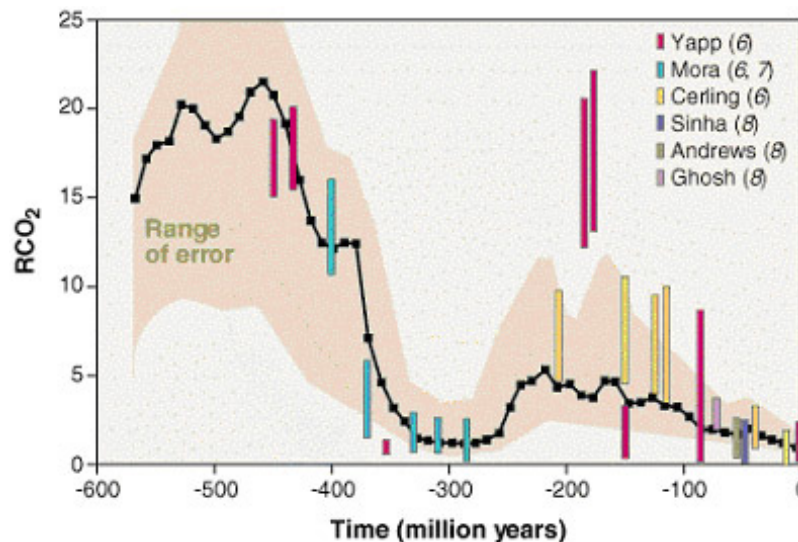
Non-vascular: Bryophytes

Vascular: Pteridophytes

Geological Evolution of Earth's Atmosphere: *Atmospheric CO₂ on Geological Time Scales*

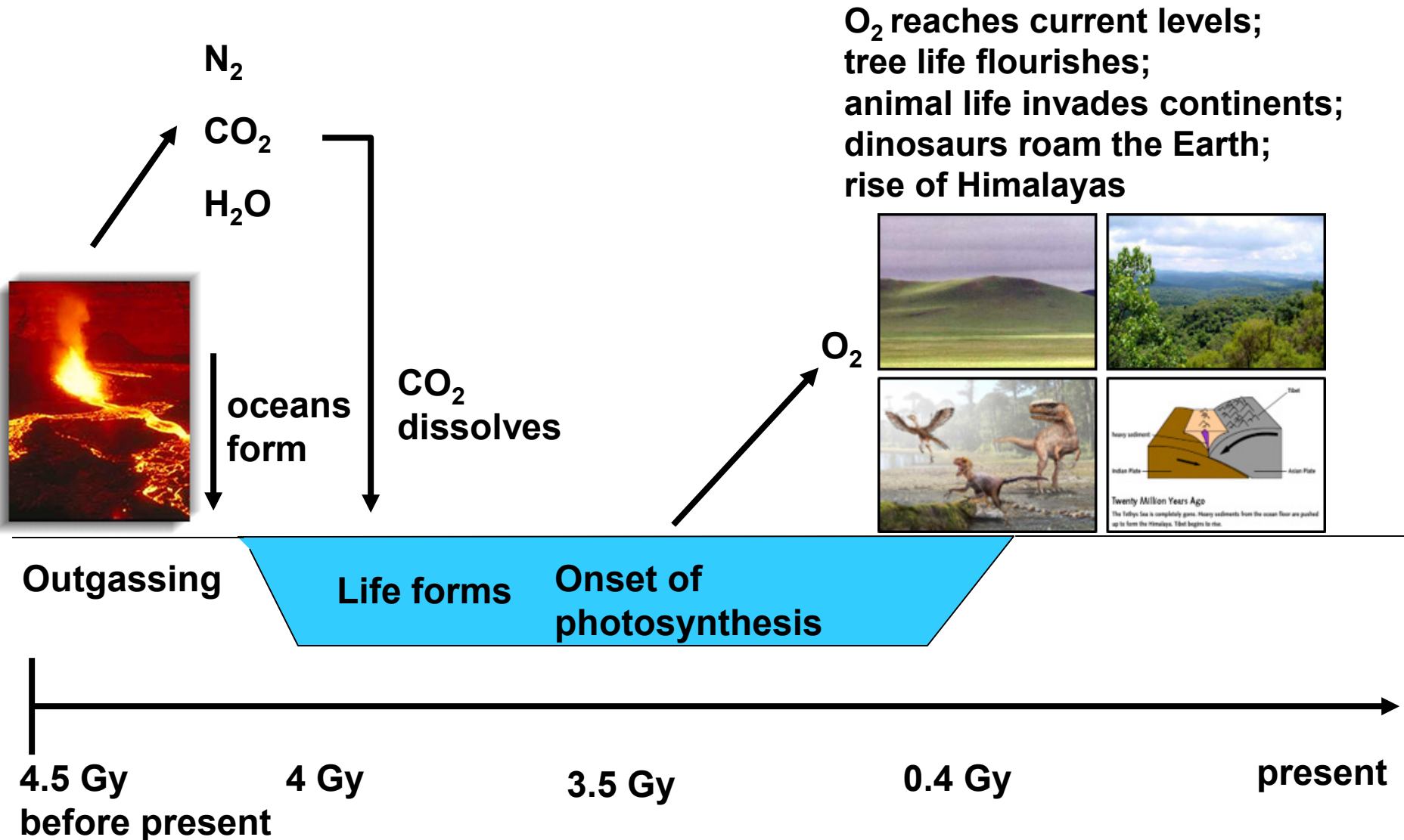
~500 to 300 My B.P.

- Development of vascular land plants
- Plants became bigger and bigger and less reliant on water
- Once buried, lignin in woody material resists decay
- Burial rate of terrestrial plant matter increases dramatically:
(evidence : $\delta^{13}\text{C}$ analysis)
- Past burial rate of vascular plant material may have been much higher than present, due to the lack (way back when) of abundant bacteria, fungi, and small soil animals that now recycle plant matter



From R. Berner, *Science*, 276, 544, 1997.

