Review of Lectures 10 to 17 AOSC / CHEM 433 & AOSC / CHEM 633 Ross Salawitch

Class Web Sites:

http://www2.atmos.umd.edu/~rjs/class/spr2022 https://myelms.umd.edu/courses/137772



https://www.videoblocks.com/video/earth-sunset-spacewalk-view-from-space-station-r7dydlcsgjd23vml0

12 April 2022

Announcements

1. Exam on Thursday

Conceptual questions only: no calculators

Closed book; no notes

- 2. Unless a prior arrangement has been made, everyone is expected to show up on time, in person, on Thursday, with a writing implement (pen or pencil)
- 3. For this exam, the 633 exam will differ only very slightly from the 433 exam and none of the 633 questions will be based on material covered only in the assigned auxiliary readings
- 4. Review of Problem Set #3 will occur quickly today, prior to completion of our grading, due to poor turnout (only 1 student) of our prior evening review

Importance of Radicals

- With a few exceptions, the only reactions between molecules that proceed at appreciable rates are those involving at least one radical
- Radicals require significant energy to form: a bond must be broken
- Radical formation is tied to absorption of photons that "photodissociate" a compound, leading to radical formation

Initiation

$$O_2$$
 + photon \rightarrow O + O

Propagation

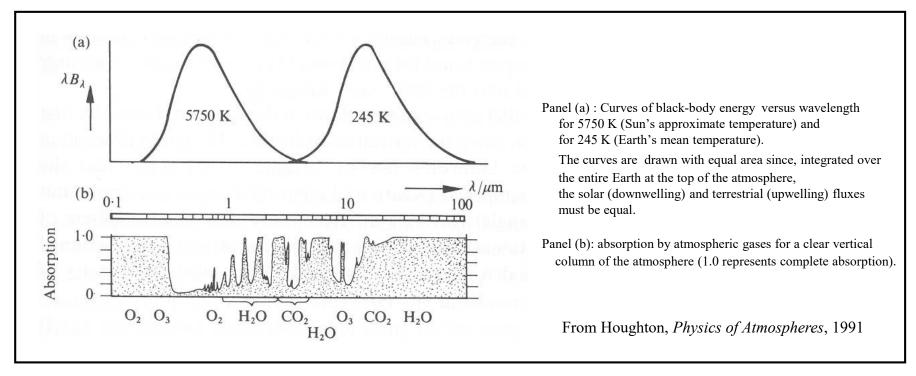
$$\begin{aligned} \text{O} + \text{O}_2 + \text{M} &\rightarrow \text{O}_3 + \text{M} \\ \text{O}_3 + \text{photon} &\rightarrow \text{O(1D)} + \text{O}_2 \\ \text{O(1D)} + \text{H}_2\text{O} &\rightarrow \text{OH} + \text{OH} \\ \text{OH} + \text{O}_3 &\rightarrow \text{HO}_2 + \text{O}_2 \\ \text{HO}_2 + \text{O} &\rightarrow \text{OH} + \text{O}_2 \end{aligned}$$

Termination

 $OH + HO_2 \rightarrow H_2O + O_2$

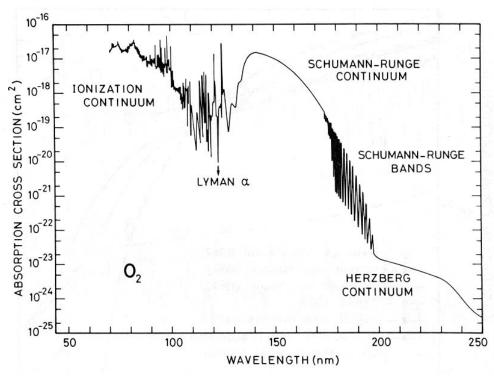
Atmospheric Radiation

• Solar irradiance (downwelling) at top of atmosphere occurs at wavelengths between ~200 and 2000 nm (~5750 K "black body" temperature)



• Absorption and photodissociation in the UV occurs due to changes in the electronic state (orbital configuration) of molecules

Absorption Cross Section of O₂



From Brasseur & Solomon, Aeronomy of the Middle Atmosphere, 1986

- O_2 can not dissociate longward of ~ 250 nm
- All of the absorption shown above is dissociative (e.g., leads to production of two O atoms)
- Structure in the O₂ cross section is related to whether the initial transition involves an unbound electronic state (smooth) or involves a specific vibrational level of an electronic state (banded, due to requirement of specific quanta of energy)

Optical Depth of O₂ Absorption

Recall the *Beer-Lambert Law*:

$$F(z,\lambda) = F_{TOA}(\lambda) e^{-\tau(z,\lambda)}$$
 (TOA : Top of Atmosphere)

where:

$$\tau(z, \lambda) = m \int_{z}^{\infty} \sigma_{\lambda} [C] dz' \qquad (\tau: optical depth)$$

$$\int_{0}^{\infty} [O_{2}] dz' \approx 4 \times 10^{24} \text{ molecules/cm}^{2}$$

Also:

$$\int_{0}^{\infty} [O_2] dz' \approx 4 \times 10^{24} \text{ molecules/cm}^2$$

O_2 Optical Depth for $\theta = 0^\circ$, $z = 0$ km					
	$\sigma_{\rm max}({\rm cm}^2)$	τ (0 km)	$e^{-\tau (0 \text{ km})}$		
Schumann-Runge Continuum	10^{-17}	4×10^7	0.		
Schumann-Runge Bands	10^{-20}	4×10^4	0.		
	3×10^{-23}	120	7.6×10^{-53}		
Herzberg Continuum	10^{-23}	40	4.2×10^{-18}		

Optical Depth of O₃ Absorption

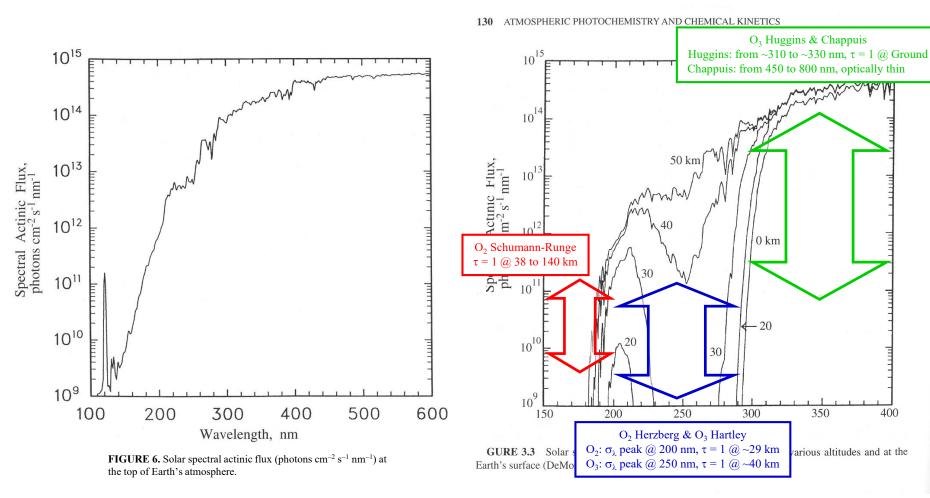
A typical mid-latitude column abundance for O₃ is 300 Dobson units (DU):

$$1 \text{ DU} = 2.687 \times 10^{16} \text{ molecules/cm}^2$$
; $300 \text{ DU} = 8 \times 10^{18} \text{ molecules/cm}^2$

$$\frac{\text{Column O}_3}{\text{Column Air}} = \frac{8 \times 10^{18}}{2 \times 10^{25}} = 0.4 \text{ parts per million } \Rightarrow \text{Ozone is a trace species!}$$