Geological Evolution of Earth's Atmosphere: *Precursors of Modern Day World*



Geological Evolution of Earth's Atmosphere: CO₂ and Temperature

What message were we trying to convey?

Climate History, 500 Million ybp to Present

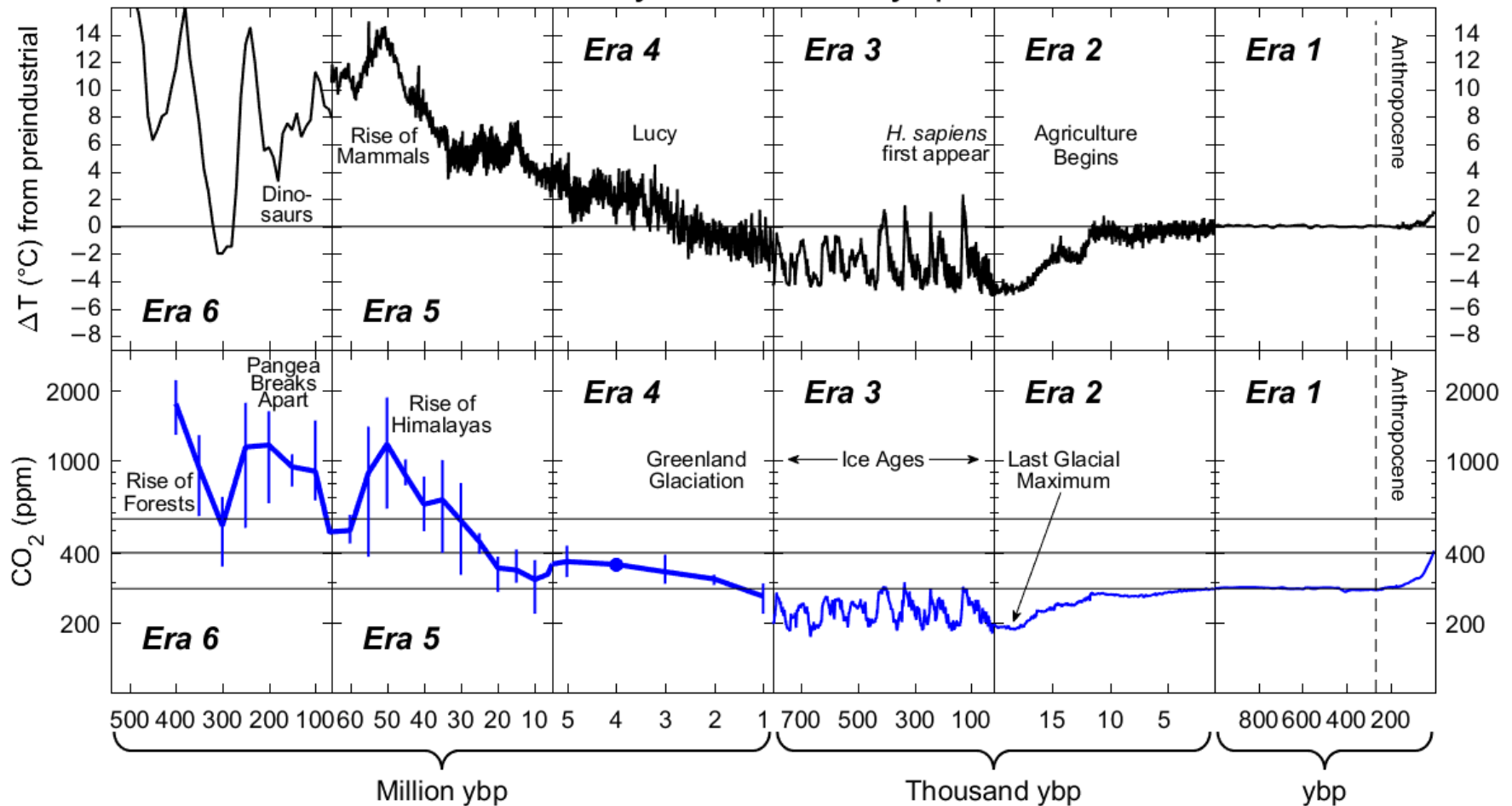
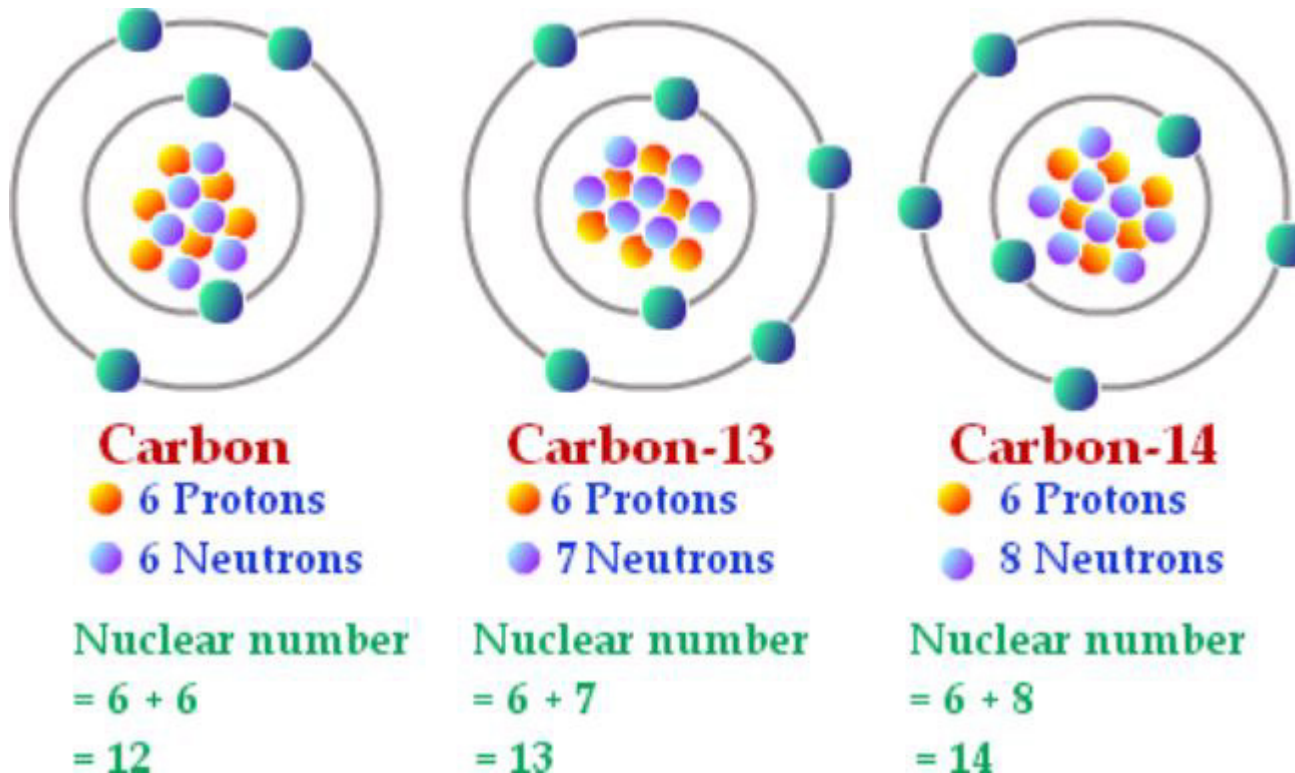


Fig 1.1, *Paris Beacon of Hope*

Geological Evolution of Earth's Atmosphere: One Day, *Everything Changed*

Biology prefers light forms of carbon:



<https://experiment.com/u/iA41fA>

Geological Evolution of Earth's Atmosphere: One Day, *Everything Changed*

By understanding how the carbon isotopic ratio of the world's surface waters changed at the K-T boundary, as recorded by the shells of preserved oceanic organisms, we could compute the fraction of the world's biosphere that must have burned on this really bad day (or soon thereafter):

Carbon isotopic evidence for biomass burning at the K-T boundary

A new interpretation of existing carbon isotopic data combined with results from a biogeochemical model suggests that burning of terrestrial biomass occurred on a global scale at the Cretaceous-Tertiary (K-T) boundary. Carbon isotopic ratios from planktonic and benthic microfossils across the K-T boundary reveal not only a breakdown in the normal surface-water to deep-water gradient of $^{13}\text{C}/^{12}\text{C}$, but also a reversal at the boundary. This reversal cannot be explained by the cessation of primary production alone. We propose that combustion of terrestrial biomass with subsequent transfer of isotopically light carbon to surface waters is the most likely cause of this anomaly. A biogeochemical model is used to quantify the extent of burning at the boundary: combustion of roughly 25% of the above-ground biomass at the end of the Cretaceous is necessary to account for the observed isotopic signal.

Ivany and Salawitch, *Geology*, 1993

Link to this paper appears in auxiliary reading for today's class

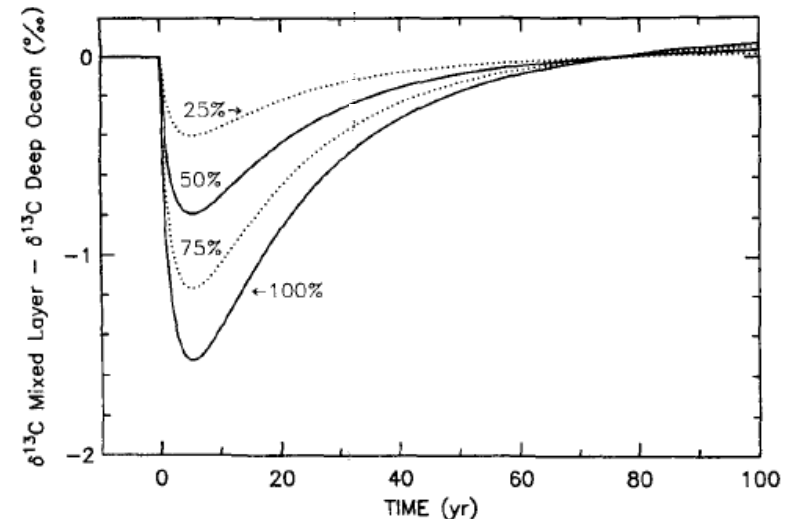
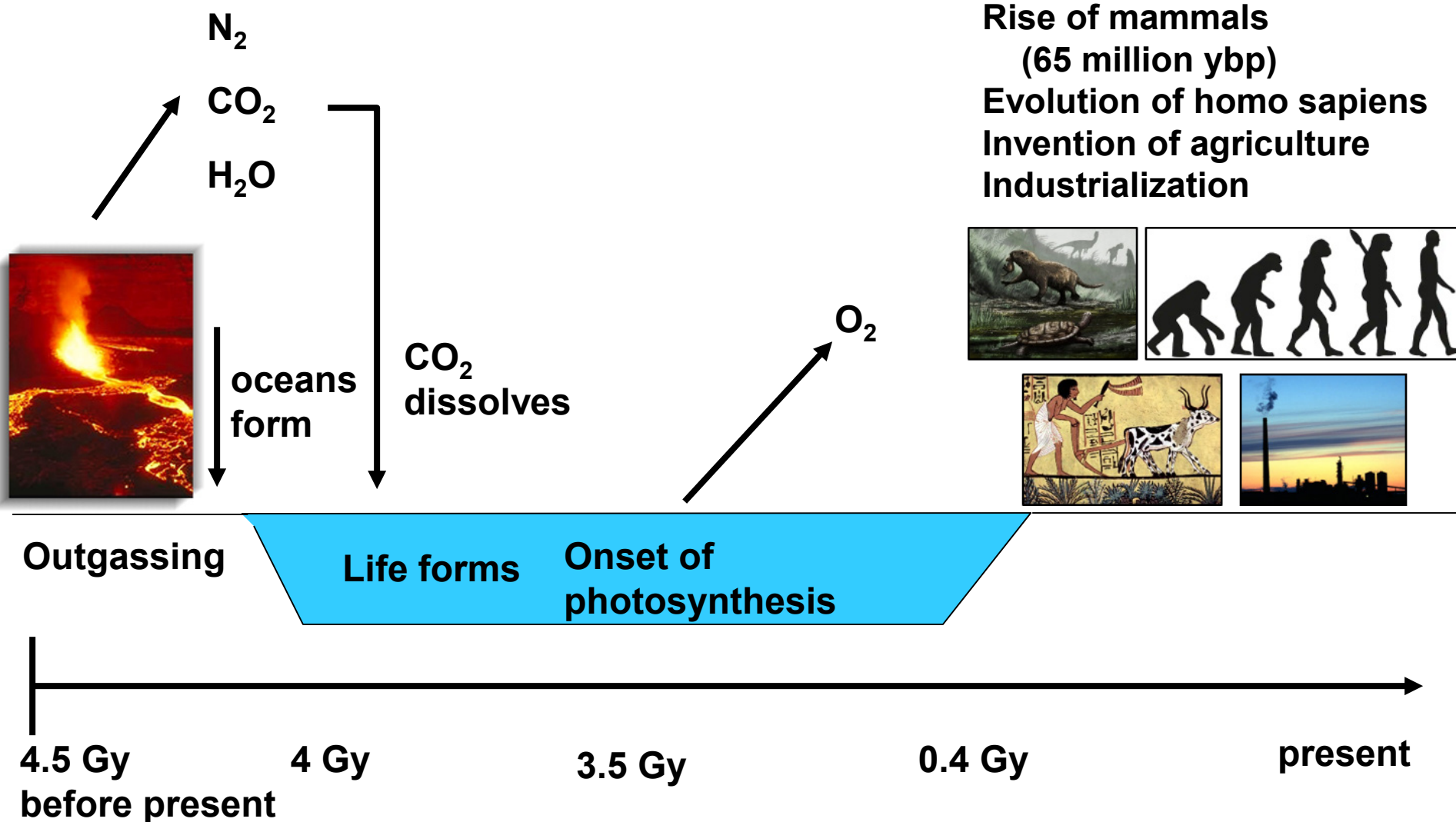