O ₃ Optio	O_3 Optical Depth for $\theta = 0^\circ$, $z = 0$ km					
	$\sigma_{\rm max} ({\rm cm}^2)$	τ (0 km)	e ^{-τ (0 km)}	O_3 Column, $\tau = 1.0$		
Hartley (~220 to 280 nm)	10^{-17}	80	1.8×10^{-35}	3.7 DU		
Huggins (~310 to 330 nm)	10^{-19}	0.8	0.45	372 DU		
Chappuis (~500 to 700 nm)	3×10^{-21}	0 .024	~1.0	12,400 DU		

Solar Spectral Actinic Flux



From DeMore et al., Chemical Kinetics and Photochemical Data for Use in Stratospheric Modeling, Evaluation No. 11, 1994.

From Seinfeld and Pandis, Atmospheric Chemistry and Physics, 1998.

O₃ Huggins & Chappuis

0 km

20

350

various altitudes and at the

400

Photolysis Frequency

For a specific spectral interval, the photolysis frequency (*partial J value*) of a gas is given by the product of its absorption cross section and the solar irradiance:

$$\begin{split} J_{gas}(z,\!\lambda) &= Quantum_Yield(\lambda) \; \sigma_{gas} \; (\lambda,\!T) \; F(z,\!\lambda) \\ &\quad Units: \; s^{-1} \; nm^{-1} \end{split}$$

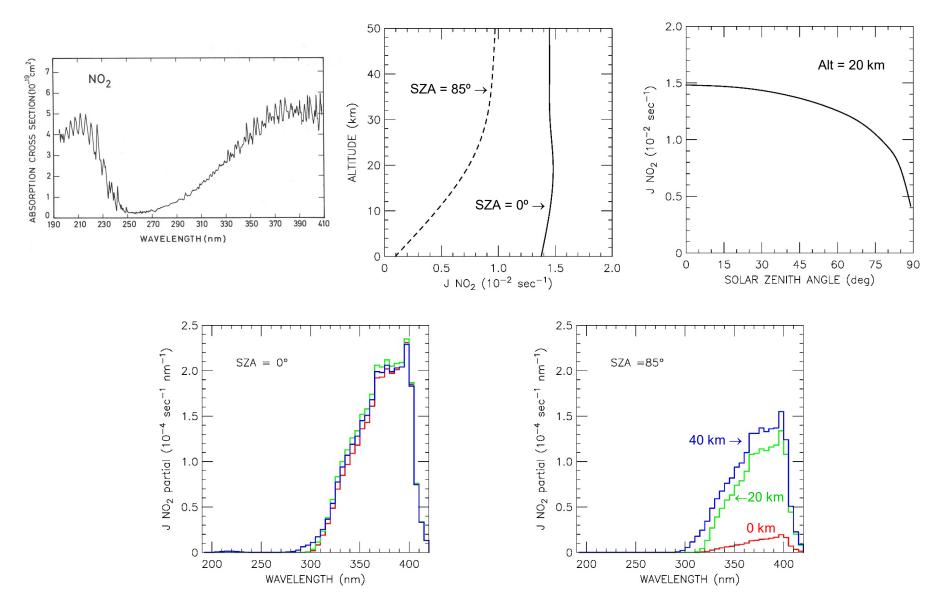
The total *photolysis frequency* (*J value*) is found by integrating $J_{gas}(z,\lambda)$ over all wavelengths for which the gas photodissociates:

$$J_{gas}(z) = \int_{\lambda}^{\lambda_{max}} J_{gas}(z, \lambda) d\lambda$$
Units: s⁻¹

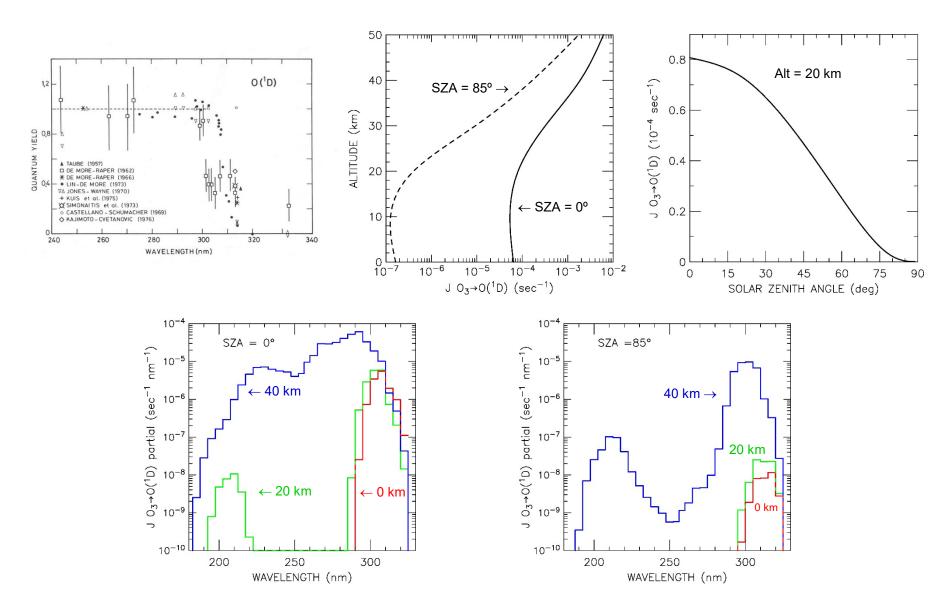
Rate of Reaction =
$$\frac{dO_3}{dt} = J$$
 [O₃]; Units of J are s⁻¹

More precisely, calculations of photolysis frequencies consider the "spectral actinic flux", which represents the amount of available photons integrated over all angles, rather than "solar irradiance". These two quantities differ because of scattering of solar radiation by gases and aerosols, and reflection of radiation by clouds and the surface.

NO₂ Photolysis



$O_3 \rightarrow O(^1D)$ Photolysis



Bimolecular Gas Phase Reactions

8.9 kcal/mole

35.1 kcal/mole

-17.8 kcal/mole

-57.8 kcal/mole

OH +
$$CH_4 \rightarrow CH_3 + H_2O$$

$$CH_3$$

 Δ Enthalpy = -13.8 kcal/mole

Exothermic!