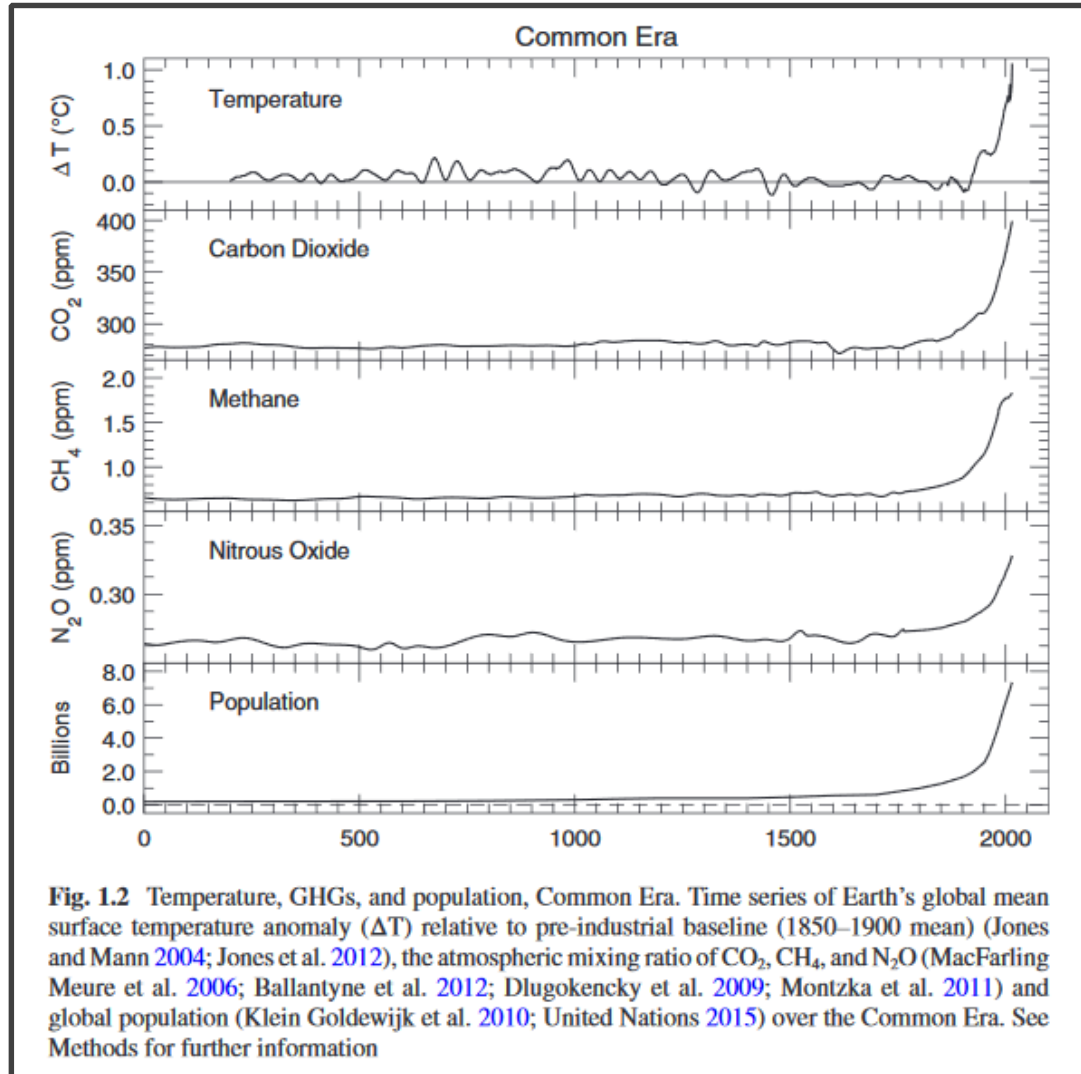
Figure 2. Variation of surface-water to deep-water gradient of $\delta^{13}\text{C}$ vs. time for simulations of biomass combustion, assumed to occur instantaneously at time zero. Results are shown for burning 25%, 50%, 75%, and 100% of above-ground biomass (10^{18} g C; $\delta^{13}\text{C} = -25.7\%$) at end of Cretaceous assuming combustion efficiency of 50% (i.e., model result for 100% combustion corresponds to injection of half of above-ground biomass carbon into atmosphere as CO_2 at time zero).

Geological Evolution of Earth's Atmosphere: *Human Influence*



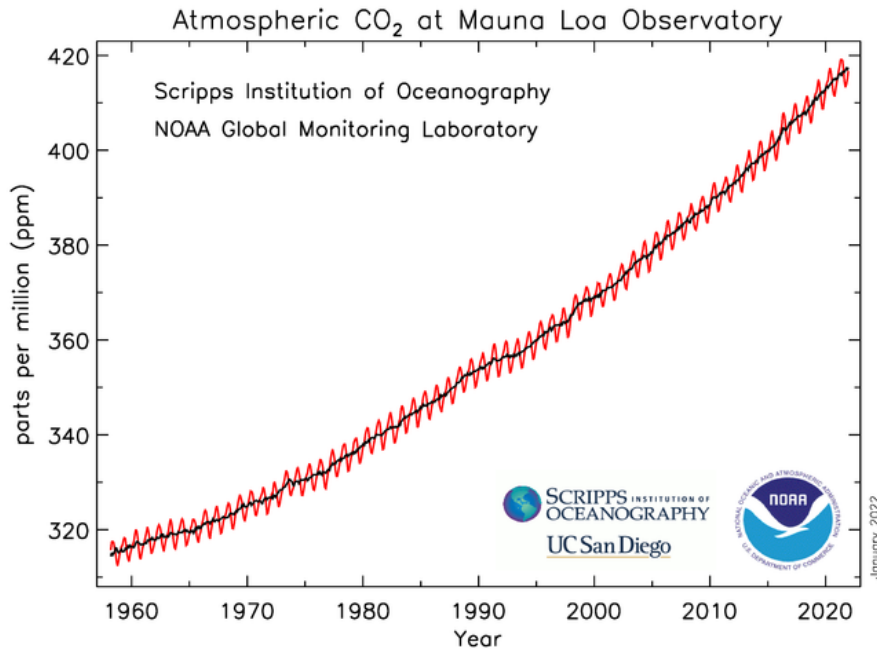
Geological Evolution of Earth's Atmosphere: *Human Influence*

What message were we trying to convey?



Earth's Atmosphere – Effect of Humans

CO₂: ~415 parts per million (ppm) and rising !



Charles Keeling, Scripps Institution of Oceanography, La Jolla, CA
https://www.esrl.noaa.gov/gmd/webdata/ccgg/trends/co2_data_mlo.png

Human drivers of global warming over the last millennium

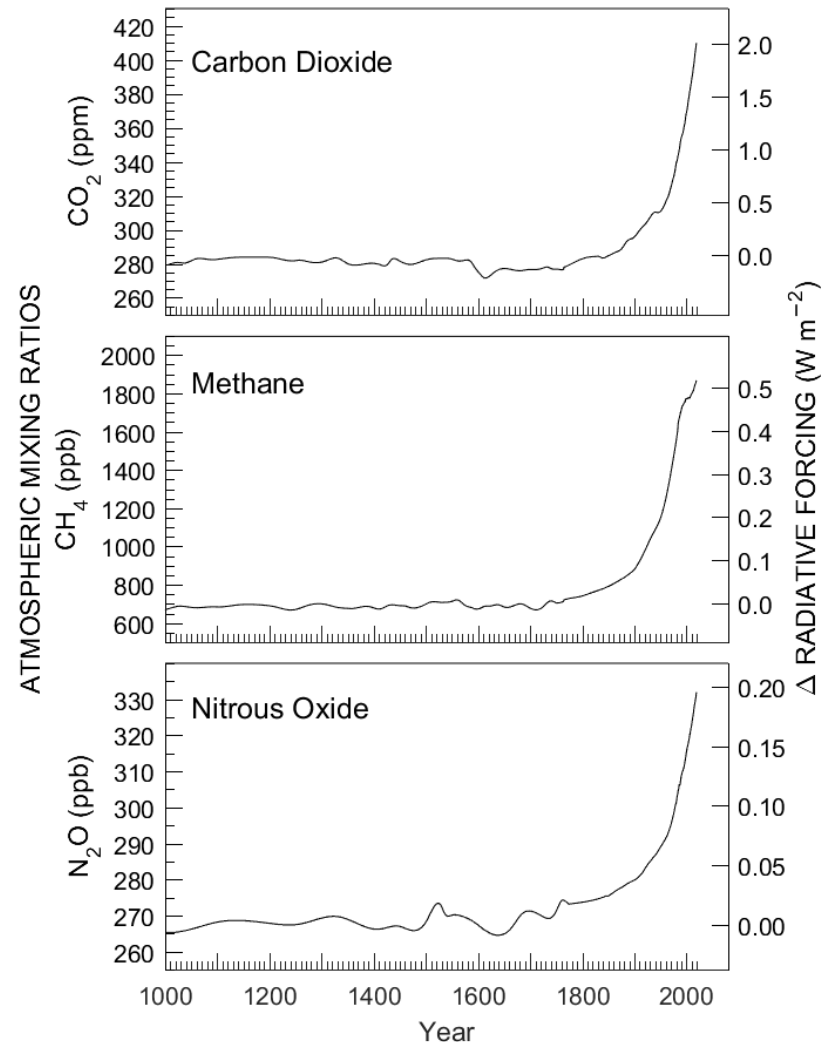
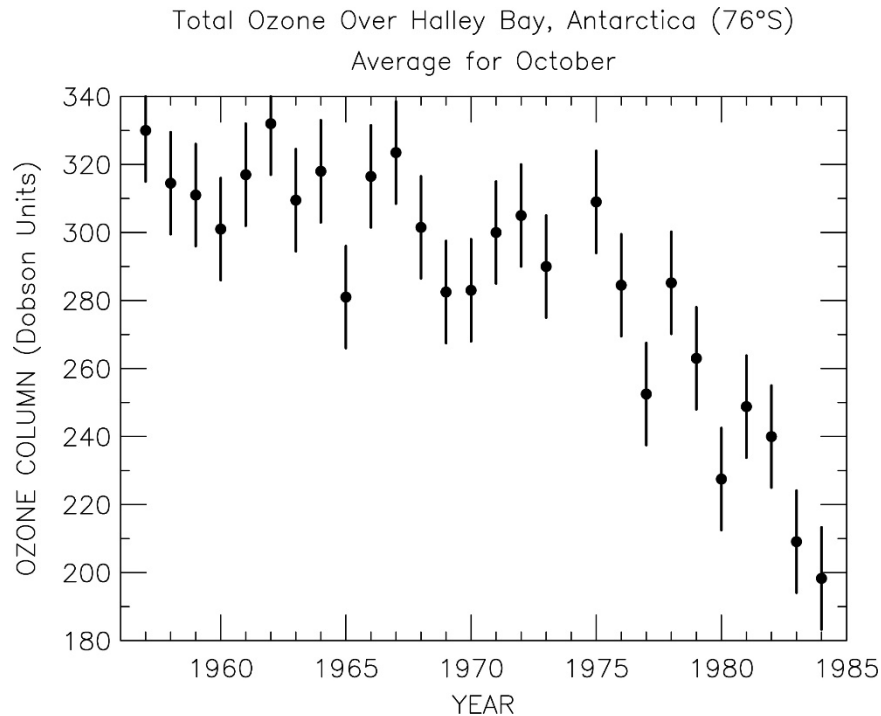