Rate of Reaction =
$$\frac{dCH_4}{dt} = k \text{ [OH][CH_4]}$$

Arrhenius Expression for rate constant:

$$k = 1.85 \times 10^{-12} \times e^{-1690/T} \text{ cm}^3 \text{ sec}^{-1}$$

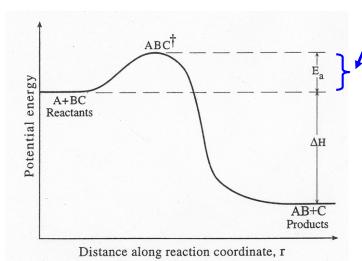


Figure 3.3 Barrier energies for the forward reaction (E_a) and the reverse reaction $(E_a + \Delta H)$.

E_△ / R ⇒ Activation Energy / Gas Constant

Energy Term

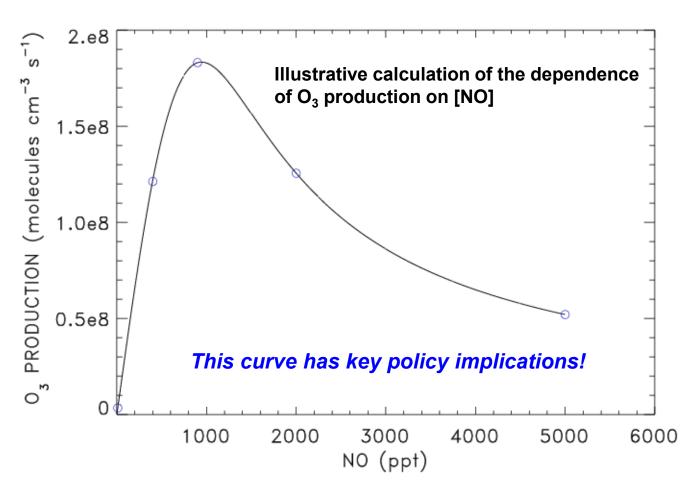
R = 8.3143×10^7 erg / (K mole) = 2.87×10^6 erg / (K gm) for air

Yung and DeMore, Photochemistry of Planetary Atmospheres, Oxford, 1999.

Tropospheric Ozone Production versus NO

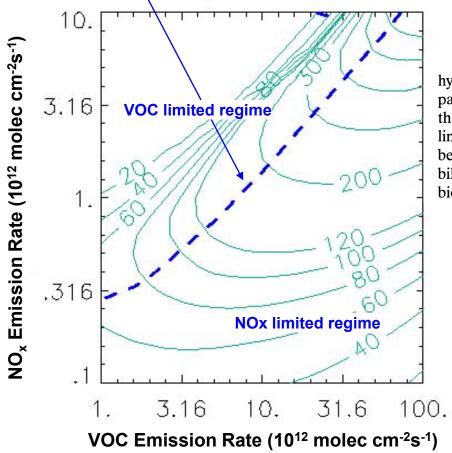
Production of Tropospheric O_3 limited by: ______ ? As NO_x rises:

[HO₂] falls faster than [NO] rises, leading to a decrease in the value _____



Tropospheric Ozone Production versus NO_x and VOCs

Ridge: local maximum for O₃ that separates the NOx-limited regime from and VOC limited regime

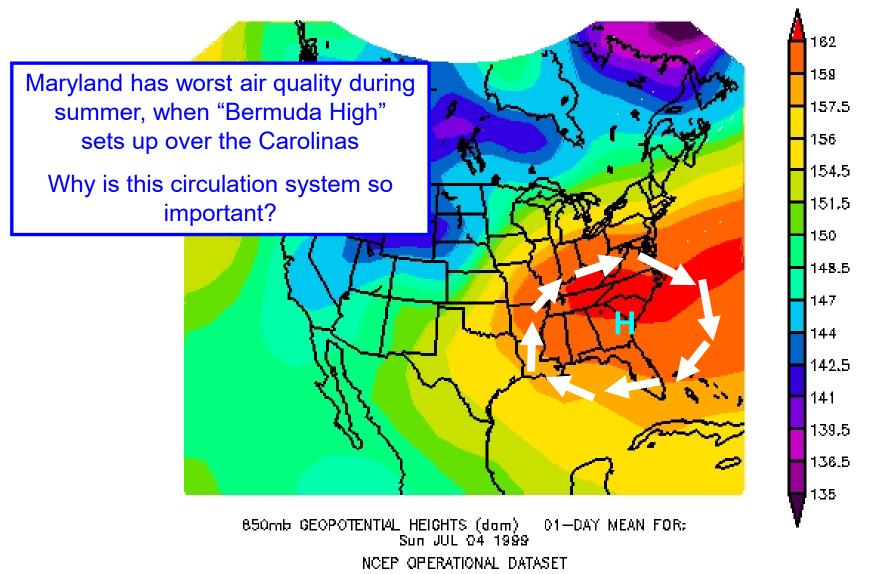


An important discovery in the past decade is that the focus on hydrocarbon emission controls to combat O_3 pollution may have been partly misdirected. Measurements and model calculations now show that O_3 production over most of the United States is primarily NO_x limited, not hydrocarbon limited. The early models were in error in part because they underestimated emissions of hydrocarbons from automobiles, and in part because they did not account for natural emission of biogenic hydrocarbons from trees and crops.

Jacob, Chapter 12, Introduction to Atmospheric Chemistry, 1999

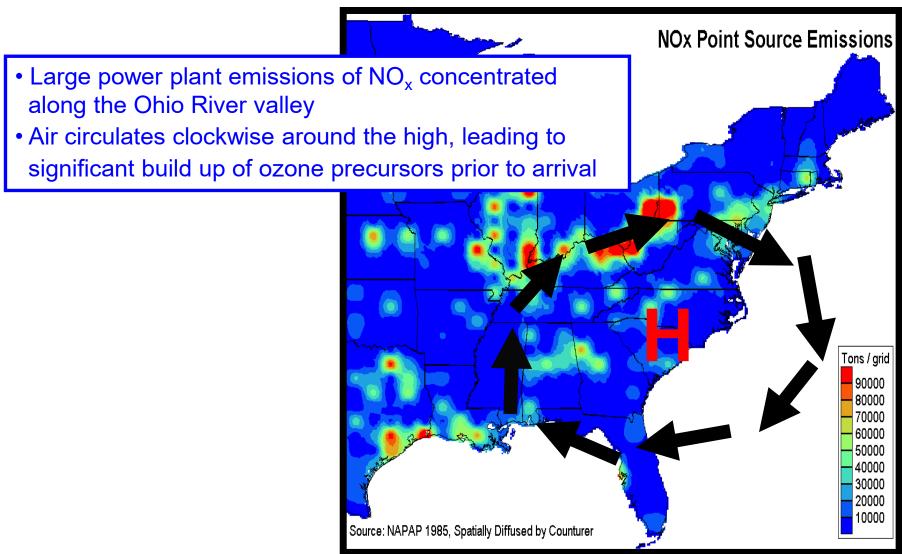
Figure: http://www-personal.umich.edu/~sillman/ozone.htm

Day-to-day meteorology (weather!) affects severity and duration of pollution episodes