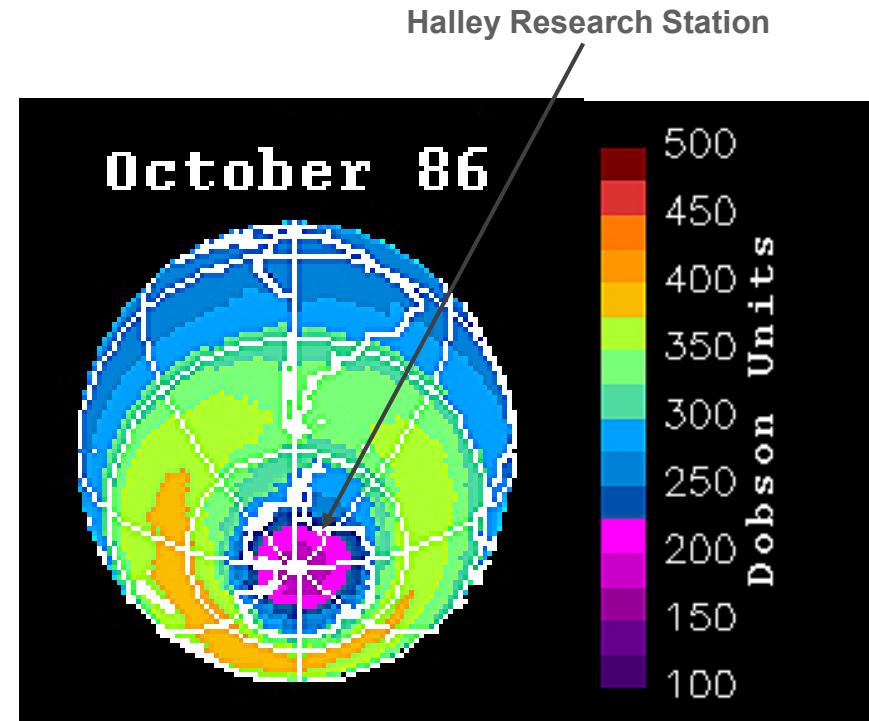
Figure courtesy Brian Bennett, Univ of Md

Earth's Atmosphere – Effect of Humans

Stratospheric Ozone – shields surface from solar UV radiation



After Farman *et al.*, Large losses of total ozone in Antarctica reveal Seasonal ClO_x/NO_x interaction, *Nature*, 315, 207, 1985.

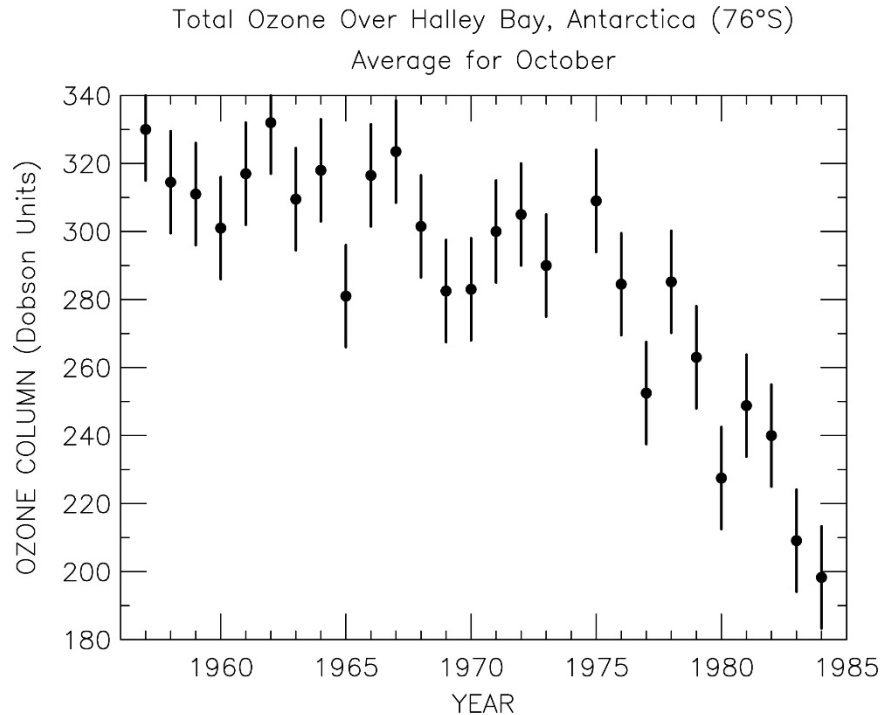


Stolarski *et al.*, *Nature*, 1986.

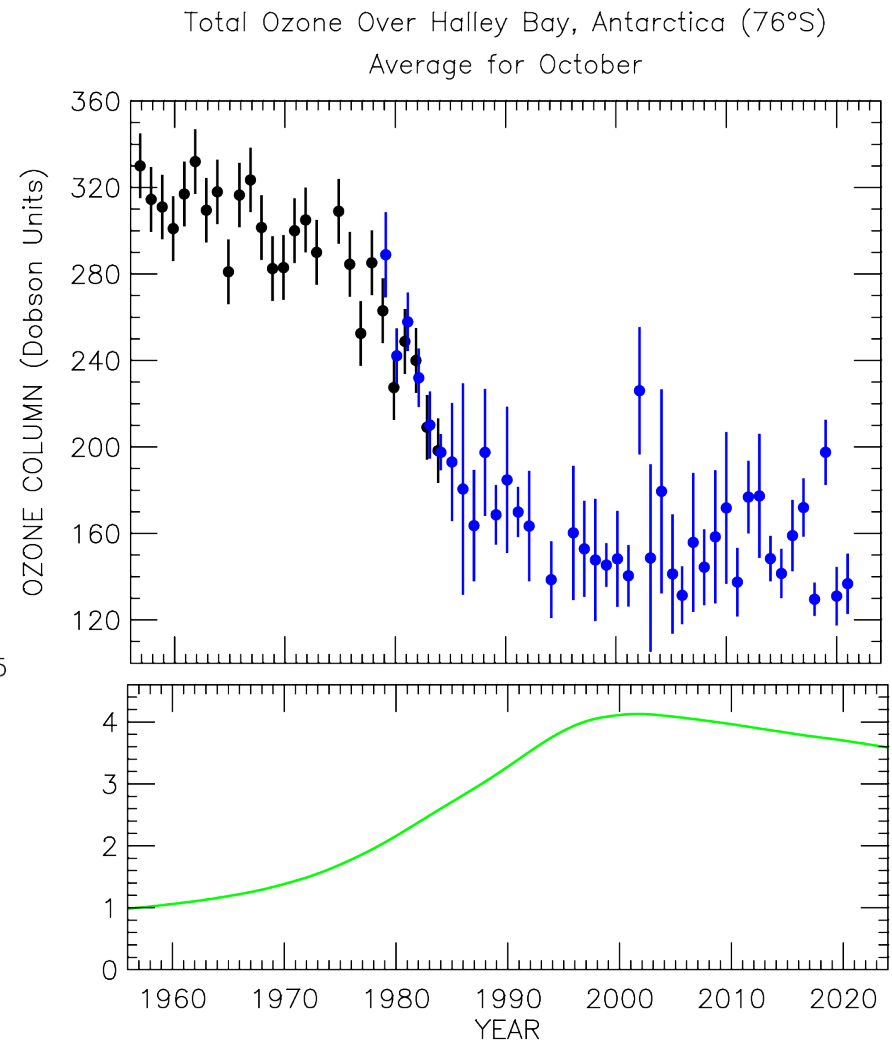
Earth's Atmosphere – Effect of Humans

Stratospheric Ozone – shields surface from solar UV radiation

Update

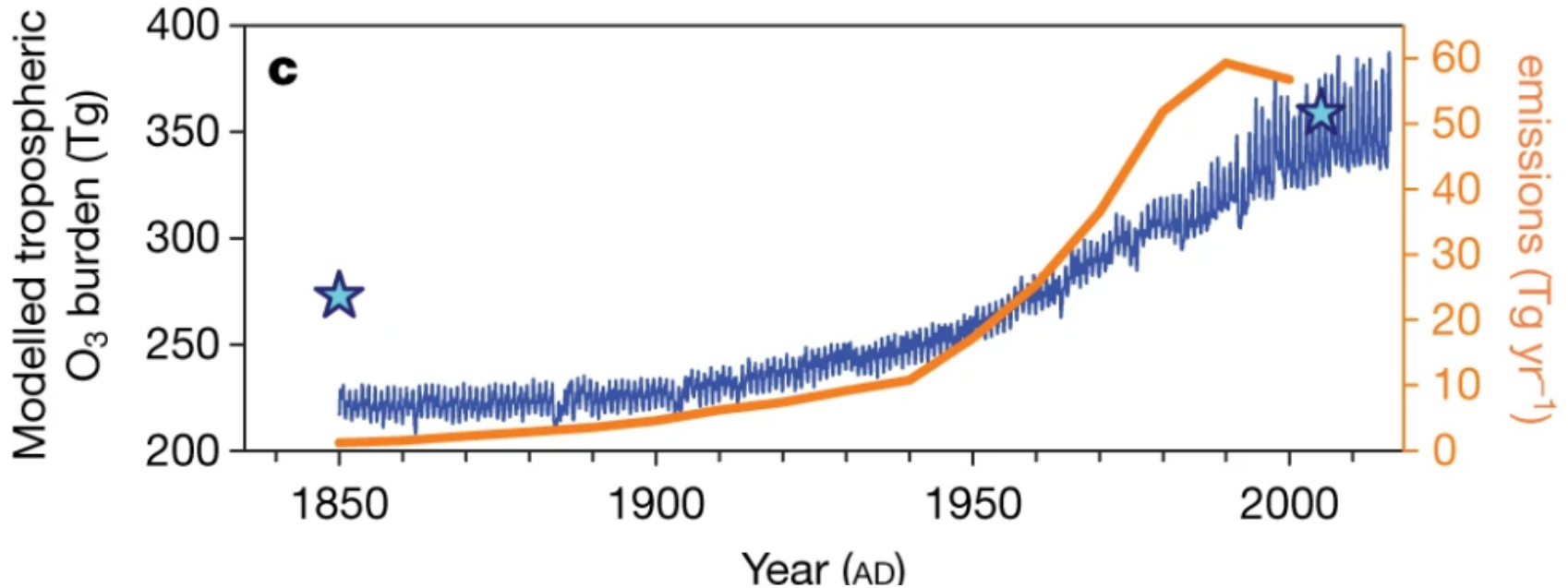


After Farman *et al.*, Large losses of total ozone in Antarctica reveal Seasonal ClO_x/NO_x interaction, *Nature*, 315, 207, 1985.