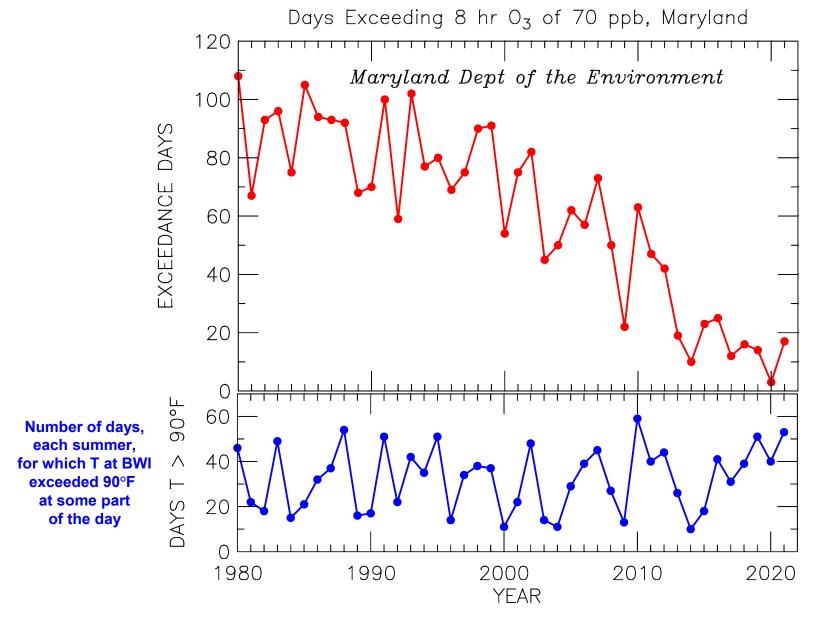
http://www.mde.state.md.us/assets/document/BJH%20-%20Basics%20on%20Ozone%20Transport.ppt

Day-to-day meteorology (weather!) affects severity and duration of pollution episodes



http://www.mde.state.md.us/assets/document/BJH%20-%20Basics%20on%20Ozone%20Transport.ppt

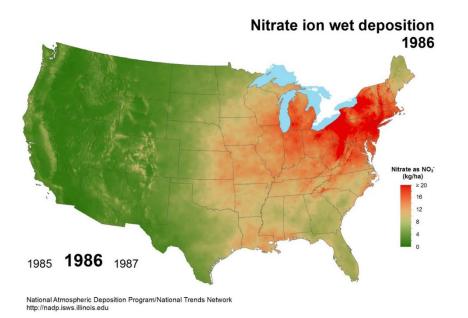
Significant Improvements in *Local* Air Quality since early 1980s



http://www.mde.state.md.us/programs/Air/AirQualityMonitoring/Pages/SeasonalReports.aspx

Nitrate Deposition (see Fig 6.12)

1986 2020



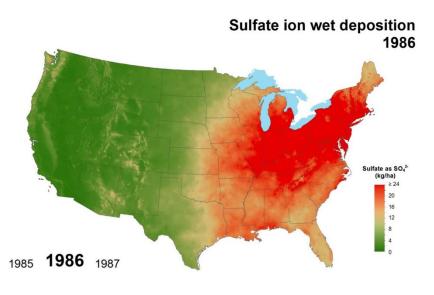


National Atmospheric Deposition Program/National Trends Network http://nadp.slh.wisc.edu

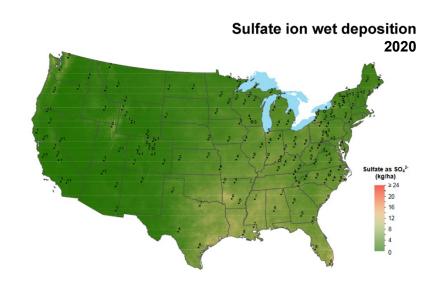
http://nadp.slh.wisc.edu

Sulfate Deposition (see Fig 6.12)

1986 2020



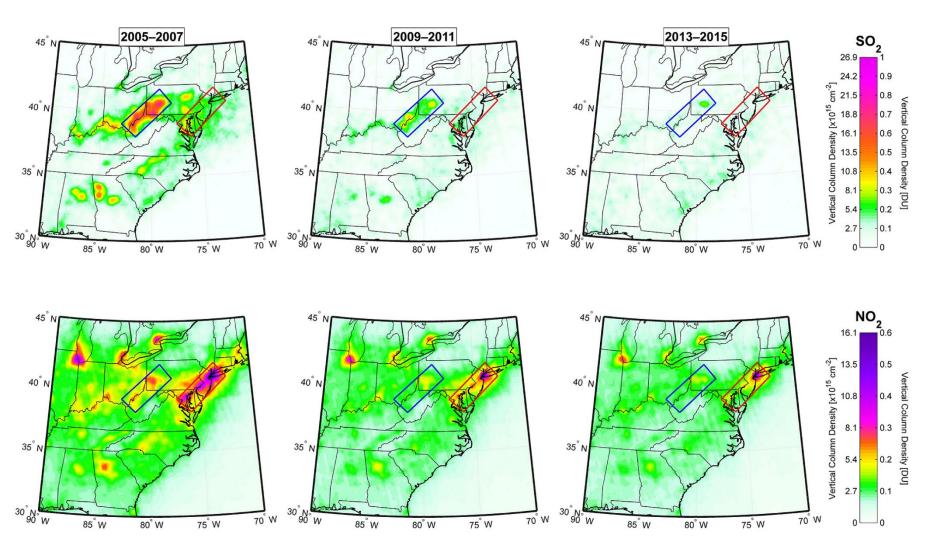
National Atmospheric Deposition Program/National Trends Network http://nadp.isws.illinois.edu



National Atmospheric Deposition Program/National Trends Network http://nadp.slh.wisc.edu

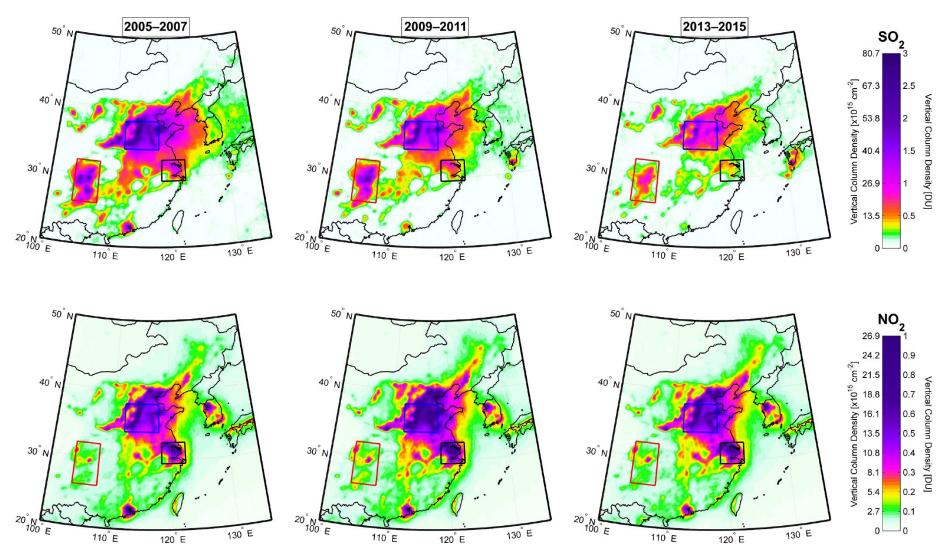
http://nadp.slh.wisc.edu

US Trends: NO₂ and SO₂



Krotkov et al., ACP, 2016

China Trends: NO₂ and SO₂



Krotkov et al., ACP, 2016

Stratospheric Ozone: Chapman Chemistry

- Production of O₃ initiated when O₂ is photodissociated by UV sunlight
- O_3 formed when resulting O atom reacts with O_2 :

$$hv + O_2 \rightarrow O + O \tag{1}$$

$$O + O_2 + M \rightarrow O_3 + M \tag{2}$$

• O₃ removed by photodissociation (UV sunlight) or by reaction with O:

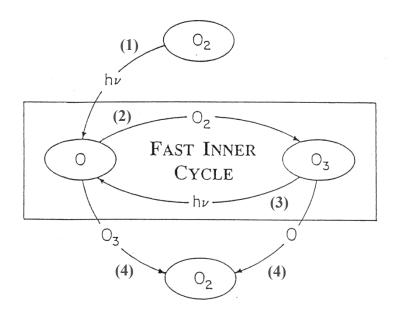
$$hv + O_3 \rightarrow O + O_2$$
 (3)

$$O + O_3 \rightarrow O_2 + O_2 \tag{4}$$

This reaction sequence was first worked out in the 1930s by Sidney Chapman, an English mathematician and geophysicist

Chapman Chemistry

- The cycling between O and O₂ (rxns 2 and 3) occurs *much* more rapidly than leakage into (rxn 1) or out of the system (rxn 4)
- The sum $O + O_3$ is commonly called "odd oxygen"



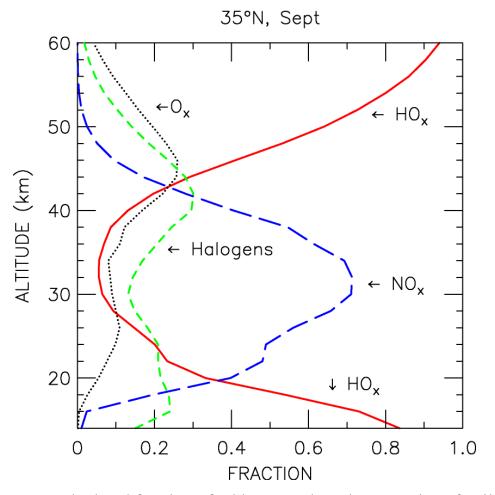
Rxn (1) produces two *odd oxygen* molecules

Rxn (4) consumes two *odd oxygen* molecules

and reactions 2 and 3 recycle *odd oxygen* molecules

Stratospheric Photochemistry: Odd Oxygen Loss By Families

Fraction of O_x Loss Due to Each Catalytic Family JPL 2002 Kinetics



Calculated fraction of odd oxygen loss due to various families of radicals

After Osterman *et al.*, *GRL*, 24, 1107, 1997; Sen *et al.*, *JGR*, 103, 3571. 1998; Sen *et al.*, *JGR*, 104, 26653, 1999.

One Atmosphere – One Photochemistry Troposphere

Stratosphere

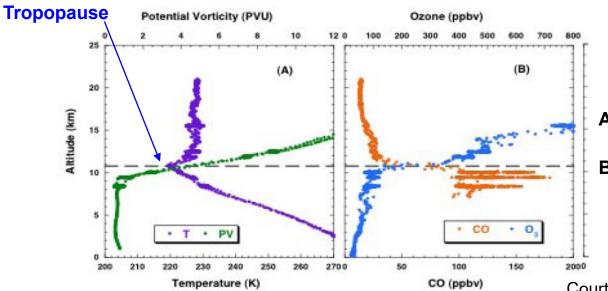
HO₂ formation: $OH + O_3 \rightarrow HO_2 + O_2$ HO_2 loss: $HO_2 + O_3 \rightarrow OH + 2 O_2$

Net: $O_3 + O_3 \rightarrow 3 O_2$

Rate HO_2 Formation = $k_{OH+O3} \times [OH][O_3] + k_{OH+CO} \times [OH][CO]$ Rate HO_2 Loss = $k_{HO2+O3} \times [HO_2][O_3] + k_{HO2+NO} \times [HO_2][NO]$ HO₂ formation: $OH + CO \xrightarrow{O_2} HO_2 + CO_2$ HO₂ loss: $HO_2 + NO \rightarrow OH + NO_2$ Followed by: $NO_2 + hv \rightarrow NO + O$ $O+O_2 + M \rightarrow O_3 + M$

Net: $CO + 2 O_2 \rightarrow CO_2 + O_3$

Rate HO_2 Formation = $k_{OH+O3} \times [OH][O_3] + k_{OH+CO} \times [OH][CO]$ Rate HO_2 Loss = $k_{HO2+O3} \times [HO_2][O_3] + k_{HO2+NO} \times [HO_2][NO]$



Above Tropopause:

Lots of O₃, little CO

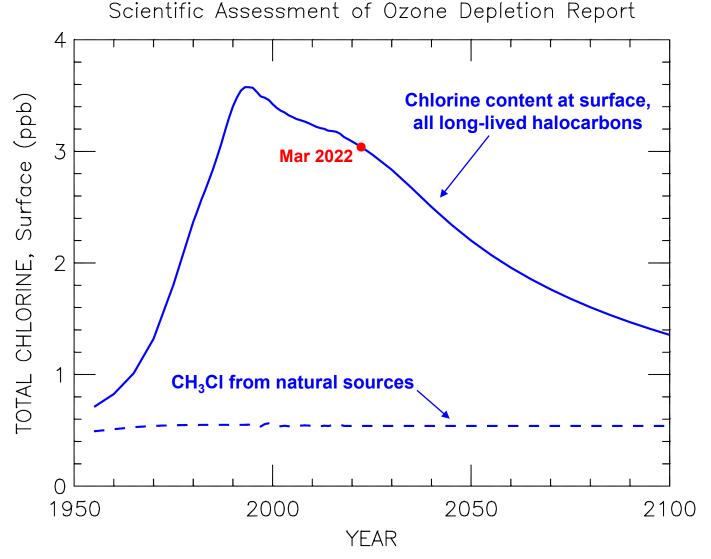
Below Tropopause:

Lots of CO, little O₃

Courtesy of Laura Pan, NCAR

Montreal Protocol Has Banned Industrial Production of CFCs & Other ODS

Projections Based on 2018 World Meteorological Organization



2018 WMO Scientific Assessment of Ozone Depletion Report:

https://www.esrl.noaa.gov/csd/assessments/ozone/2018

Chlorine Abundance, Mid-Latitude Stratosphere

Measurements of Chlorine Gases from Space

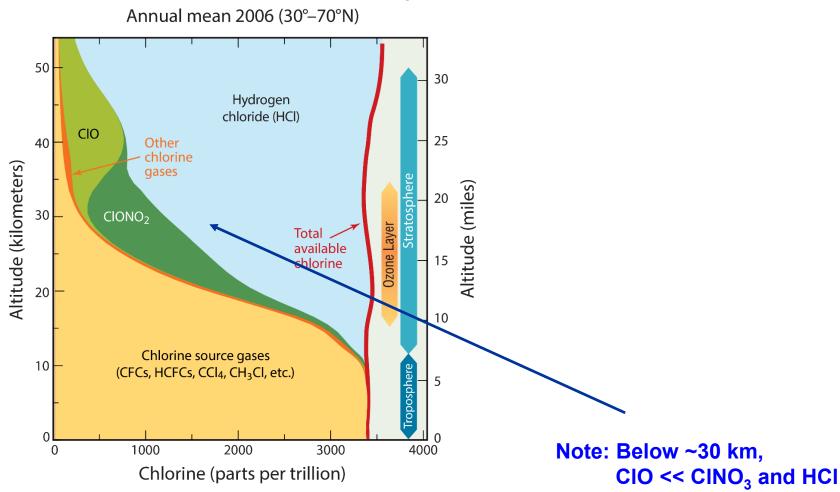


Fig Q7-2, WMO/UNEP Twenty QAs Ozone

Chlorine Source Gases

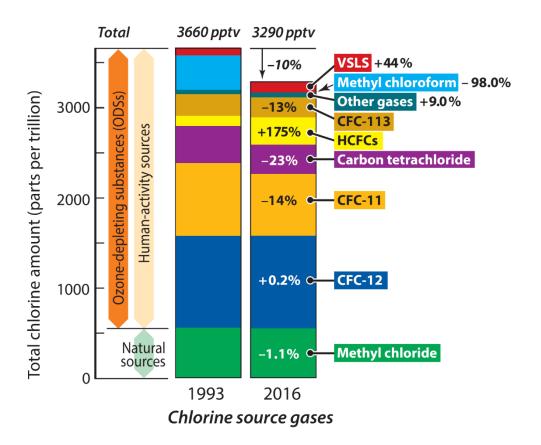


Fig Q6-1, WMO/UNEP Twenty QAs Ozone

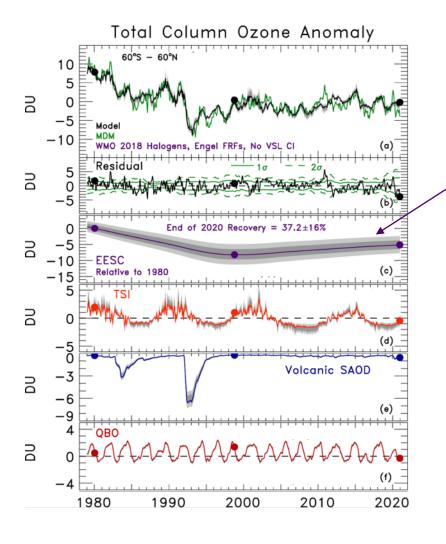
Time series of chlorine content of organic halocarbons that reach the stratosphere. Past values based on direct atmospheric observation. Future values based on projections that include the lifetime for removal of each halocarbon.

Table 6-4, WMO/UNEP 2018

Gas	Atmospheric Lifetime (years)	Ozone Depletion Potential (ODP) ^b	
Halogen Source Gases			
Chlorine Gases		_	
CFC-11 (CCI ₃ F)	52	1	
Carbon tetrachloride (CCI ₄)	32	0.87	
CFC-113 (CCI ₂ FCCIF ₂)	93	0.81	
CFC-12 (CCl ₂ F ₂)	102	0.73	
Methyl chloroform (CH ₃ CCl ₃)	5.0	0.14	
HCFC-141b (CH ₃ CCl ₂ F)	9.4	0.102	
HCFC-142b (CH ₃ CCIF ₂)	18	0.057	
HCFC-22 (CHF ₂ CI)	12	0.034	
Methyl chloride (CH ₃ Cl)	0.9	0.015	
Bromine Gases			
Halon-1301 (CBrF ₃)	65	15.2	
Halon-1211 (CBrClF ₂)	16	6.9	
Methyl bromide (CH ₃ Br)	0.8	0.57	
Hydrofluorocarbons (HFCs	;)		
HFC-23 (CHF ₃)	228	0	
HFC-143a (CH ₃ CF ₃)	51	0	
HFC-125 (CHF ₂ CF ₃)	30	0	
HFC-134a (CH ₂ FCF ₃)	14	0	
HFC-32 (CH ₂ F ₂)	5.4	0	
HFC-152a (CH ₃ CHF ₂)	1.6	0	
HFO-1234yf (CF ₃ CF=CH ₂)	0.03	0	

Mid-Latitude Ozone Depletion

Total column ozone anomaly is deseasonalized, cosine latitude weighted average of total column ozone collected between 60°S and 60°N, relative to the mean total column abundance over the entire time period.



"Expected" recovery of near global ozone layer for end of 2019 relative to maximum depletion since 1980, driven by atmospheric halogens

Circles (•) placed at 1980, column minimum due to EESC, & end of 2020

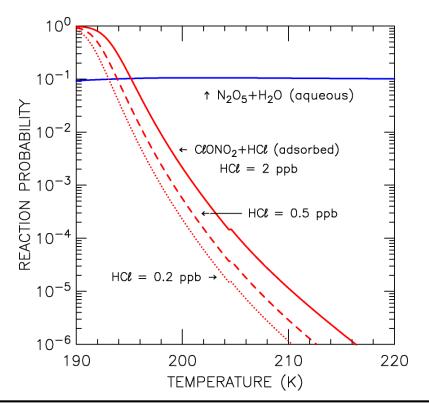
Multiple linear regression of total ozone column anomaly as a function of equivalent effective stratospheric chlorine (EESC), total solar irradiance (TSI), stratospheric aerosol optical depth (SAOD), and the quasi-biennial oscillation (QBO) has long been used to quantitatively assess factors that drive variations in the thickness of the ozone layer.

Note: EESC = Inorganic Stratospheric Chlorine + 60× Inorganic Stratospheric Bromine

McBride et al., In Prep, 2022

Heterogeneous Chemistry, Mid-Latitude vs Polar Regions

In all cases, γ must be measured in the laboratory



Reaction probabilities given for various surface types, with formulations of various degrees of complexity, in **Section 5** of the JPL Data Evaluation.

Atmospheric Chemistry and Physics by Seinfeld and Pandis provides extensive treatment of aqueous phase chemistry, properties of atmospheric aerosol, organic aerosols, etc.