Earth's Atmosphere – Effect of Humans

Tropospheric Ozone – oxidant, lung irritant, harmful to crops



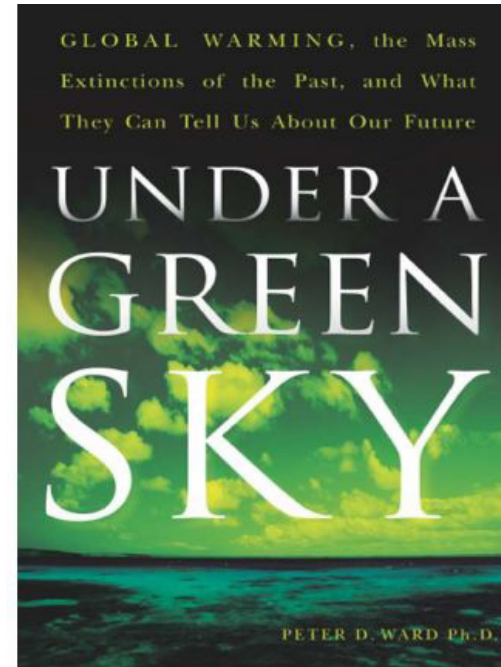
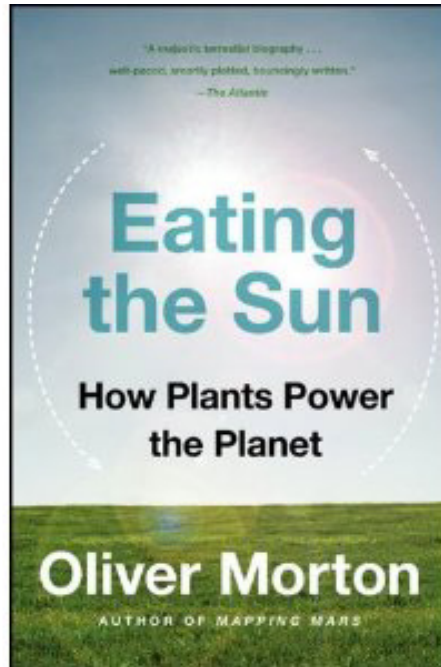
Modelled increase in tropospheric O₃ burden from the GEOS-Chem/MERRA2 model (cyan stars) and GISS-E2.1 model (blue line), as well as estimates of historical emissions of NO_x (orange line).

Yueng *et al.*, *Nature*, 2019

<https://www.nature.com/articles/s41586-019-1277-1>

Source Material

These books are a great resource for how photosynthesis works as well as the history of atmospheric composition

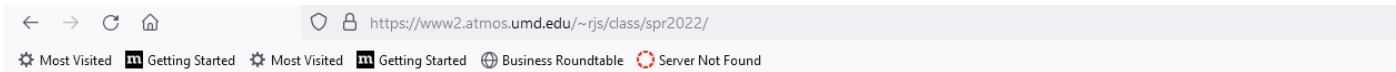


http://www.amazon.com/Eating-Sun-Plants-Power-Planet/dp/0007163657/ref=sr_1_1?s=books&ie=UTF8&qid=1359325940&sr=1-1&keywords=eating+the+sun

http://www.amazon.com/Under-Green-Sky-Warming-Extinctions/dp/0061137928/ref=sr_1_1?s=books&ie=UTF8&qid=1359326345&sr=1-1&keywords=under+a+green+sky

and provided some of the source material for much of this lecture

Next Lecture: Course Overview



2. Schedule

Date	Lecture Topic	Required Reading	Admis. Tickets	Lecture Notes	Problem Sets*	Additional Readings	Learning Outcome
01/25	Class Overview	No reading for first meeting	No AT	Lecture 0 Video			No Quiz
01/27	Geological Evolution of Earth's Atmosphere	Paris Beacon of Hope Sec 1.1, 1.2 (intro), and 1.2.1 (11.5 pages)	AT 1	Lecture 1 Video		Ivany and Salawitch, Geology, 1993 NOVA: The Day The Dinosaurs Died	Quiz 1
02/01	Overview of Global Warming, Air Quality, & Ozone Depletion	IPCC 2007 FAQ (1.1, 1.2, 1.3, 2.1, & 3.1) (11 pages) EPA AQI Brochure (11 pages) 20 QAs Ozone (Q1, 2, 7, & 14) (11 pages) Paris Beacon of Hope Sec 1.2.2 (3 pages)	AT 2	Lecture 2 Video		Kerr, Science, 2007 * Bell et al., EHP, 2006 * Montzka et al., Nature, 2018 Naming Convention for CFCs & Halons Entire IPCC 2007 FAQ Entire 20 QAs Ozone Movie Clip	Quiz 2

Readings: IPCC 2007 FAQ 1.1, 1.2, 1.3, 2.1, & 3.1 (11 pages)
 EPA Air Quality Guide (11 pages)
 20 QAs Ozone Layer Q1, 2, 7, and 14 (11 pages)
 Paris Beacon of Hope, Sect 1.2.2 (3 pages)

Note: 36 pages, about our norm

Admission Ticket for Lecture 2 is posted on ELMS