Polar Ozone Loss

- COLD TEMPERATURES → POLAR STRATOSPHERIC CLOUDS (PSCs)
- Reactions on PSC surfaces lead to elevated <u>CIO</u>

```
HCI + CINO<sub>3</sub> \rightarrow Cl<sub>2</sub> (gas) + HNO<sub>3</sub> (solid)

CINO<sub>3</sub> + H<sub>2</sub>O \rightarrowHOCI + HNO<sub>3</sub>

Cl<sub>2</sub> + SUNLIGHT + O<sub>3</sub> \rightarrow CIO

HOCI + SUNLIGHT + O<sub>3</sub> \rightarrow CIO

HNO<sub>3</sub> SEDIMENTS (PSCs fall due to gravity)
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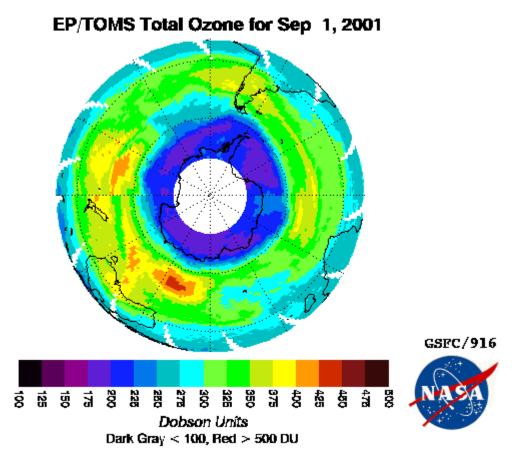
- ELEVATED CIO + SUNLIGHT DESTROYS O₃
- BrO: Reaction Partner For CIO \Rightarrow Additional O₃ Loss



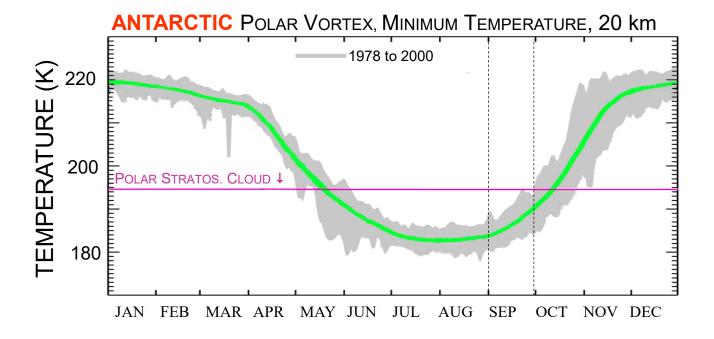
Polar Vortex Circulation

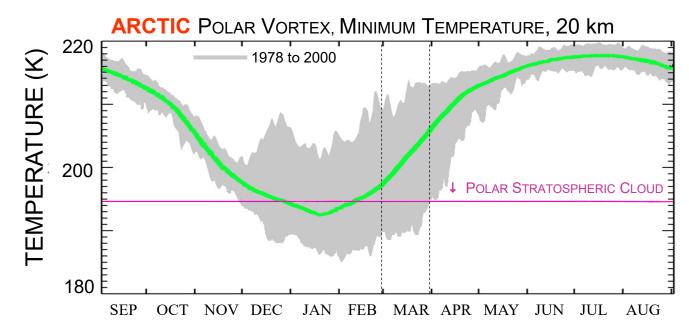
During winter:

- radiative cooling leads to cold air in polar stratosphere
- large scale low pressure region develops over pole
- strong "polar night jet" develops, isolating air at high latitudes from air at low latitudes
- T continues to fall in the "vortex like" circulation near the pole



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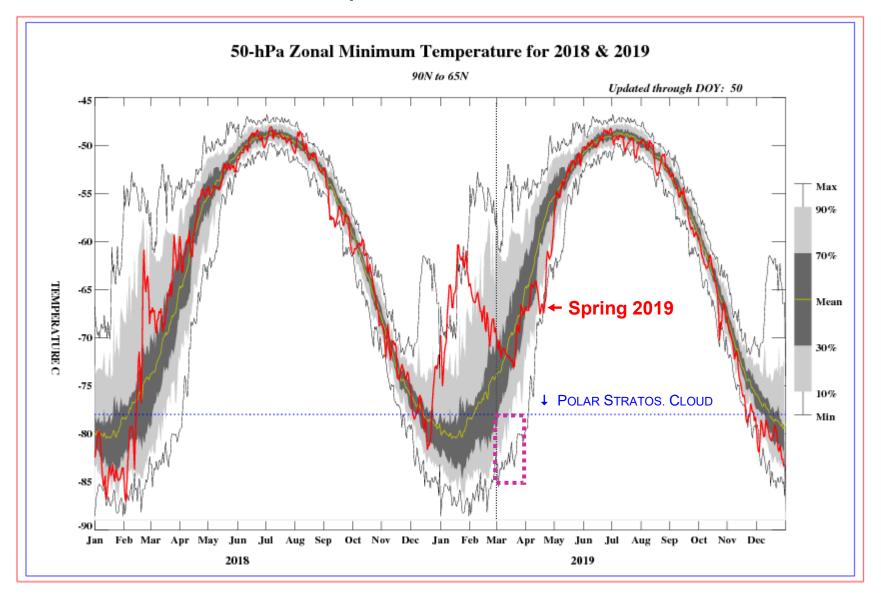




NASA/GSFC

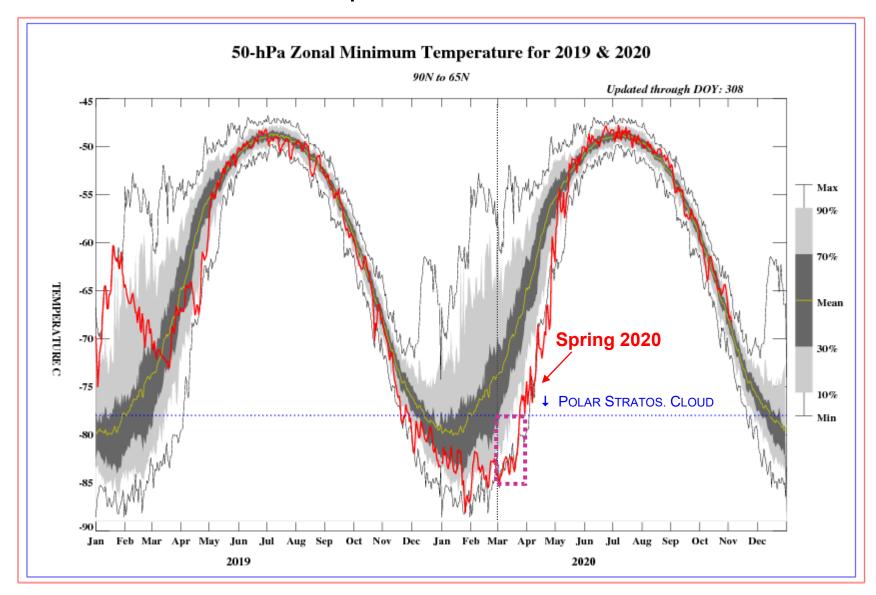
Data Courtesy P. Newman,

Arctic Temperature: Mar 2019



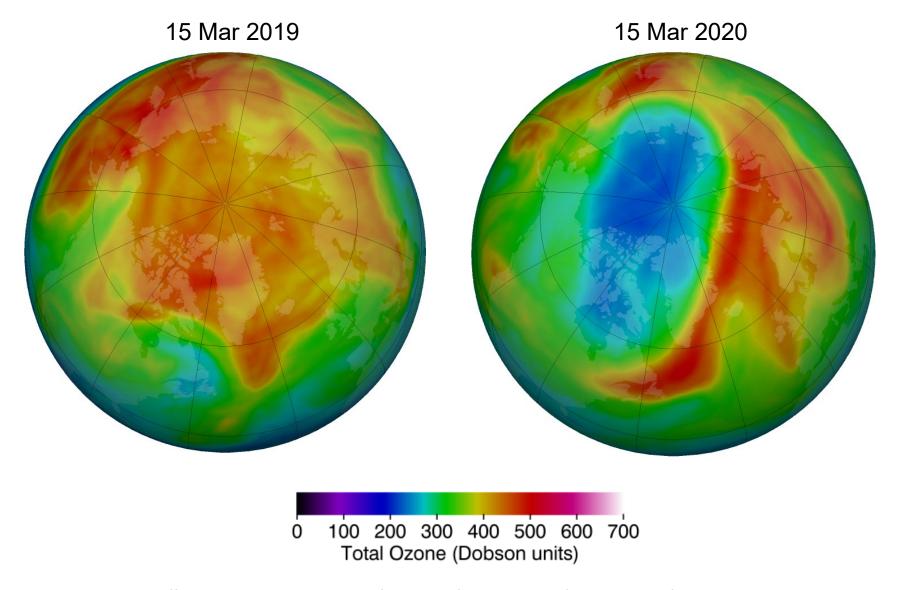
http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/50mbnhlo.png

Arctic Temperature: Mar 2020



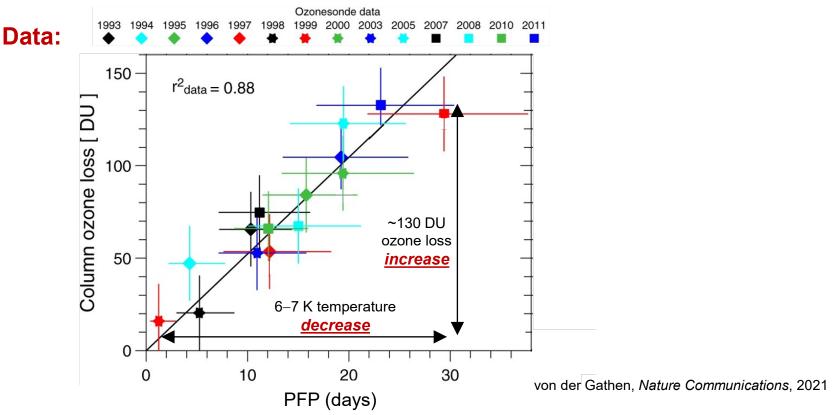
http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/50mbnhlo.png

Arctic Ozone: 2019 and 2020



http://www.cpc.ncep.noaa.gov/products/stratosphere/temperature/50mbnhlo.png

Arctic Ozone Loss Varies as a function of PSC Formation Potential



 Surprisingly simple relationship between chemical loss of column ozone and volume of air exposed to PSC formation potential over winter, where

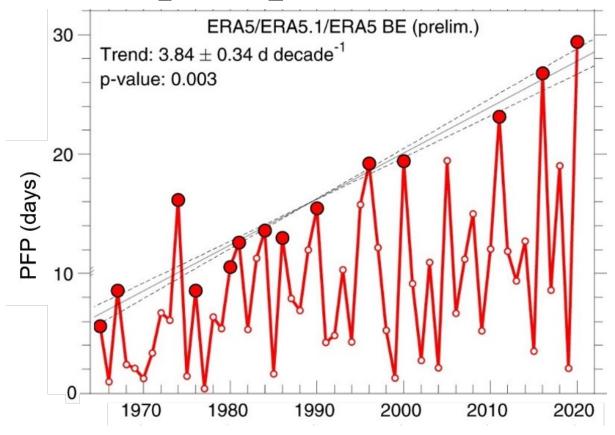
PFP =
$$\int_{1 \text{ Nov}}^{30 \text{Apr}} \frac{V_{\text{PSC}}(t)}{V_{\text{VORTEX}}(t)} dt$$
; PFP stands for PSC Formation Potential

and V_{PSC} is the volume of the vortex where T is cold enough to allow for formation of PSCs, and V_{VORTEX} is the volume of the Arctic vortex

Relation leads to estimate of ~20 DU additional loss of ozone per degree Kelvin cooling of Arctic stratosphere

Cold Arctic Winters Tend to Exhibit Larger PFP as a Function of Time

More Data: PFP is PSC Formation Potential



von der Gathen, Nature Communications, 2021

PSC Formation Potential in Arctic Vortex

based on 55 years of data from the European Centre for Medium-Range Weather Forecasts (ECMWF)

SOLID CIRCLES denote local maxima in PFP relative to a trend line

Future Ozone: ODSs, CO₂, CH₄ and N₂O

Global Total Ozone Changes in Response to Ozone Depleting Substances and Greenhouse Gases

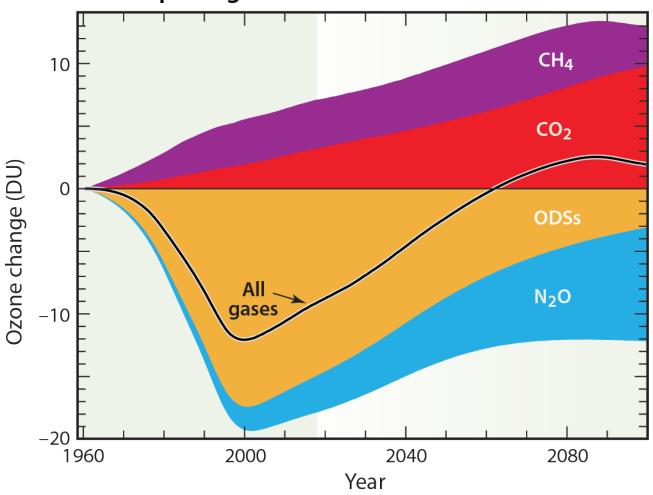
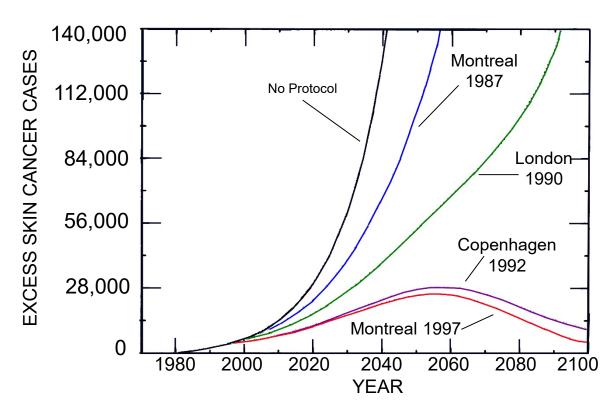


Fig Q20-3, 20 QAs, WMO (2019)

EXCESS SKIN CANCER CASES IN THE UNITED STATES, PER YEAR, DUE TO OZONE DEPLETION FOR VARIOUS CFC EMISSION SCENARIOS



Longstreth et al., J. of Photochemistry and Photobiology B, 46, 20–39, 1998.

See also Slaper *et al.*, Estimates of ozone depletion and skin cancer incidence to examine the Vienna Convention achievements, *Nature*, *384*, 256–258, 1996, who state:

The no-restrictions and Montreal Protocol scenarios produce a runaway increase in skin cancer incidence, up to a quadrupling and doubling, respectively, by year 2100.